

# 2D and 3D Multiscale Geometric Transforms

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# What is a good representation for data?

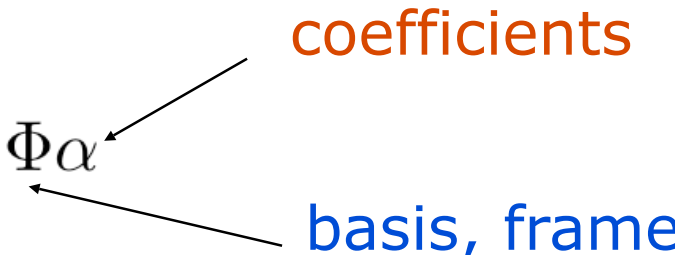
- Computational harmonic analysis seeks representations of a signal as linear combinations of basis, frame, dictionary, element :

$$\Phi = \{\phi_1, \dots, \phi_K\}$$

$$s = \sum_{k=1}^K \alpha_k \phi_k = \Phi \alpha$$

coefficients

basis, frame



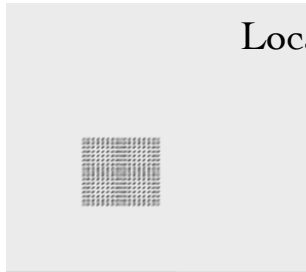
- Fast calculation of the coefficients
- Analyze the signal through the statistical properties of the coefficients
- Approximation theory uses the sparsity of the coefficients.

**Sparsity Model 1:** we consider a dictionary which has a fast transform/reconstruction operator:

$$\Phi = \{\phi_1, \dots, \phi_K\}$$

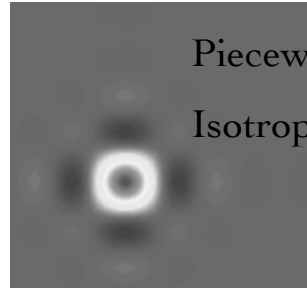
$$s = \sum_{k=1}^K \alpha_k \phi_k = \Phi \alpha$$

Local DCT Stationary textures



Locally oscillatory

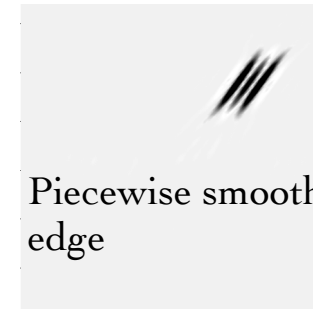
Wavelet transform



Piecewise smooth

Isotropic structures

Curvelet transform



Piecewise smooth, edge

**Sparsity Model 2:** Morphological Diversity:

$$\phi = [\phi_1, \dots, \phi_L], \quad \alpha = \{\alpha_1, \dots, \alpha_L\}, \quad s = \phi \alpha = \sum_{k=1}^L \phi_k \alpha_k$$

**Sparsity Model 3:** we adapt/learn the dictionary directly from the data

Model 3 can be also combined with model 2:



# Sparsity Model 1: Multiscale Transforms

## Critical Sampling

(bi-) Orthogonal WT  
Lifting scheme construction  
Wavelet Packets  
Mirror Basis

## Redundant Transforms

Pyramidal decomposition (Burt and Adelson)  
**Undecimated Wavelet Transform**  
**Isotropic Undecimated Wavelet Transform**  
Complex Wavelet Transform  
Steerable Wavelet Transform  
Dyadic Wavelet Transform  
Nonlinear Pyramidal decomposition (Median)

## **New Multiscale Construction**

Contourlet  
Bandelet  
Finite Ridgelet Transform  
Platelet  
(W-)Edgelet  
Adaptive Wavelet  
Groupelet

**Ridgelet**  
**Curvelet** (Several implementations)  
Wave Atom

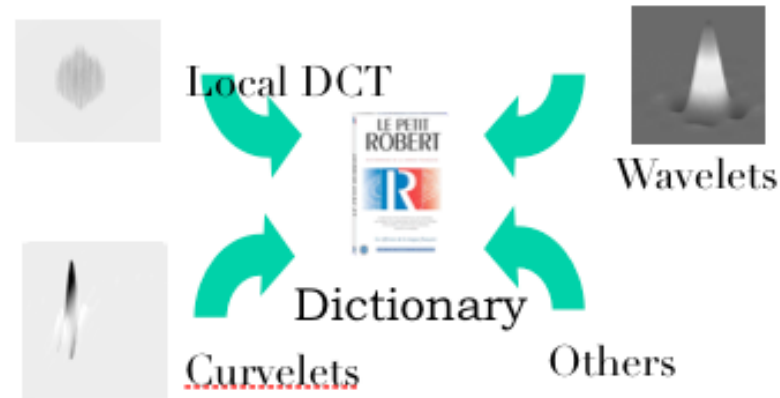
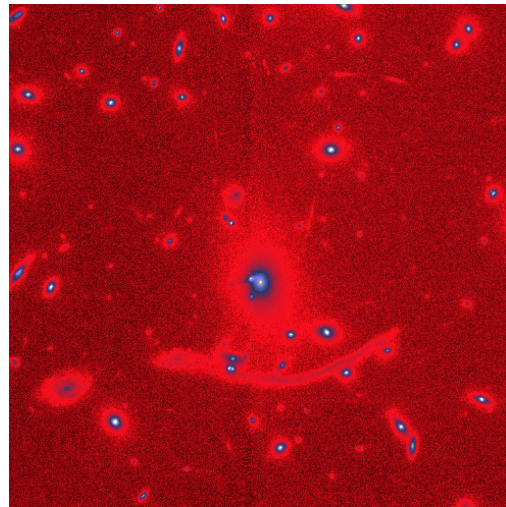
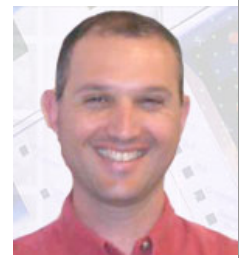


# Morphological Diversity



•J.-L. Starck, M. Elad, and D.L. Donoho, *Redundant Multiscale Transforms and their Application for Morphological Component Analysis*, *Advances in Imaging and Electron Physics*, 132, 2004.

•J.-L. Starck, M. Elad, and D.L. Donoho, *Image Decomposition Via the Combination of Sparse Representation and a Variational Approach*, *IEEE Trans. on Image Proces.*, 14, 10, pp 1570--1582, 2005.



$$\phi = [\phi_1, \dots, \phi_L], \quad \alpha = \{\alpha_1, \dots, \alpha_L\}, \quad s = \phi\alpha = \sum_{k=1}^L \phi_k \alpha_k$$

**Sparsity Model 2:** we consider a signal as a sum of  $K$  components  $s_k$ ,  $s = \sum_{k=1}^K s_k$  each of them being sparse in a given dictionary :

$$s_k = \Phi_k \alpha_k$$
$$s = \sum_{k=1}^K s_k = \sum_{k=1}^K \Phi_k \alpha_k = \Phi \alpha$$

**Advantages of model 1:** extremely fast.

**Advantages of model 2:**

- more flexible to model 1.
- The coupling of local DCT+curvelet is well adapted to a relatively large class of images.

**Advantages of model 3:**

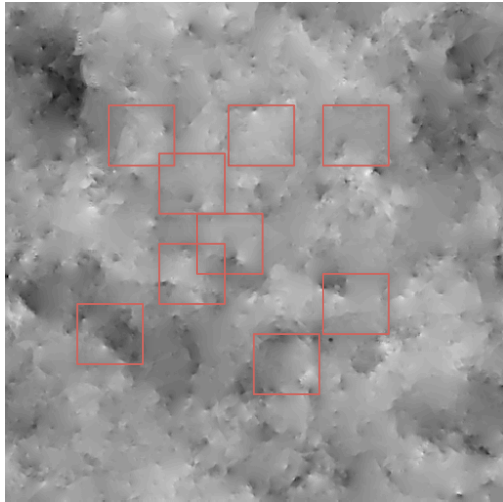
atoms can be obtained which are well adapted to the data, and which could never be obtained with a fixed dictionary.

**Drawback of model 3 versus model 1,2:**

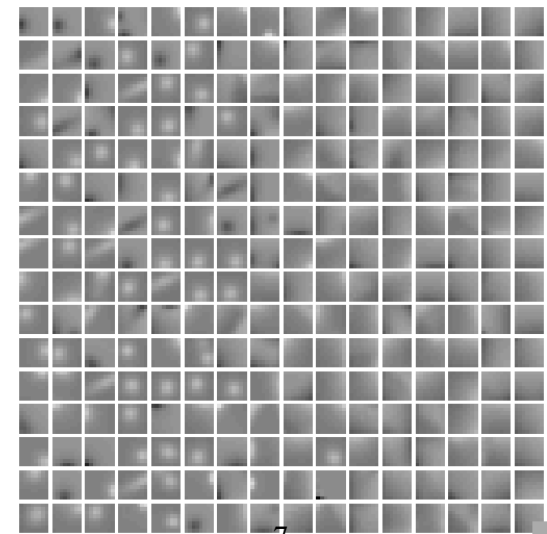
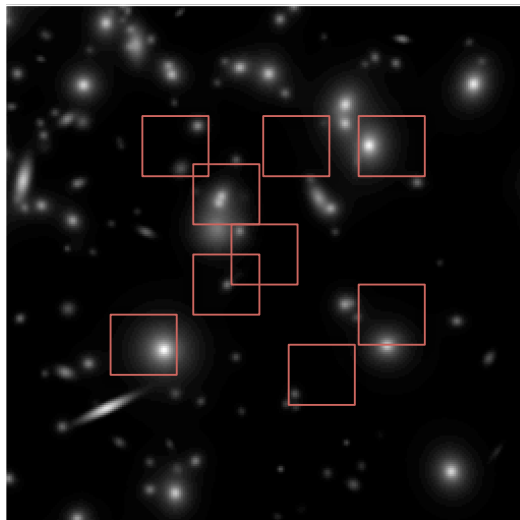
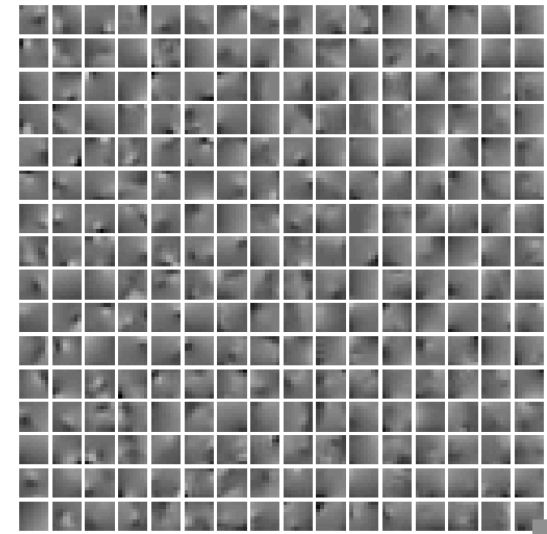
We pay the price of dictionary learning by being less sensitive to detect very faint features.

Complexity: Computation time, parameters, etc

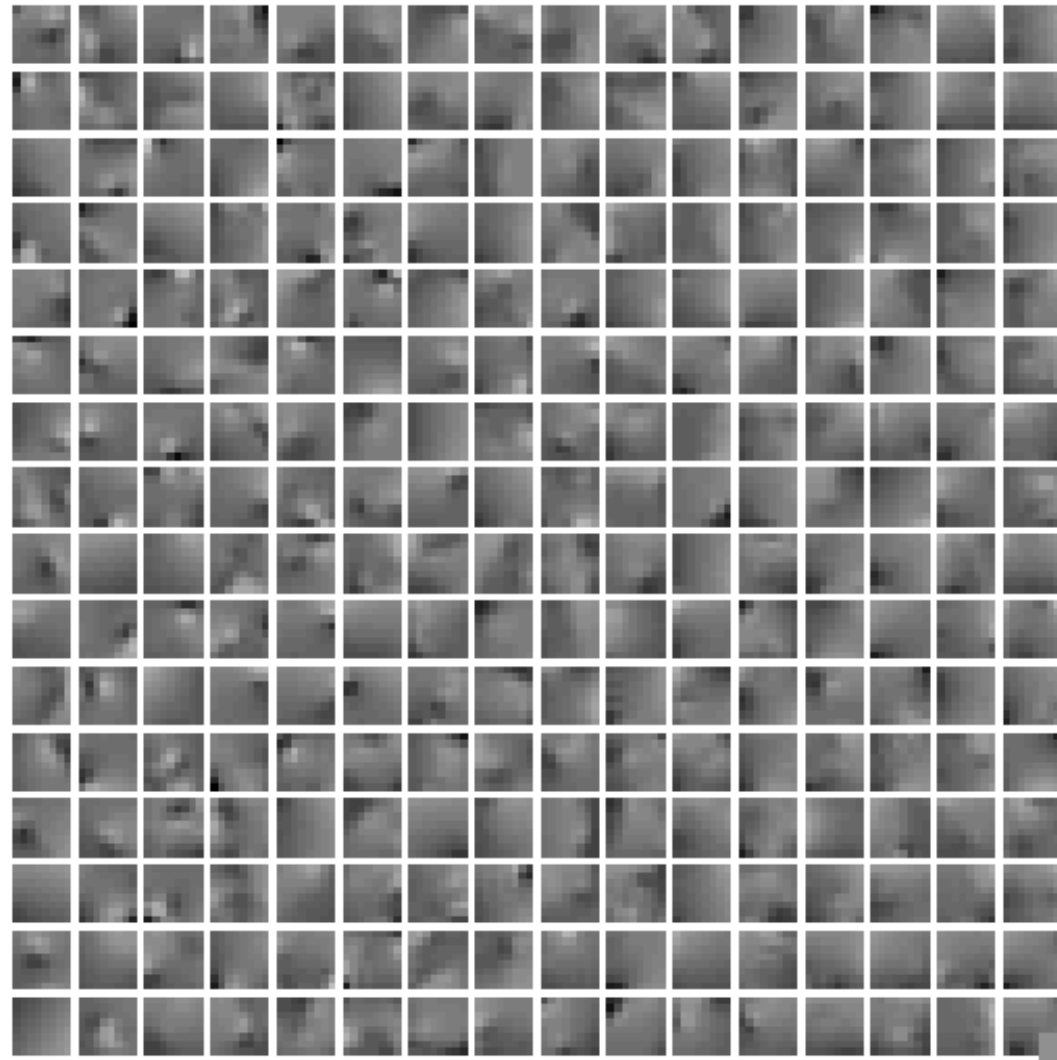
## Sparsity Model 3: Dictionary Learning



Training basis.



## Sparsity Model 3: Dictionary Learning



## **2D and 3D Multiscale Geometric Transforms**

- Ridgelet 2D
- Curvelet 2D
- BeamCurvelet 3D
- RidCurvelet 3D
- FastCurvelet 3D
- 3D Morphological Diversity

# Problems related to the WT

1) Edges representation:

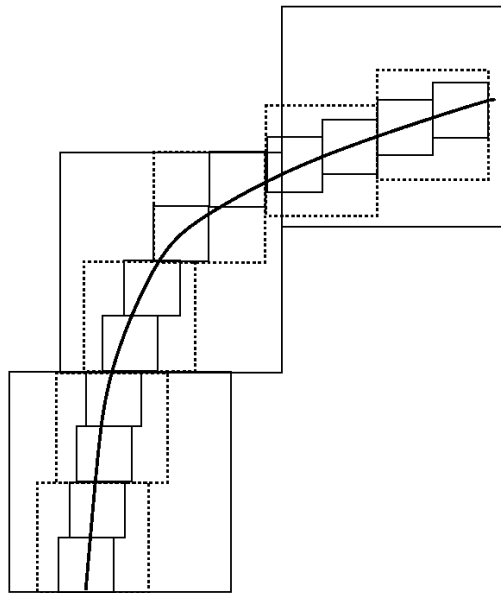
if the WT performs better than the FFT to represent edges in an image, it is still not optimal.

2) There is only a fixed number of directional elements independent of scales.

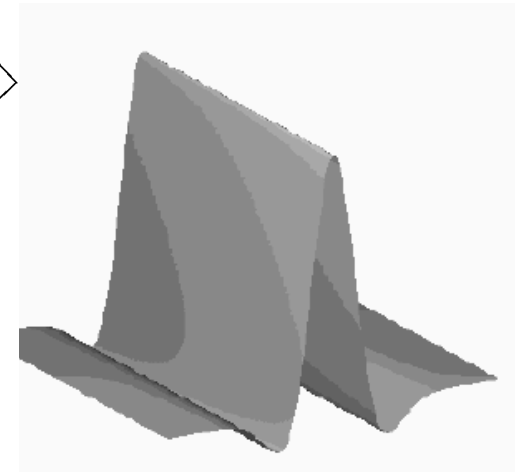
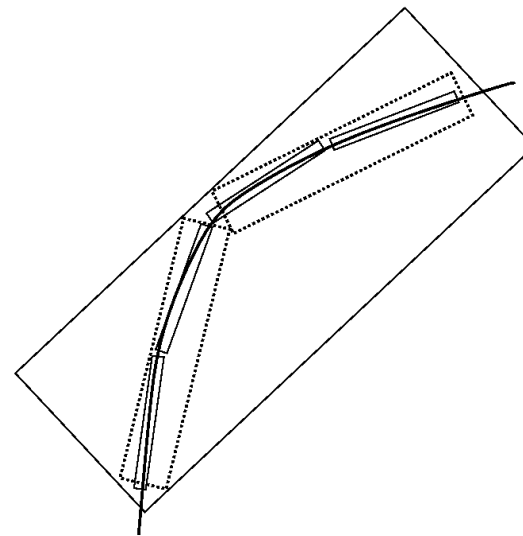
3) Limitation of existing scale concepts: there is no highly anisotropic elements.

# Wavelets and edges

- many wavelet coefficients are needed to account for edges i.e. singularities along lines or curves :

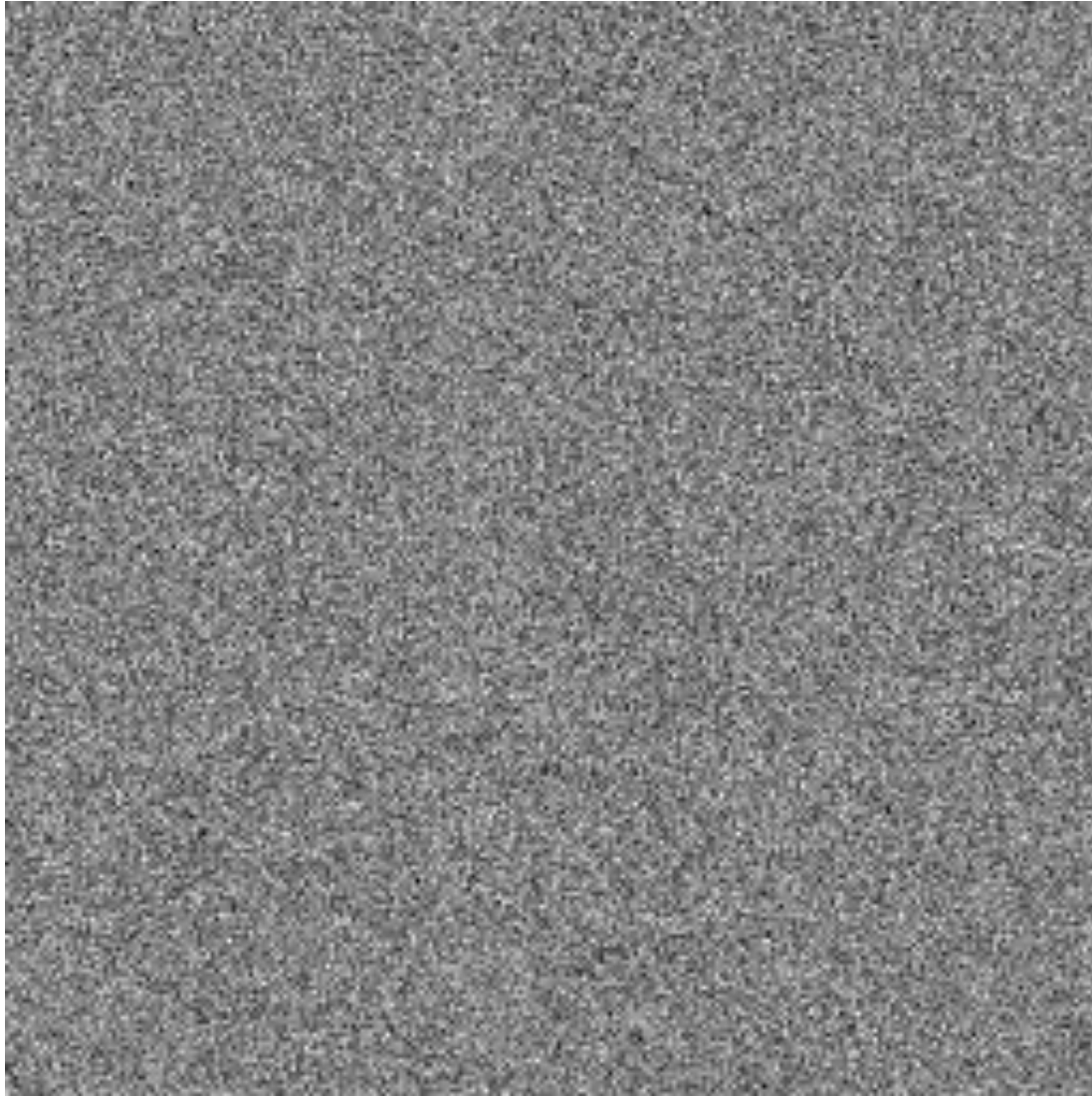


- need dictionaries of strongly anisotropic atoms :

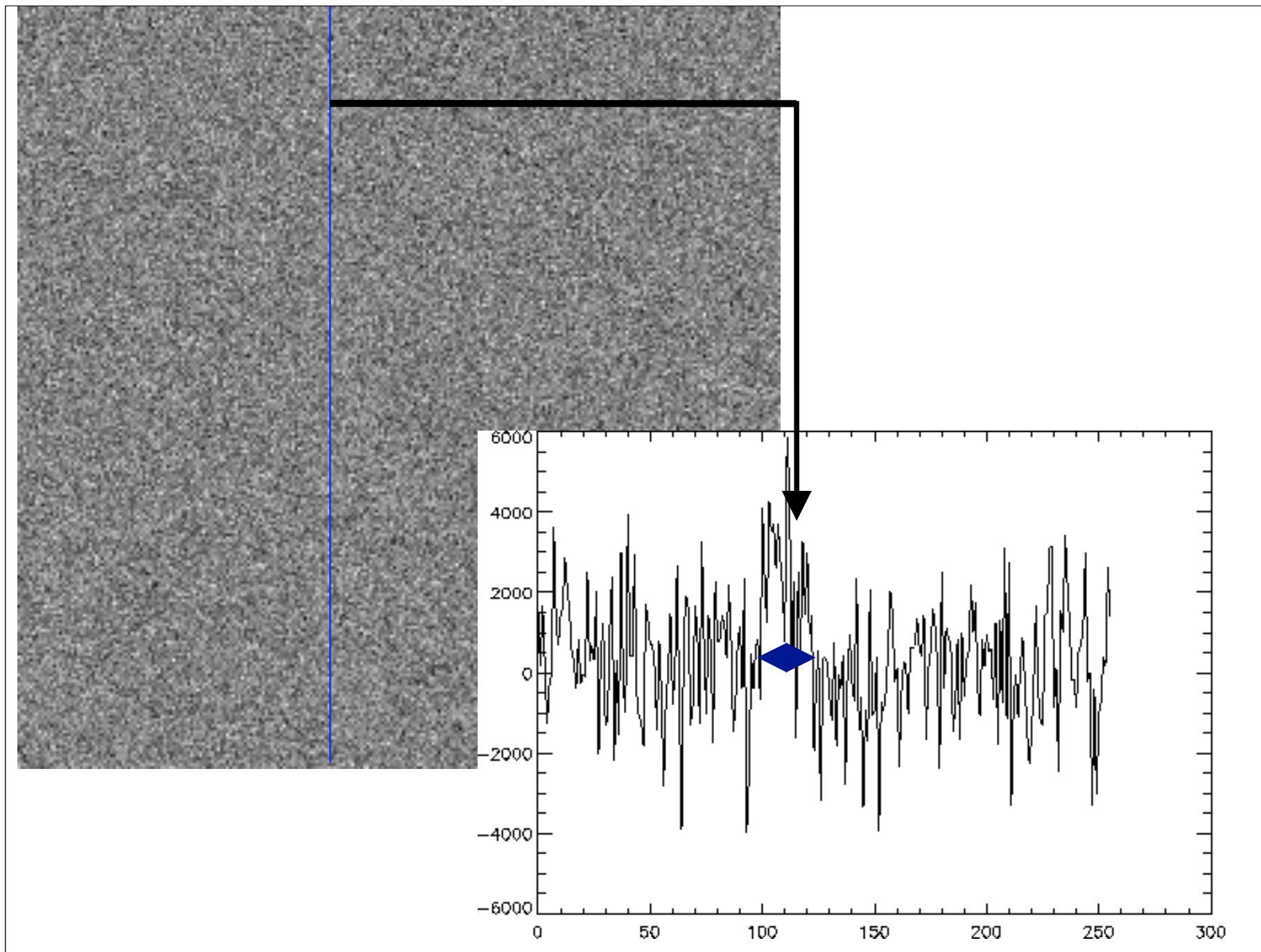


ridgelets, curvelets, contourlets, bandelettes, etc.

SNR = 0.1



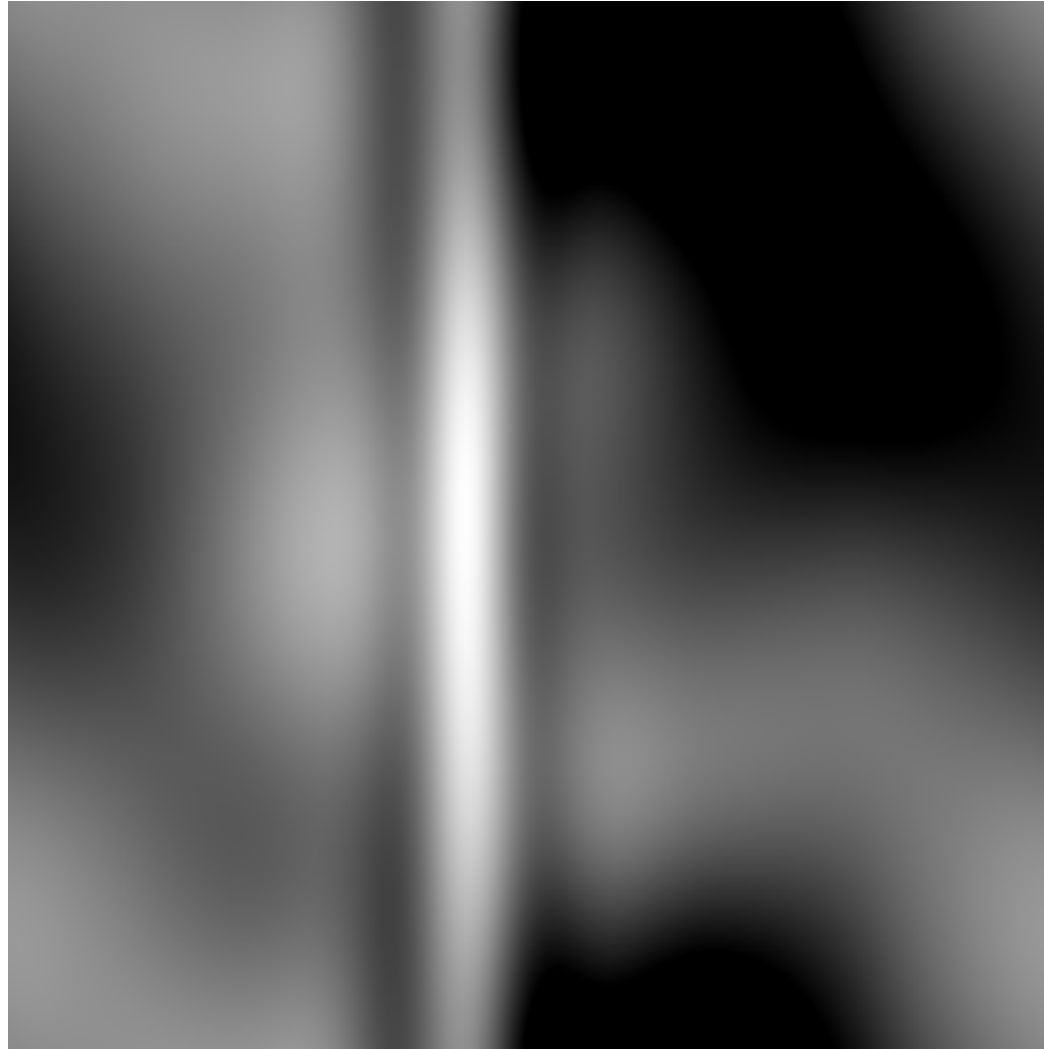




## Undecimated Wavelet Filtering (3 sigma)



## Ridgelet Filtering (5sigma)



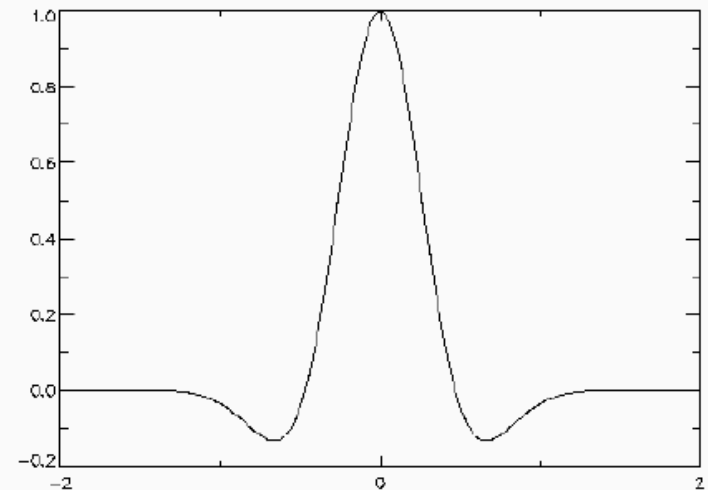
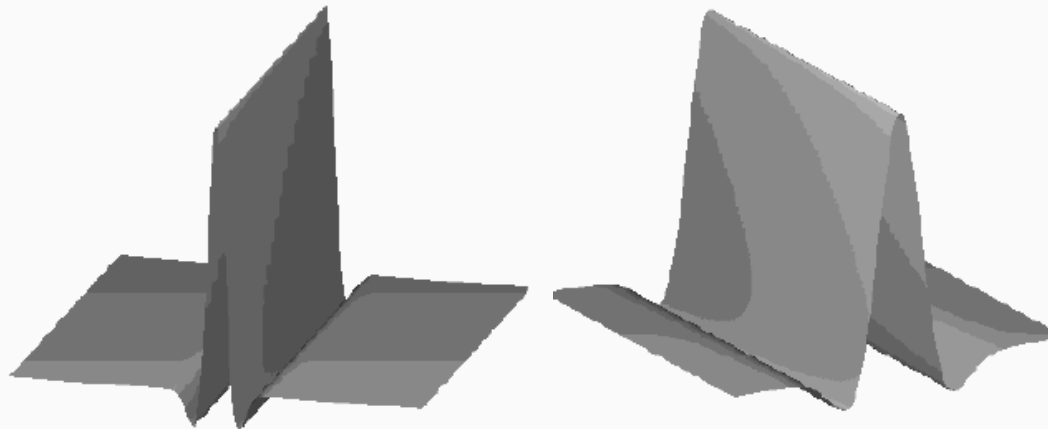


# Continuous Ridgelet Transform

Ridgelet Transform (Candes, 1998):  $R_f(a,b,\theta) = \int \psi_{a,b,\theta}(x) f(x) dx$

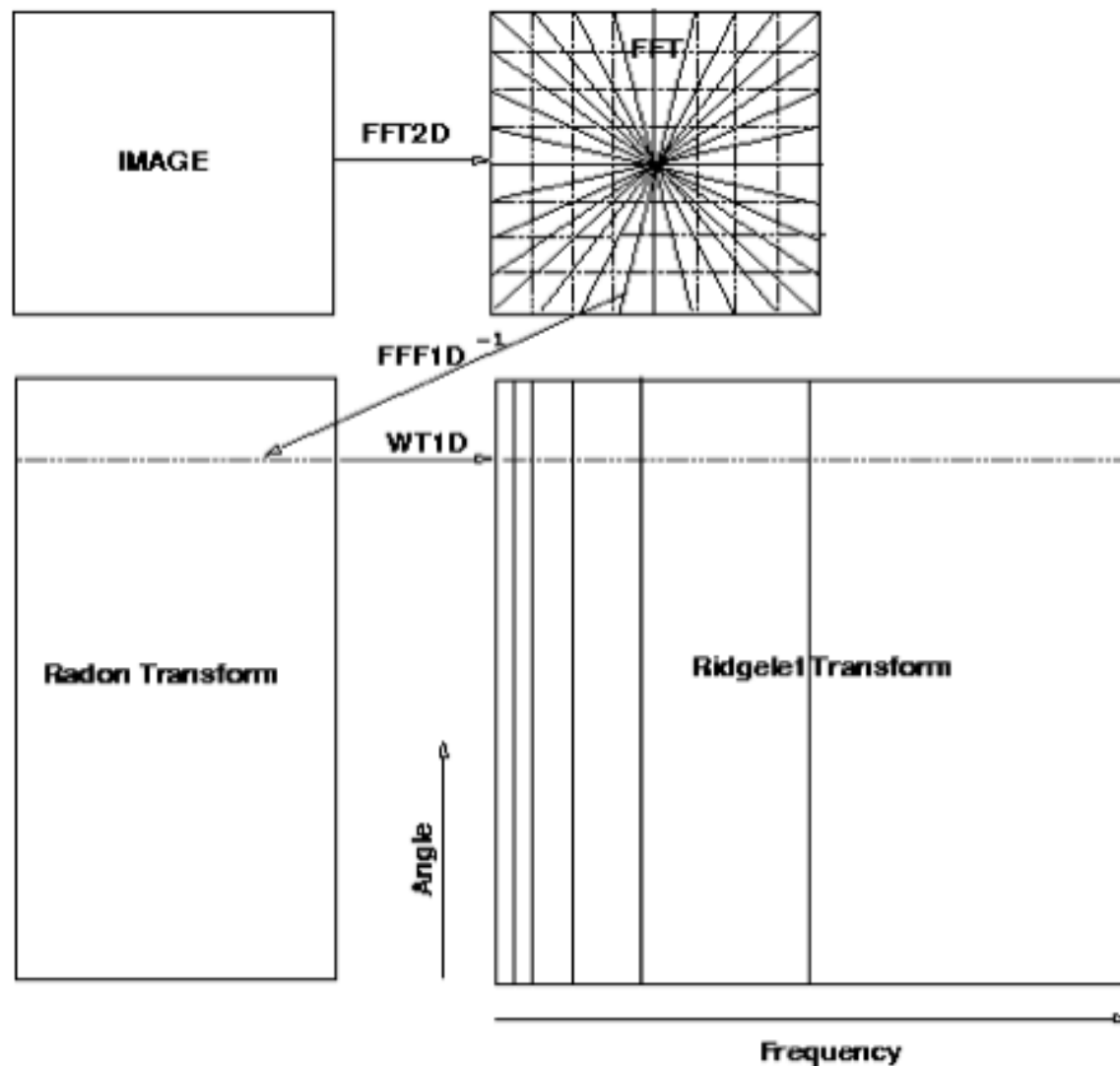
Ridgelet function:  $\psi_{a,b,\theta}(x) = a^{-\frac{1}{2}} \psi\left(\frac{x_1 \cos(\theta) + x_2 \sin(\theta) - b}{a}\right)$

The function is constant along lines. Transverse to these ridges, it is a wavelet.

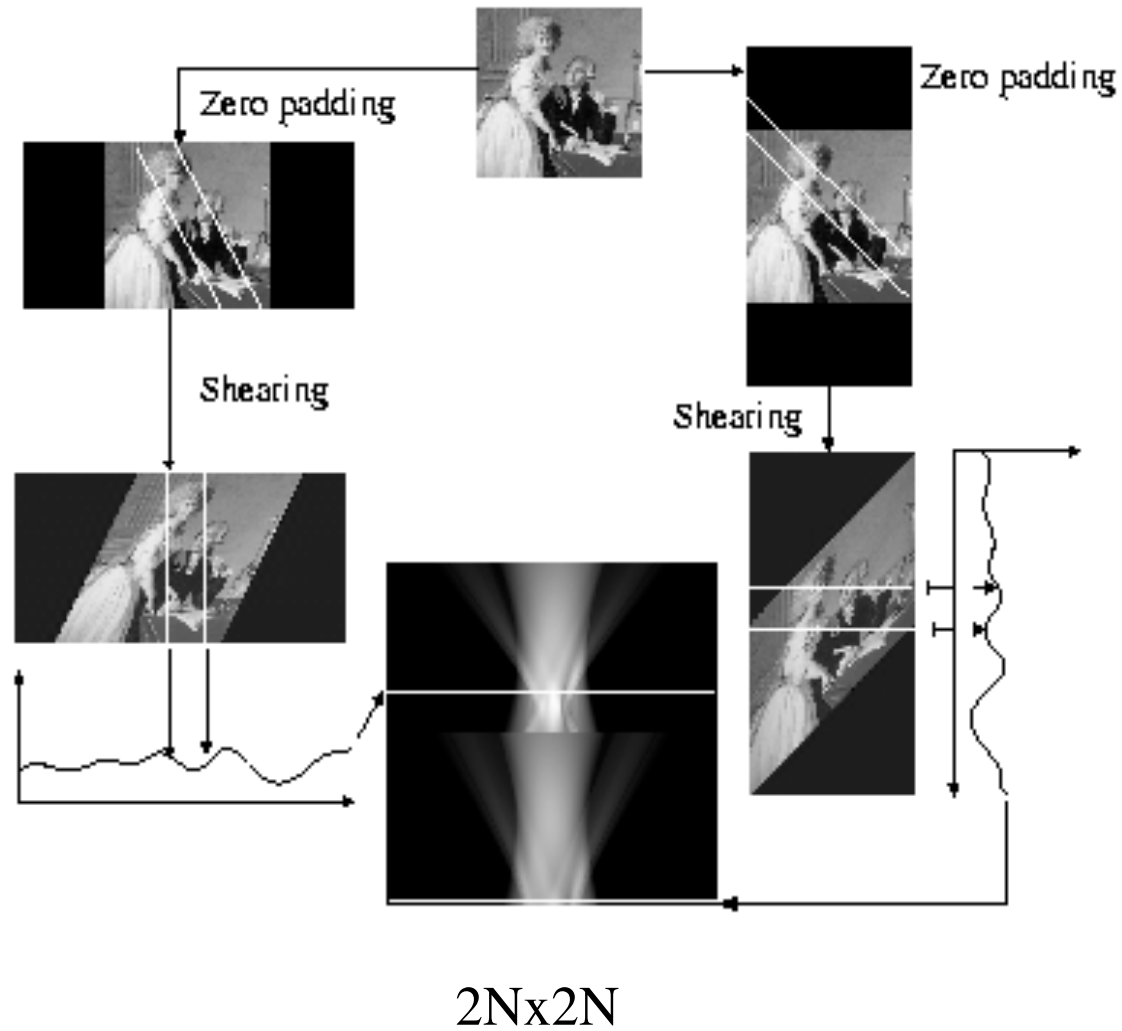


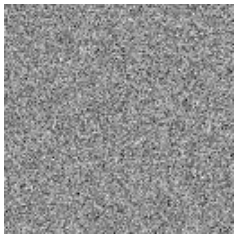
The ridgelet coefficients of an object  $f$  are given by analysis

of the Radon transform via: 
$$R_f(a,b,\theta) = \int Rf(\theta,t)\psi\left(\frac{t-b}{a}\right)dt$$



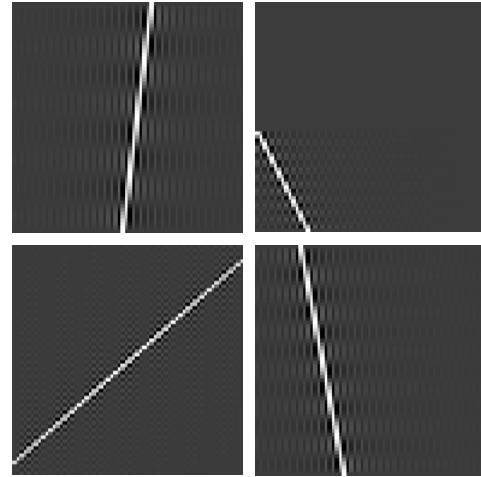
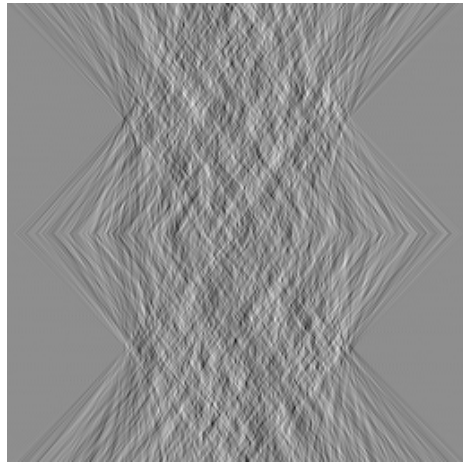
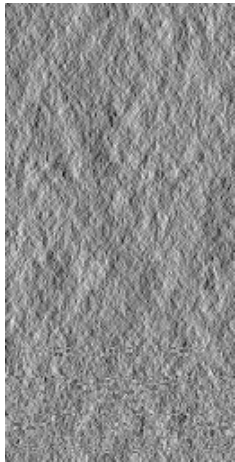
# Slant Stack Radon Transform (Averbuch et al, 2001) CUR01-SSR



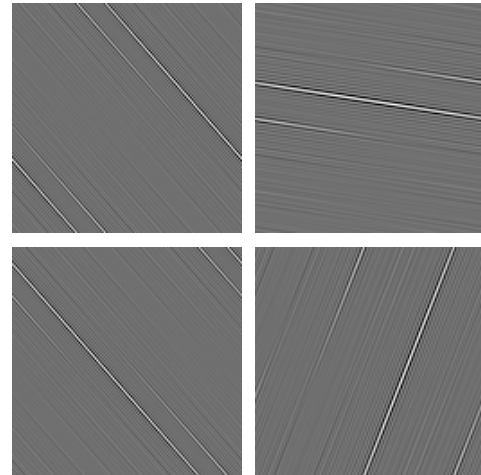


Linogram

Slant Stack Radon



Backprojection

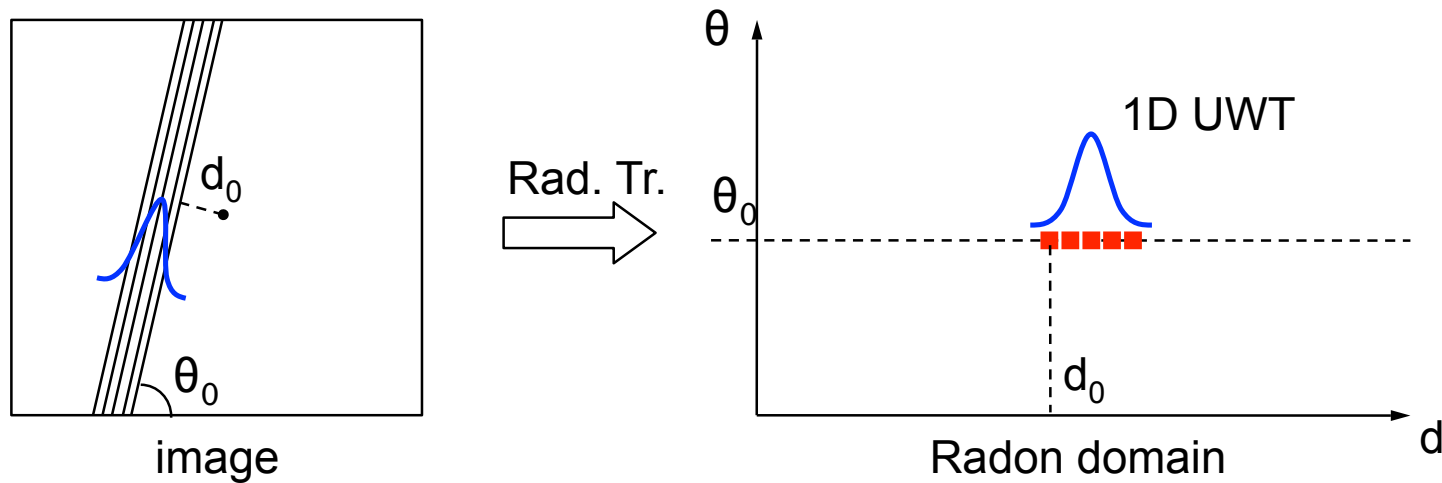


Backprojection



## Ridgelet Transform

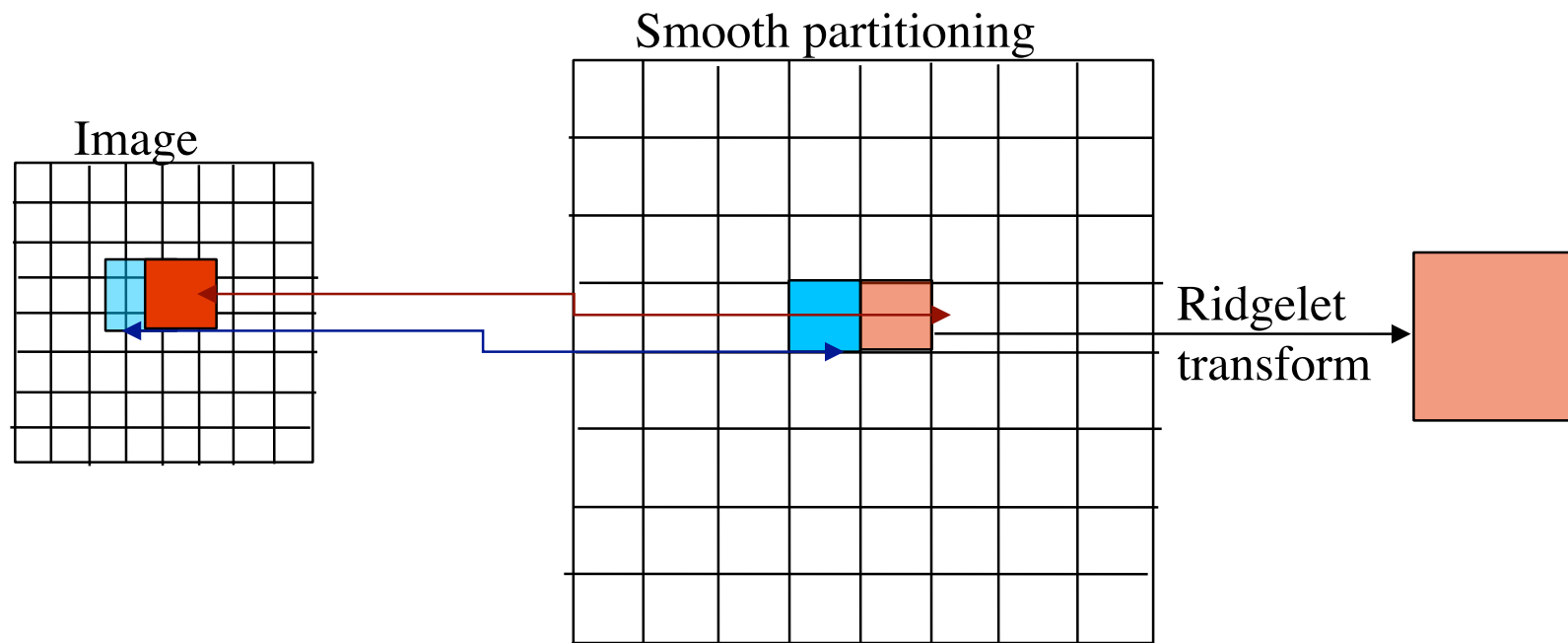
- Ridgelet transform: Radon + 1D Wavelet



1. Rad. Tr.
2. For each line, apply the same denoising scheme as before



# LOCAL RIDGELET TRANSFORM



The partitioning introduces a redundancy, as a pixel belongs to 4 neighboring blocks.



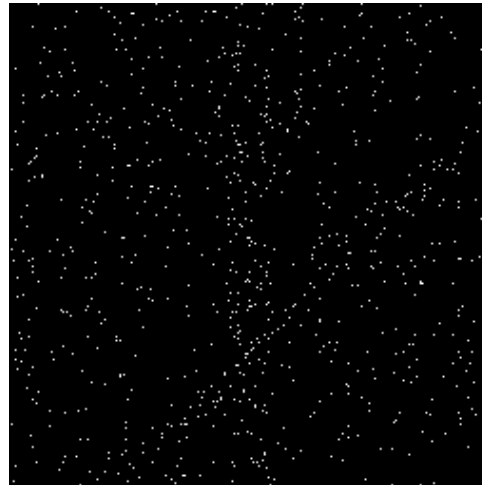
## Poisson Noise and Line-Like Sources Restoration (MS-VST + Ridgelet)



*B. Zhang, M.J. Fadili and J.-L. Starck, "Wavelets, Ridgelets and Curvelets for Poisson Noise Removal" ,ITIP, Vol 17, No 7, pp 1093--1108, 2008.*



underlying intensity image



simulated image of counts



restored image  
from the left image of counts

### Max Intensity

background = 0.01

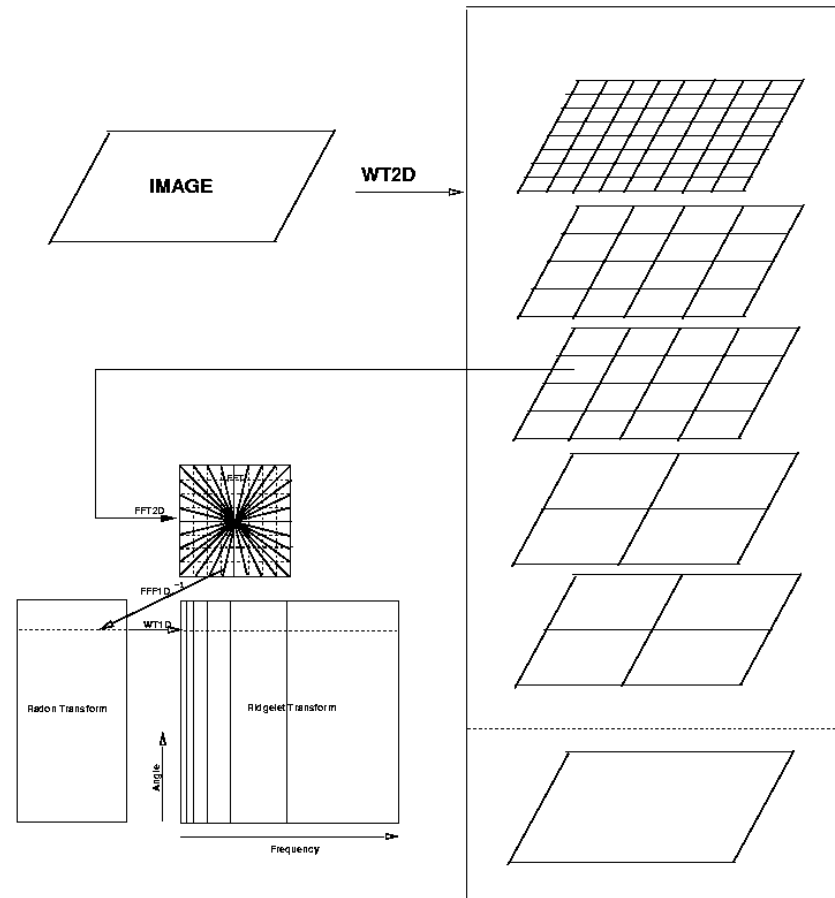
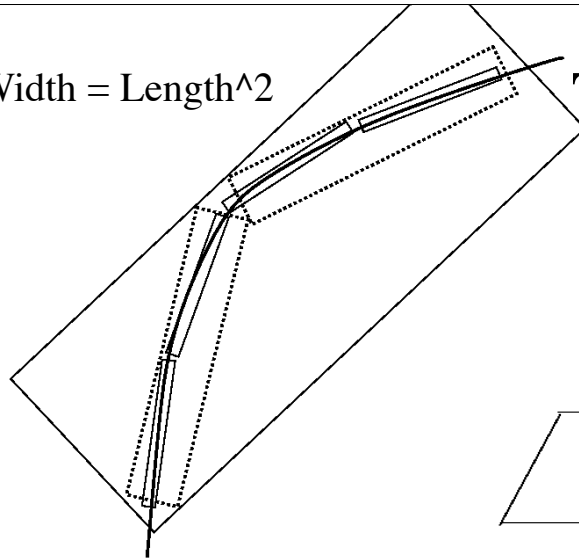
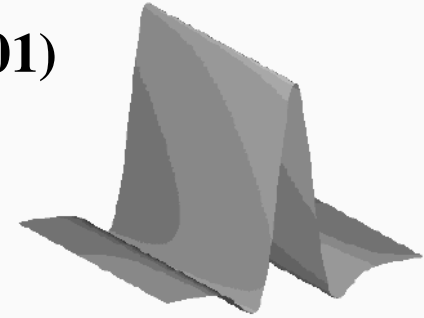
vertical bar = 0.03

inclined bar = 0.04

Width = Length<sup>2</sup>

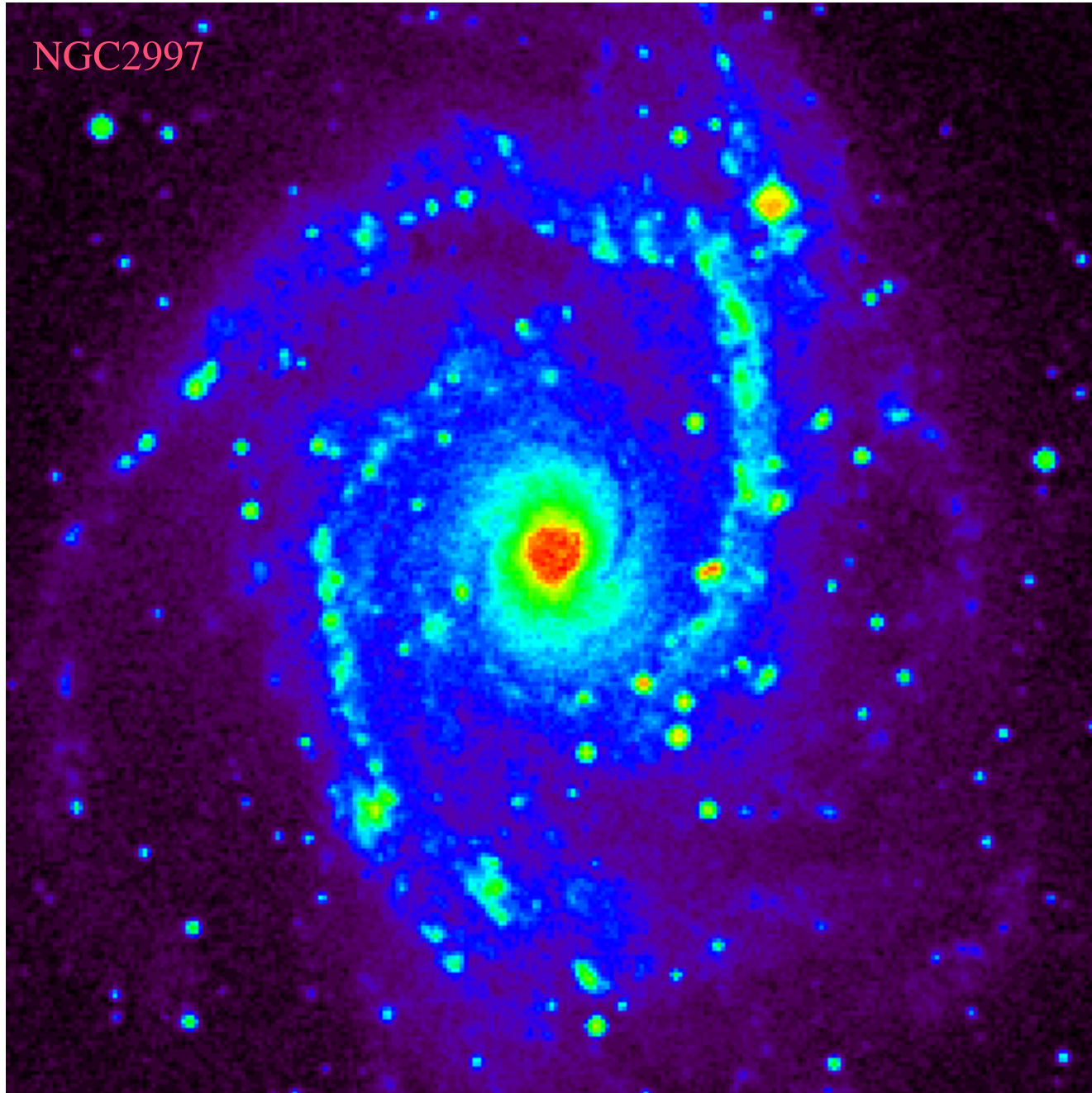
# The Curvelet Transform (CUR01)

*J.-L. Starck, E. Candes, D.L. Donoho The Curvelet Transform for Image Denoising, IEEE Transaction on Image Processing, 11, 6*



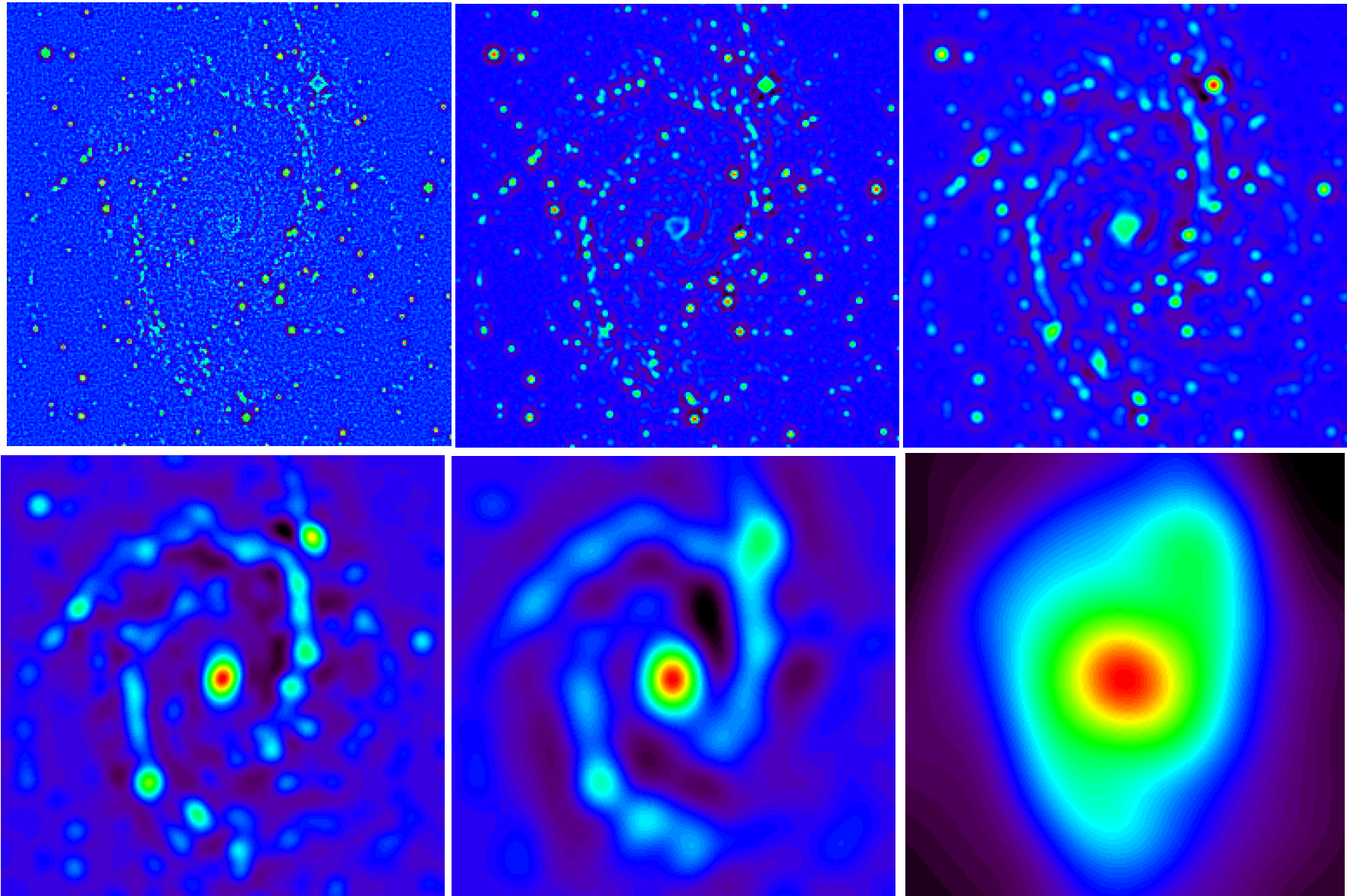
Redundancy  $16J + 1$  for  $J$  wavelet scales.  
Complexity  $O(N^2(\log N)^2)$  for  $N \times N$  images.

NGC2997



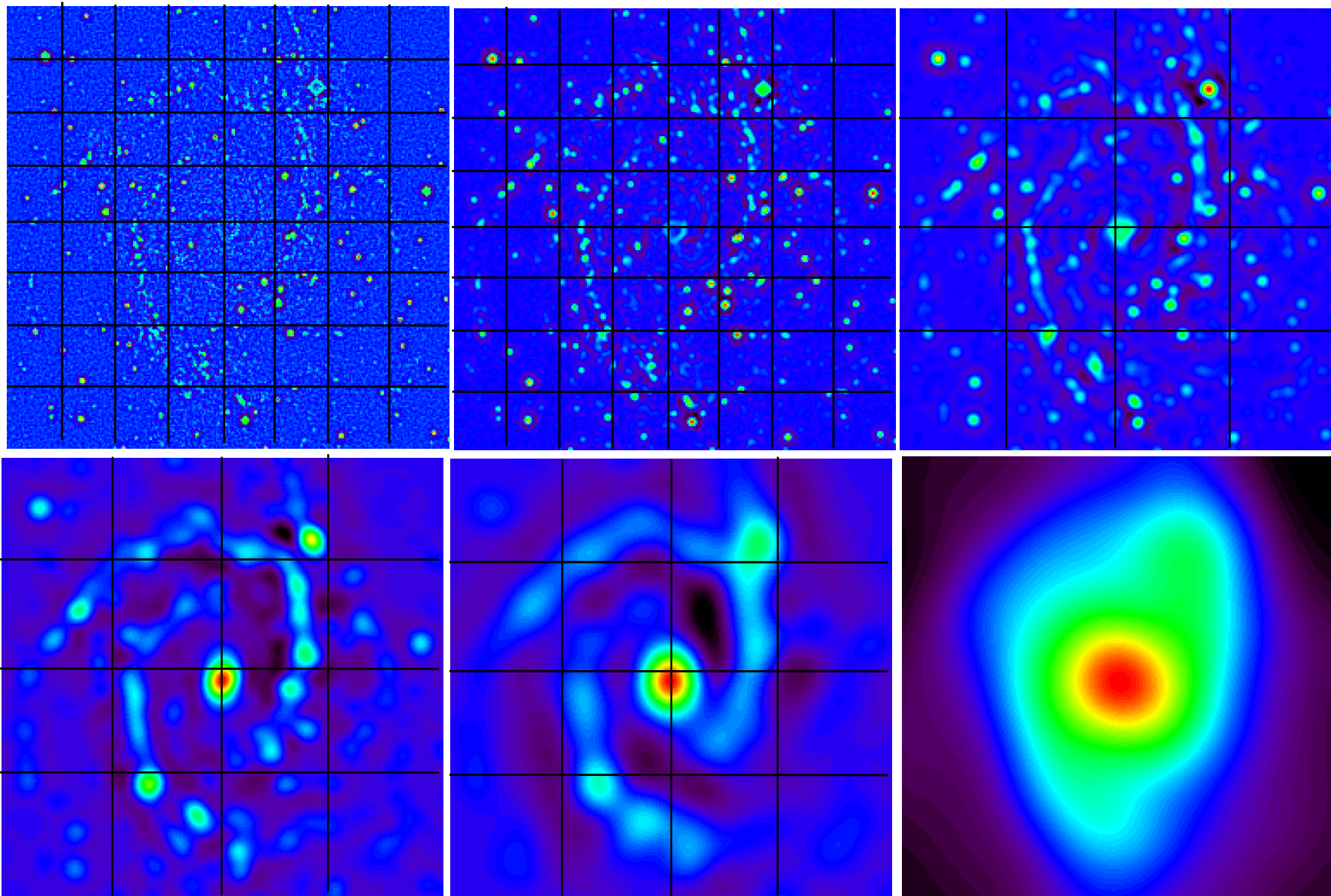
**Undecimated Isotropic WT:**

$$I(k,l) = c_{J,k,l} + \sum_{j=1}^J w_{j,k,l}$$





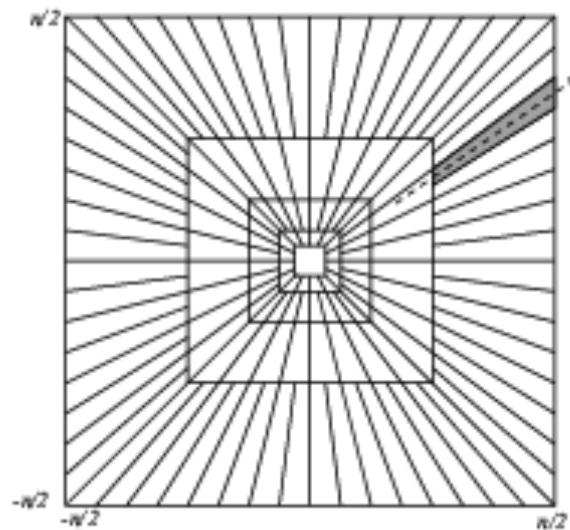
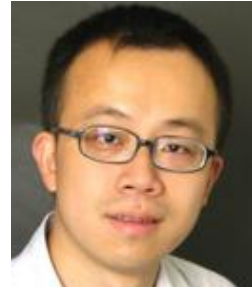
# PARTITIONING



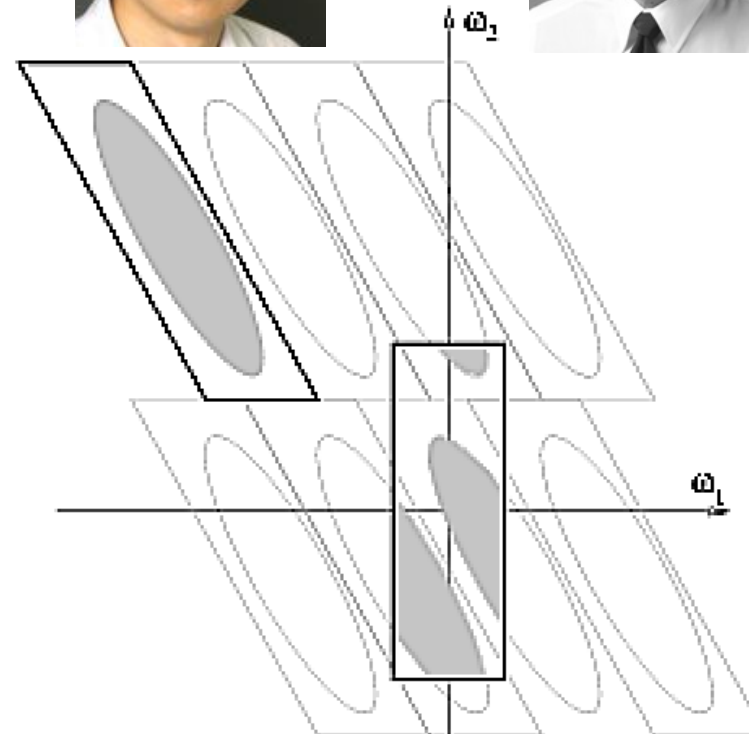
# The Fast Curvelet Transform, Candes et al, 2005

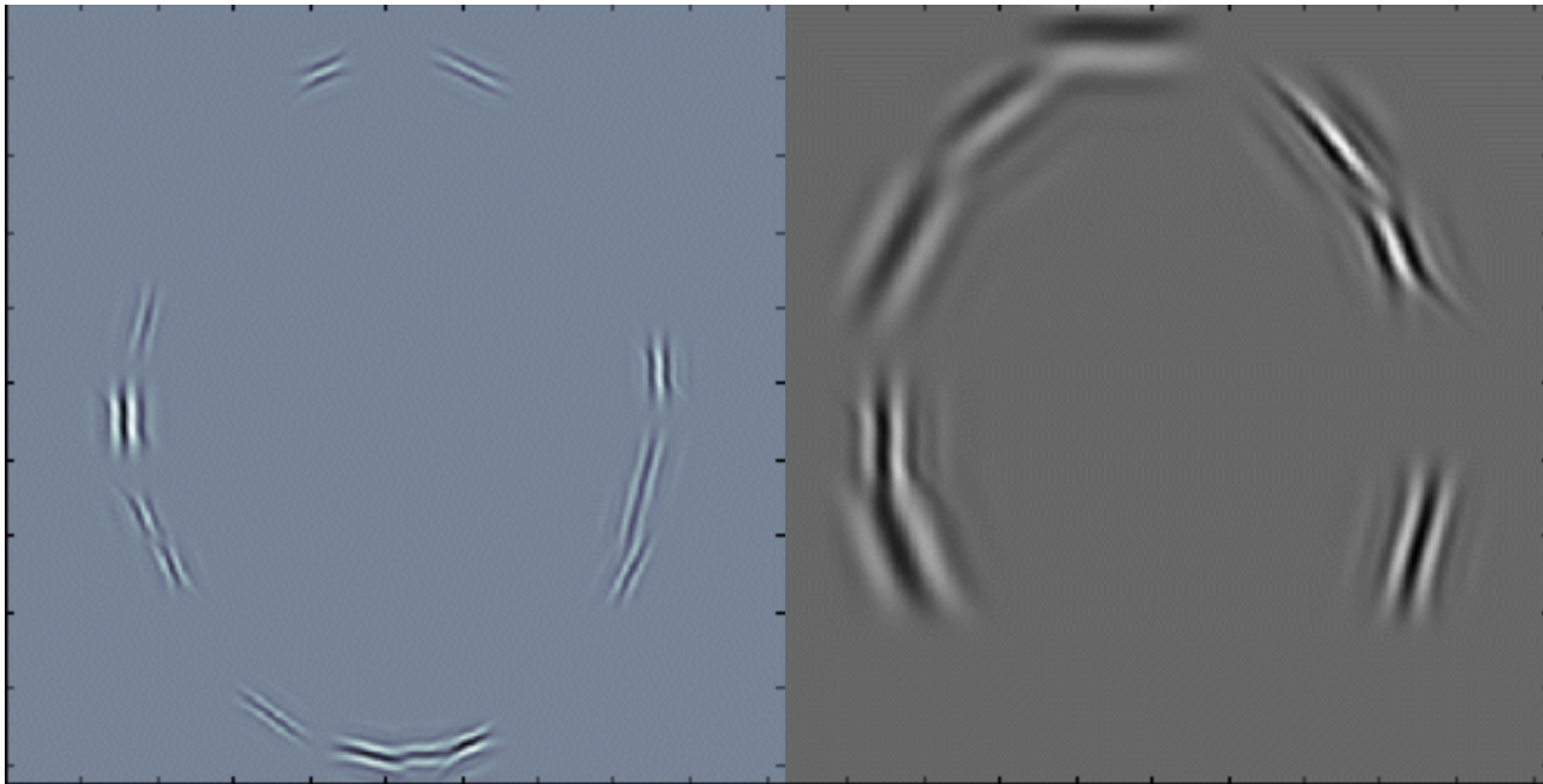
CUR03 - Fast Curvelet Transform using the USFFT

CUR04 - Fast Curvelet Transform using the Wrapping and 2DFFT



(a)





- J.-L. Starck, E. Candes, and D.L. Donoho, "**The Curvelet Transform for Image Denoising**", IEEE Transactions on Image Processing , 11, 6, pp 670 -684, 2002.
- J.-L. Starck, M.K. Nguyen and F. Murtagh, "**Wavelets and Curvelets for Image Deconvolution: a Combined Approach**", Signal Processing, 83, 10, pp 2279-2283, 2003.
- J.-L. Starck, E. Candes, and D.L. Donoho, "**Astronomical Image Representation by the Curvelet Transform**", Astronomy and Astrophysics, 398, 785--800, 2003.
- J.-L. Starck, F. Murtagh, E. Candes, and D.L. Donoho, "**Gray and Color Image Contrast Enhancement by the Curvelet Transform**", IEEE Transaction on Image Processing, 12, 6, pp 706--717, 2003.



# CONTRAST ENHANCEMENT USING THE CURVELET TRANSFORM

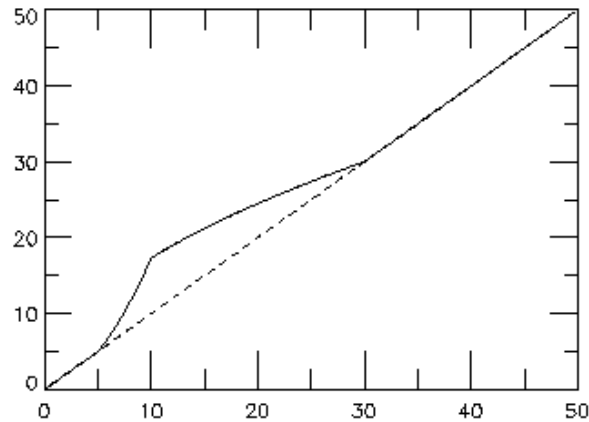
J.-L. Starck, F. Murtagh, E. Candes and D.L. Donoho, "Gray and Color Image Contrast Enhancement by the Curvelet Transform",

IEEE Transaction on Image Processing, 12, 6, 2003.

$$\tilde{I} = C_R(y_c(C_T I))$$

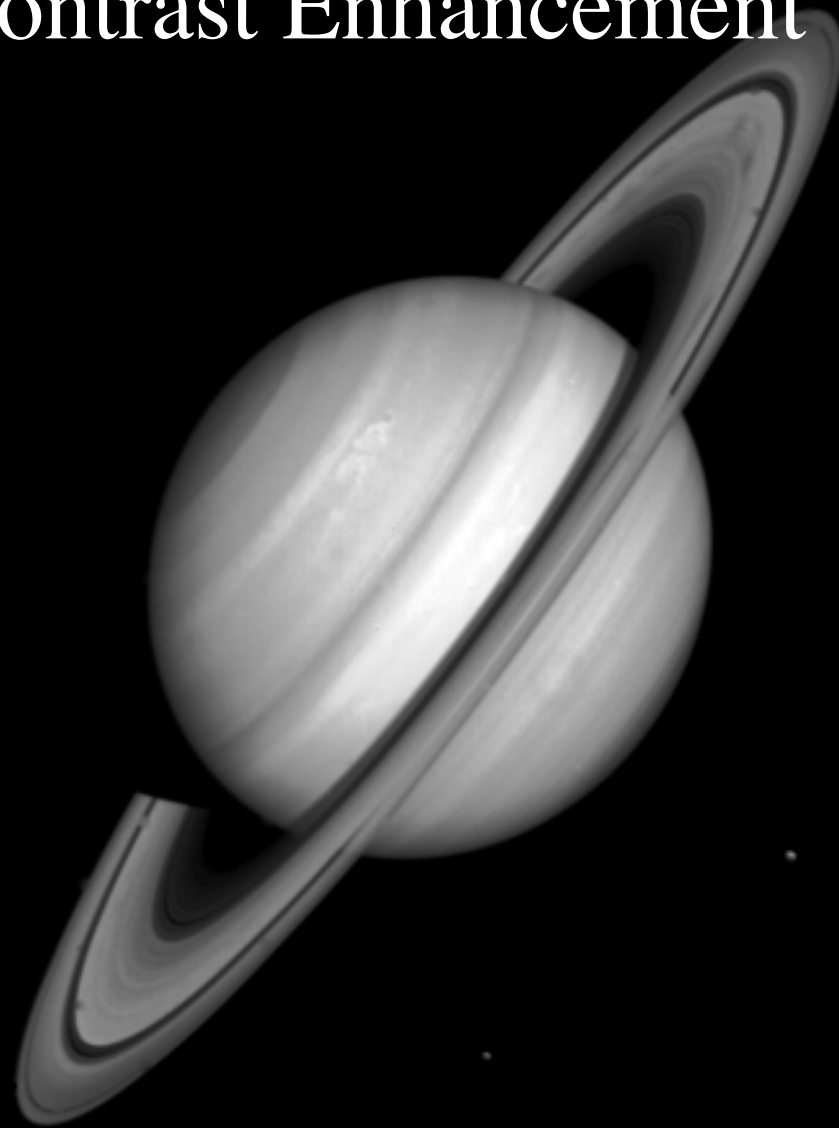
$$\left\{ \begin{array}{ll}
 y_c(x, \sigma) = 1 & \text{if } x < c\sigma \\
 y_c(x, \sigma) = \frac{x - c\sigma}{c\sigma} \left(\frac{m}{c\sigma}\right)^p + \frac{2c\sigma - x}{c\sigma} & \text{if } x < 2c\sigma \\
 y_c(x, \sigma) = \left(\frac{m}{x}\right)^p & \text{if } 2c\sigma \leq x < m \\
 y_c(x, \sigma) = \left(\frac{m}{x}\right)^s & \text{if } x > m
 \end{array} \right.$$

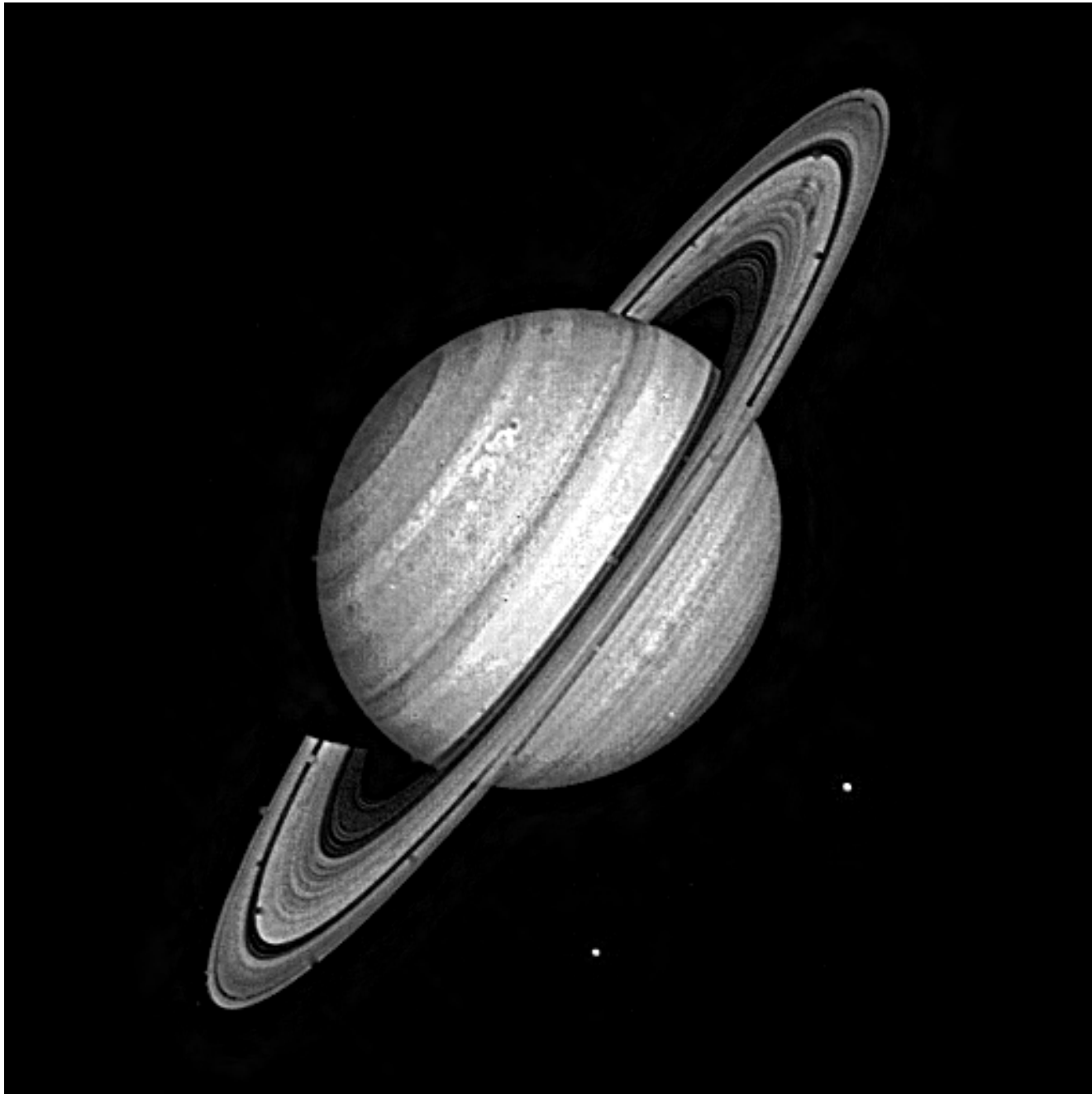
*Modified  
curvelet  
coefficient*



*Curvelet coefficient*

# Contrast Enhancement

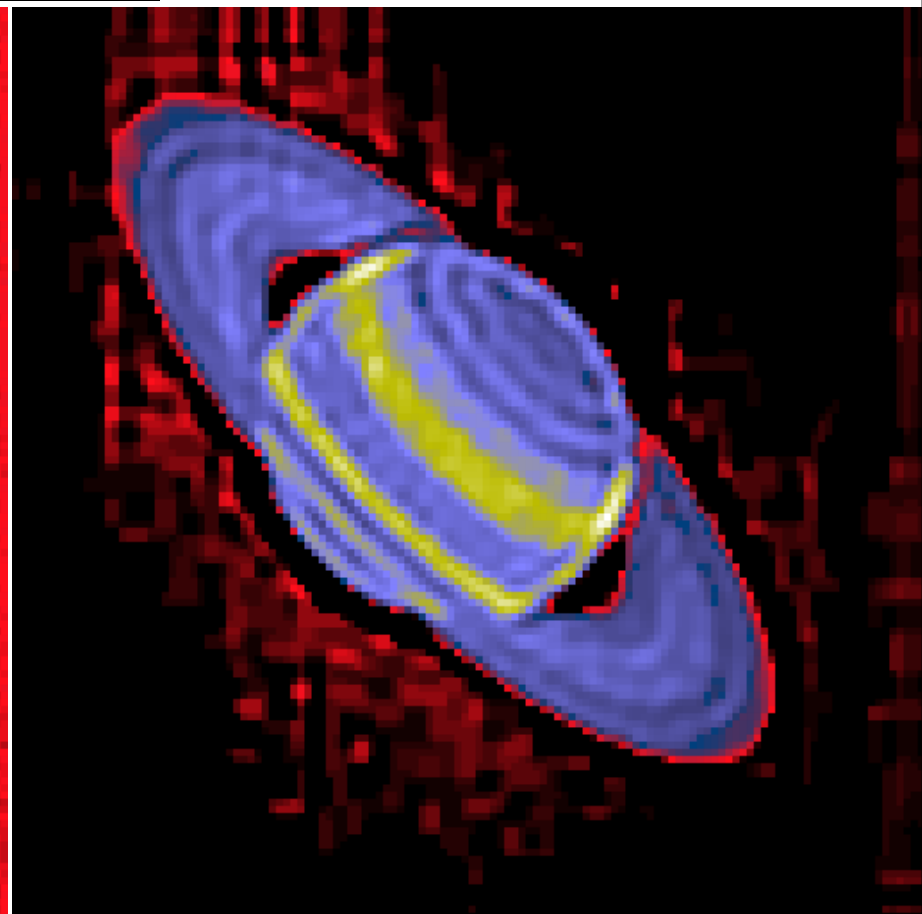
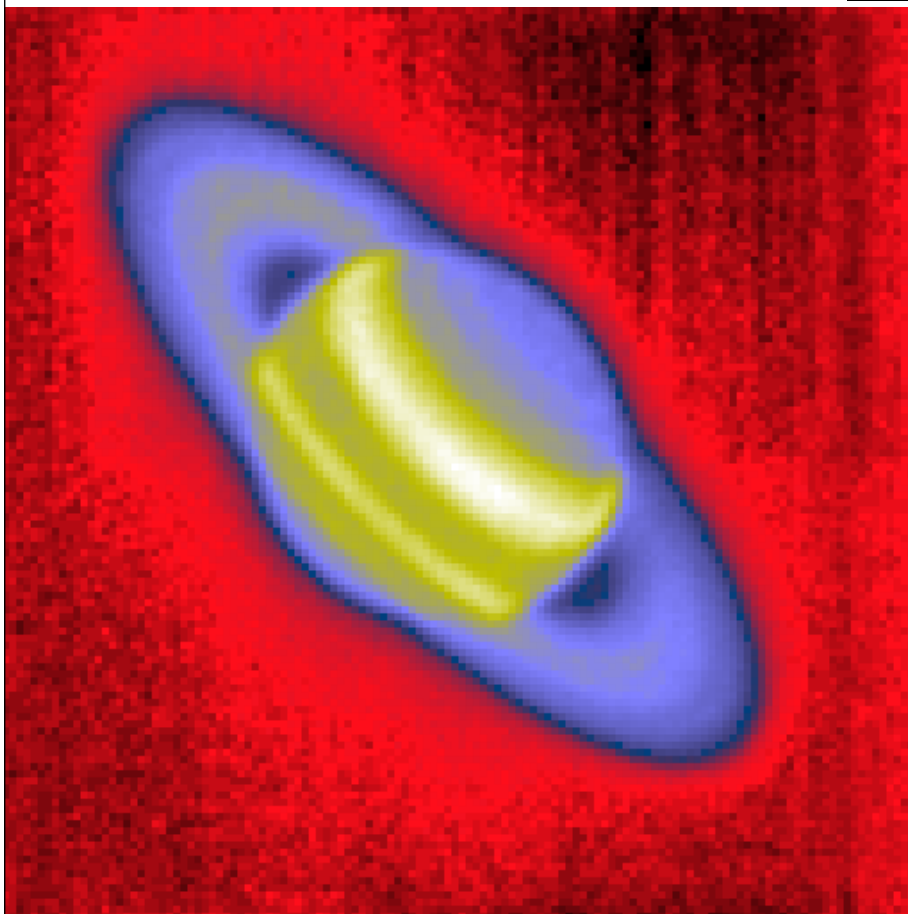
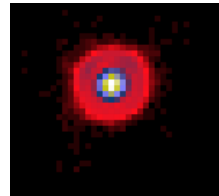






# DECONVOLUTION

- E. Pantin, J.-L. Starck, and F. Murtagh, "Deconvolution and Blind Deconvolution in Astronomy", in *Blind image deconvolution: theory and applications*, pp 277--317, 2007.
- J.-L. Starck, F. Murtagh, and M. Bertero, "The Starlet Transform in Astronomical Data Processing: Application to Source Detection and Image Deconvolution", Springer, *Handbook of Mathematical Methods in Imaging*, in press, 2011.



# PSNR

$$\left( = -10 \log_{10} \left( \frac{\text{Variance}(\text{Error})}{255^2} \right) \right)$$

	Lena sig=20	Lena sig=42	Nimes = 10
DATA	22.09	15.66	28.15
DWT	28.31	24.42	29.09
Contourlet(1.5)	28.61	25.42	
CUR03 (4)	30.73	27.71	
CUR04 (4)	30.91	27.92	31.49
PWT(4)	30.56	26.99	31.62
Complex WT (4)	30.97	27.46	31.45
UWT	31.36	28.66	32.68
CUR01	31.51	28.74	32.60
CUR01+ SSR	31.65	28.83	

}	<b>CUR01+UWT</b>	<b>32.11</b>	<b>28.90</b>
	CUR04+UWT	32.02	28.82

→ ***Very High Quality Image Restoration***, in *Signal and Image Processing IX, San Diego, 1-4 August, 2001*,  
Eds Laine, Andrew F.; Unser, Michael A.; Aldroubi, Akram, Vol. 4478, pp 9-19, 2001.



# 3D Multiscale Geometric Transforms

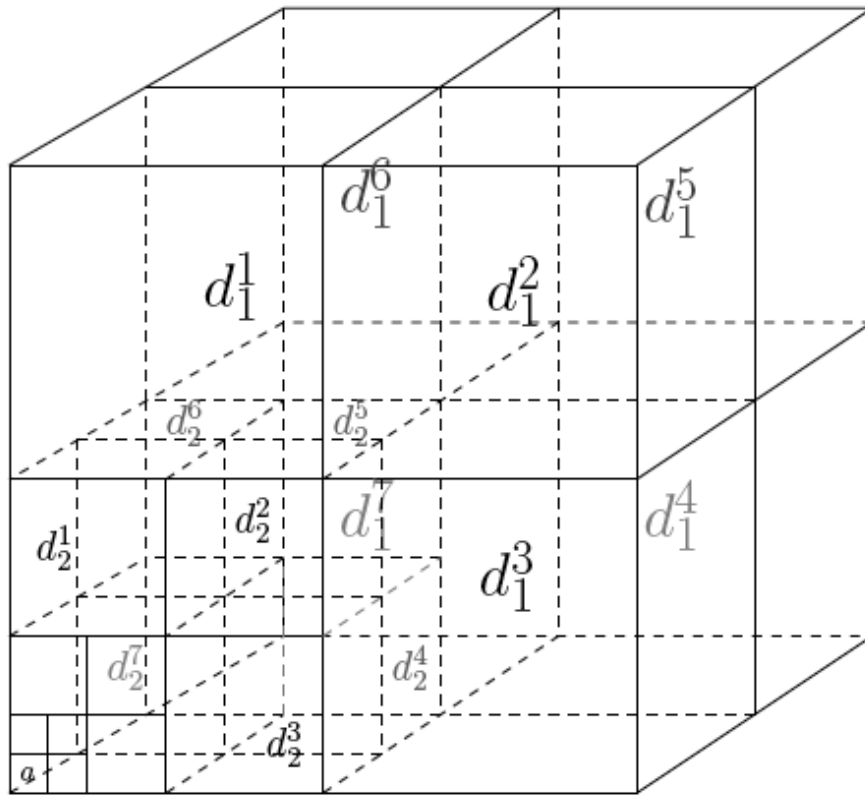
A. Waiselle, J.L. Starck and M.J. Fadili, "[3D curvelet transforms and astronomical data restoration](#)", **Applied and Computational Harmonic Analysis**, Vol. 28, No. 2, pp. 171-188, 2010.

A. Waiselle, J.L. Starck, M.J. Fadili, "[3D Data Denoising and Inpainting with the Fast Curvelet transform](#)", **JMIV**, in press, 2011.

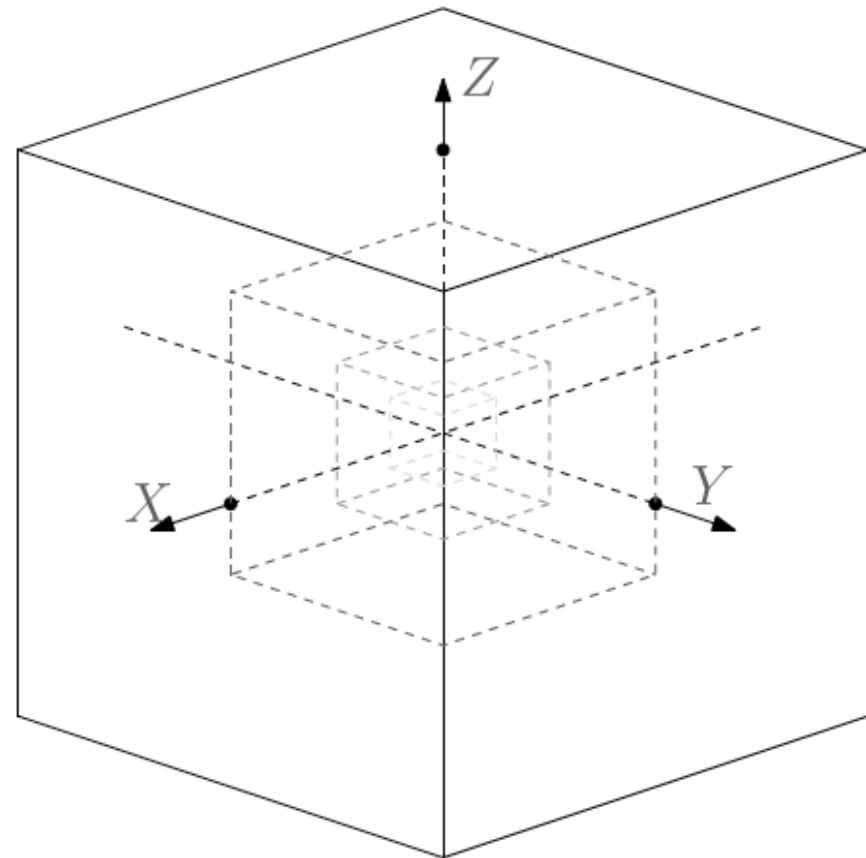
Software: <http://jstarck.free.fr/cur3d.html>

Curvelet 0 | 2D  $\implies$  3D  
FastCurvelet 3D

# 3D Wavelets



Orthogonal Wavelets



Meyer Wavelets

Pavage de l'espace de Fourier pour deux transformées en ondelettes tridimensionnelles.

# 3D extension of Curvelet

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- As in 2D, the 3D first generation curvelet transform we develop is based on the **3D ridgelet** transform applied to **localized blocks** of the output of a **3D wavelet transform**.
- The essential ingredient is the **projection slice theorem**: the  $m$ -D FT of the projection of a  $d$ -D function onto an  $m$ -D linear submanifold is equal to an  $m$ -D central slice of the  $d$ -D FT parallel to the submanifold.
- Two 3D extensions of the ridgelet transform:
  - Projections along **lines** (3D partial Radon transform,  $d=3$ ,  $m=2$ ): **BeamCurvelets**.
  - Projecting along **planes** (3D Radon transform,  $d=3$ ,  $m=1$ ): **RidCurvelets**.

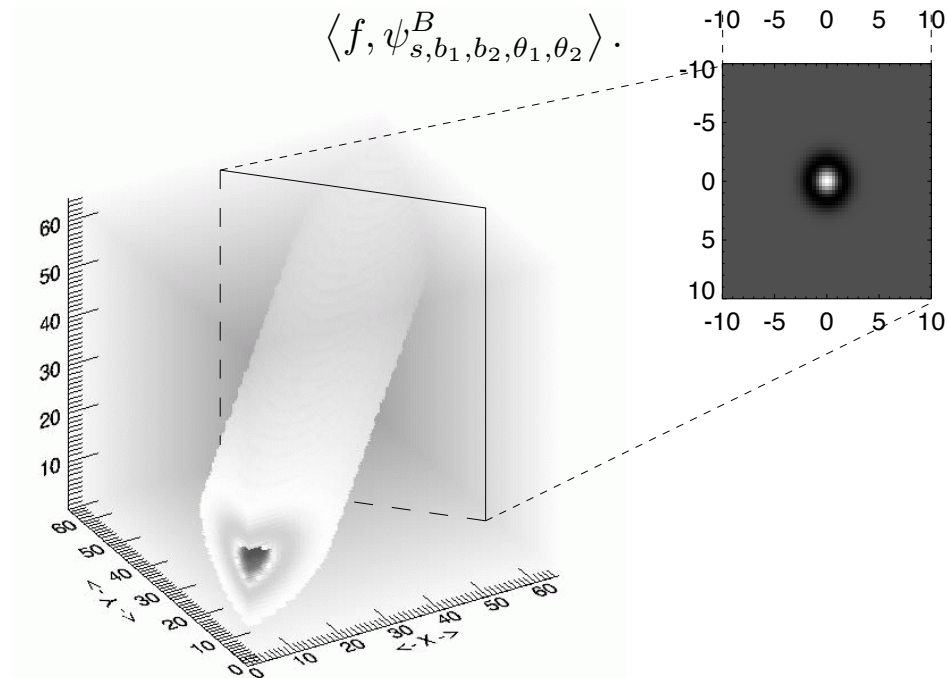


# 3D beamlet transform

- $\gamma \in L_2(\mathbb{R}^2)$  with zero-mean and has sufficient decay (2D wavelet).
- For each scale  $s > 0$ , position  $(b_1, b_2) \in \mathbb{R}^2$  and orientation  $(\theta_1, \theta_2) \in [0, 2\pi) \times [0, \pi)$ , define the 2D beamlet  $\psi_{s,b_1,b_2,\theta_1,\theta_2}^B : \mathbb{R}^3 \rightarrow \mathbb{R}$  by

$$\psi_{s,b_1,b_2,\theta_1,\theta_2}^B(\mathbf{x}) = s^{-1/2} \cdot \gamma\left(\frac{-x \sin \theta_1 + y \cos \theta_1 - b_1}{s}, \frac{(x \cos \theta_1 \sin \theta_2 + y \sin \theta_1 \sin \theta_2 - z \cos \theta_2 - b_2)}{s}\right).$$

- The 3D beamlet transform of  $f(\mathbf{x})$ ,  $\mathbf{x} \in \mathbb{R}^3$  is the set of coefficients



# 3D BeamCurvelet transform

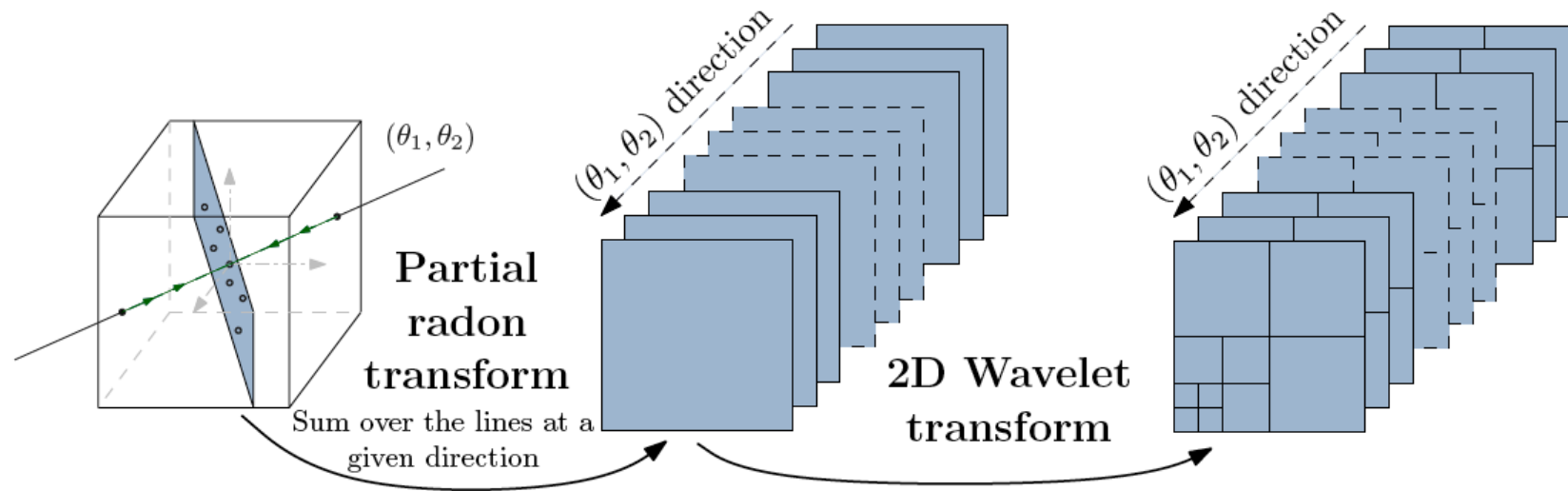
- The 3D BeamCurvelet transform of  $f(\mathbf{x})$ :

$$\langle (T_{\mathbf{Q}_{j,\mathbf{k}}})^{-1} w_{\mathbf{Q}_{j,\mathbf{k}}} \Delta_j(f), \psi_{s,b_1,b_2,\theta_1,\theta_2}^B \rangle = \langle f, \Delta_j(w_{\mathbf{Q}_{j,\mathbf{k}}} T_{\mathbf{Q}_{j,\mathbf{k}}} \psi_{s,b_1,b_2,\theta_1,\theta_2}^B) \rangle ,$$

- spatial scale  $j$ ,
- spatial location  $\mathbf{k} = (k_x, k_y, k_z)$ ,
- ridge scale  $s$ ,
- ridge location  $(b_1, b_2)$ ,
- angular location  $(\theta_1, \theta_2)$ .

A. Woiselle, J.L. Starck and M.J. Fadili, "[3D curvelet transforms and astronomical data restoration](#)", **Applied and Computational Harmonic Analysis**, Vol. 28, No. 2, pp. 171-188, 2010.

# Discrete 3D Beamlet transform



# Discrete 3D BeamCurvelet transform

**Algorithm:** Fourier-based implementation.

**Data:** A data cube and a block size  $B$ .

**Result:** BeamCurvelet transform.

**begin**

    Apply a 3D isotropic wavelet transform.

**for**  $j = 1$  to  $J$  **do**

        Smooth partition of the subband into block cubes of size  $B$ .

**for each block do**

            Apply a 3D FFT.

            Extract planes passing through the origin at every angle  $(\theta_1, \theta_2)$ .

**for each plane**  $(\theta_1, \theta_2)$  **do**

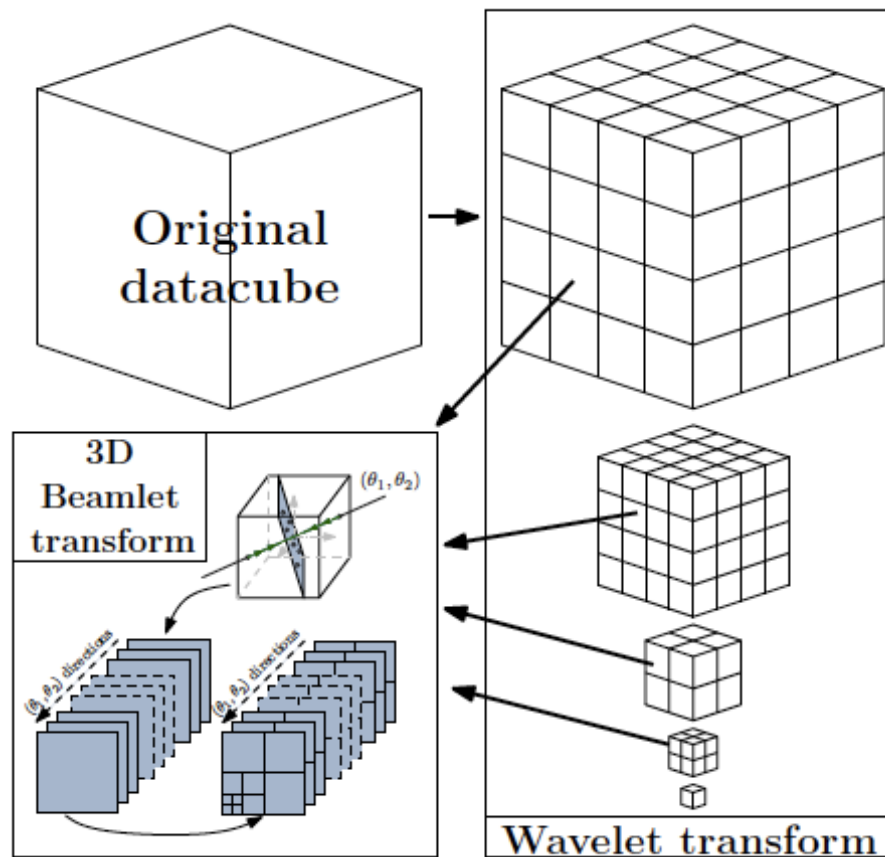
                Apply an inverse 2D FFT.

                Apply a 2D wavelet transform to get the

                BeamCurvelet coefficients.

**if**  $j$  is odd **then** according to the parabolic scaling:  $B \leftarrow 2B$ .

**end**



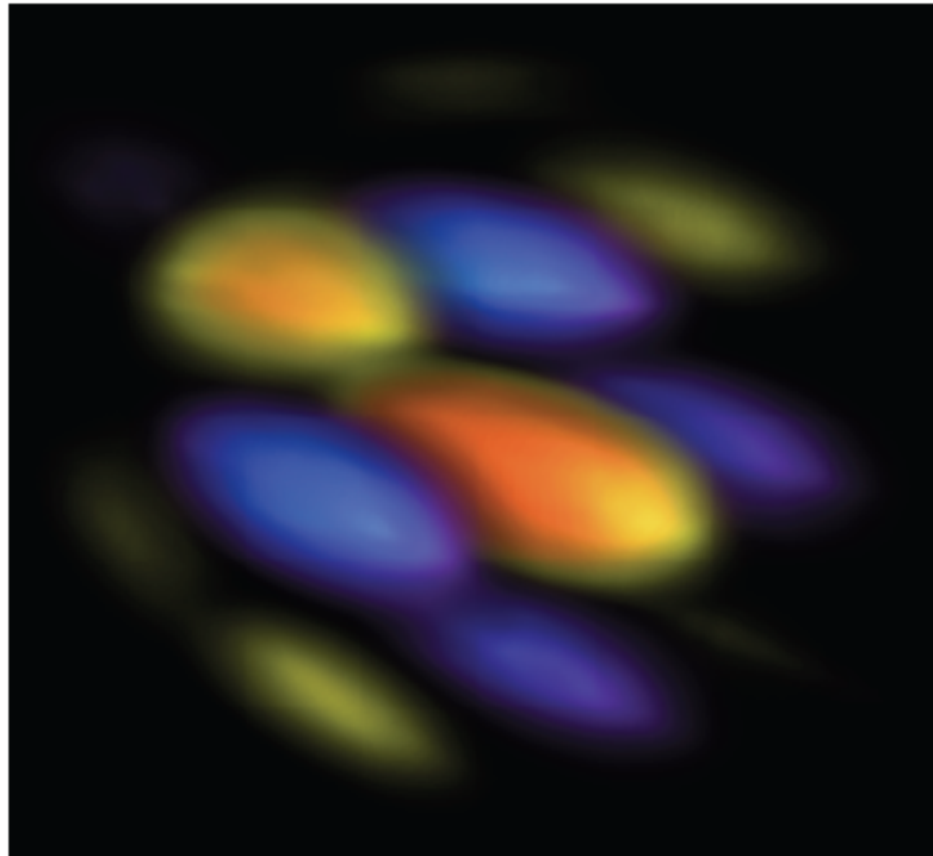
Redundancy  $\approx 3\rho^3 B$ ,  $\rho \in [1, 2]$ .

Complexity  $O(N^3(\log N)^2)$  for  $N \times N \times N$  volume.

# 3D BeamCurvelets

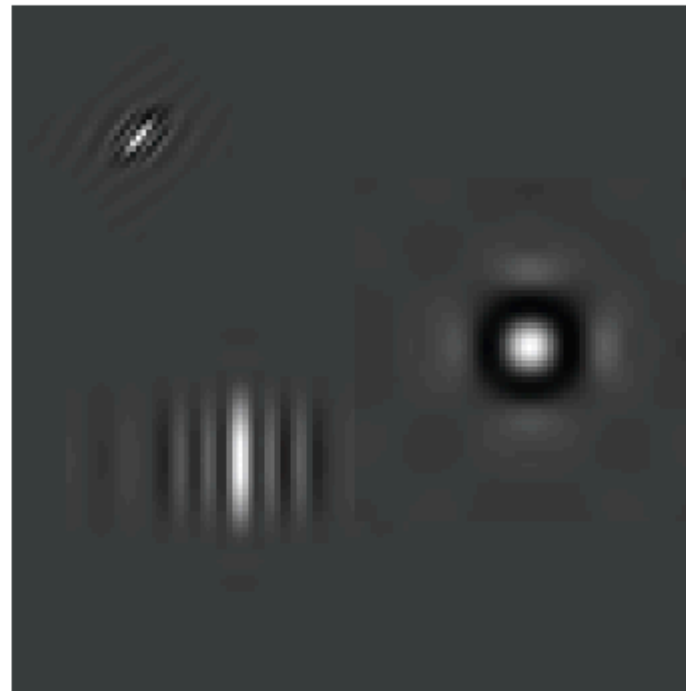
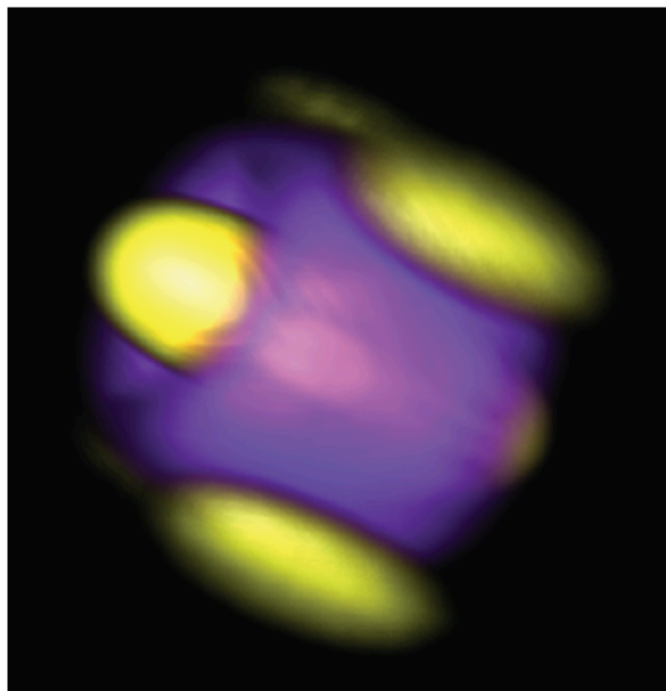
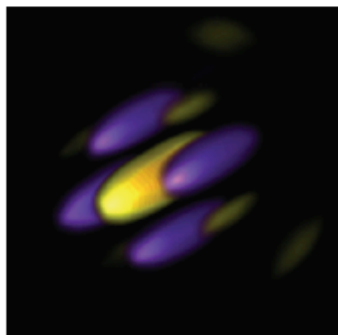
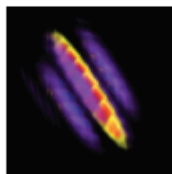
---

- It is constant along segments of direction  $(\theta_1, \theta_2)$ , and a 2D wavelet function transverse to this direction.
- Adapted to filamentary structures in 3D.



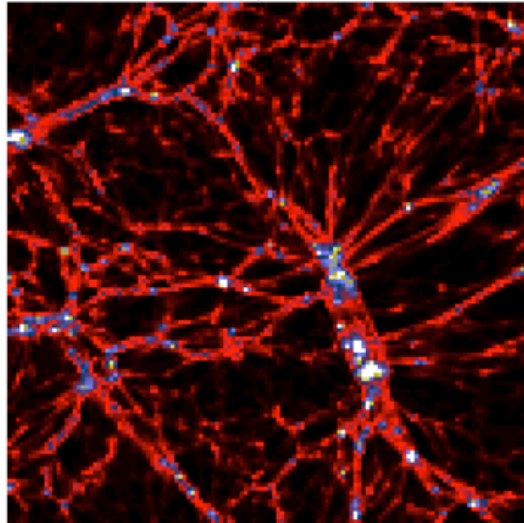
# 3D BeamCurvelets

---

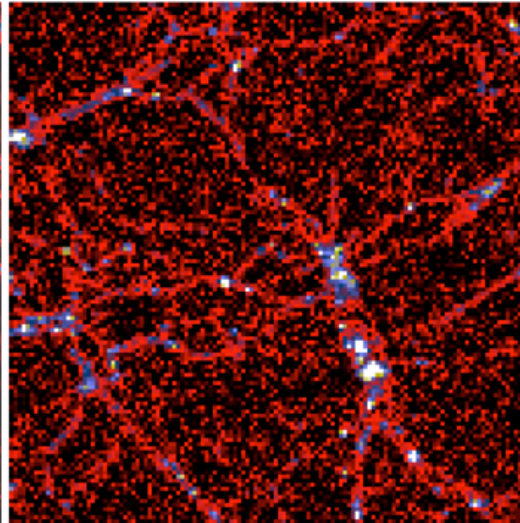


# 3D BeamCurvelets

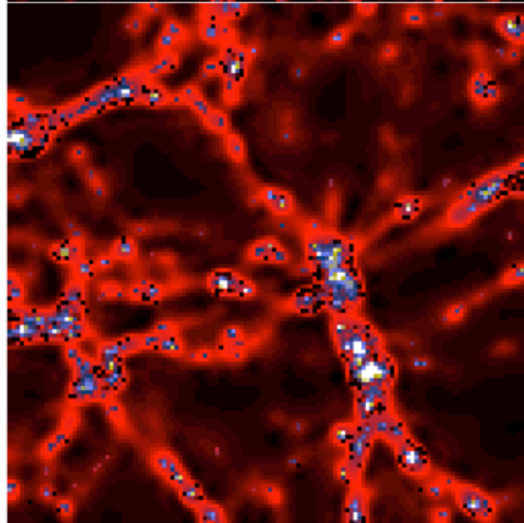
Simulation



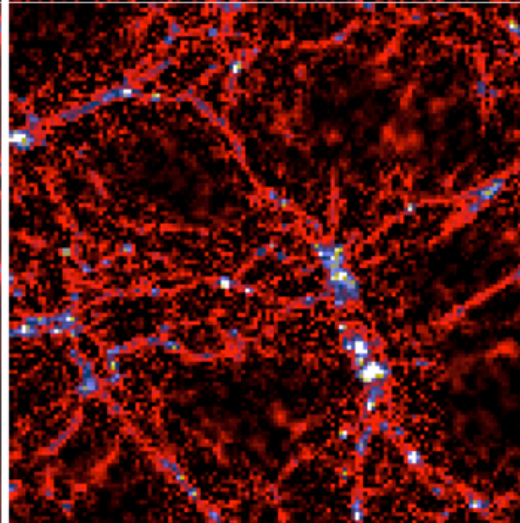
Noise Data



Wavelet  
Thresholding



BeamCurvelet  
Thresholding

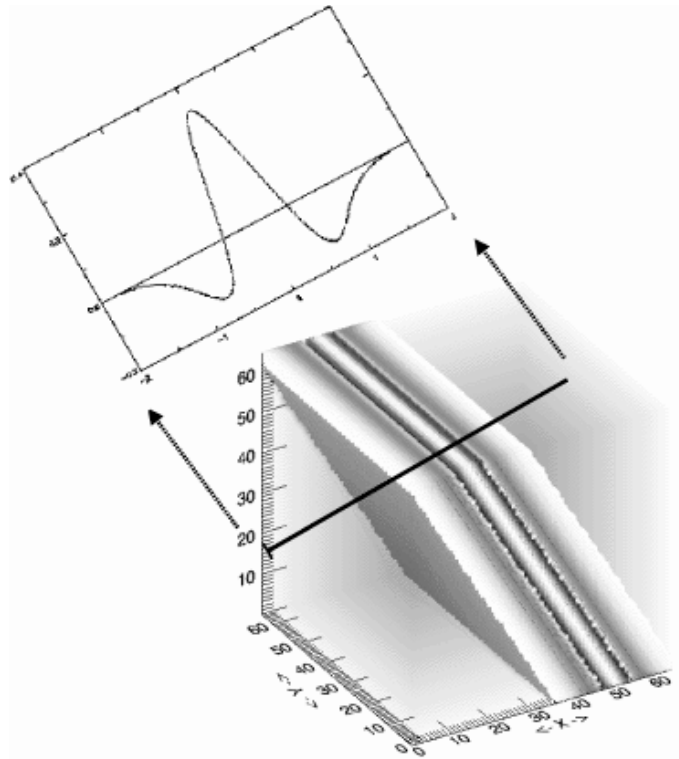


# 3D ridgelet transform

- $\psi \in L_2(\mathbb{R})$  with zero-mean and has sufficient decay.
- For each scale  $s > 0$ , position  $b \in \mathbb{R}$  and orientation  $(\theta_1, \theta_2) \in [0, 2\pi) \times [0, \pi)$ , define the 2D ridgelet  $\psi_{s,b,\theta_1,\theta_2}^R : \mathbb{R}^3 \rightarrow \mathbb{R}$  by

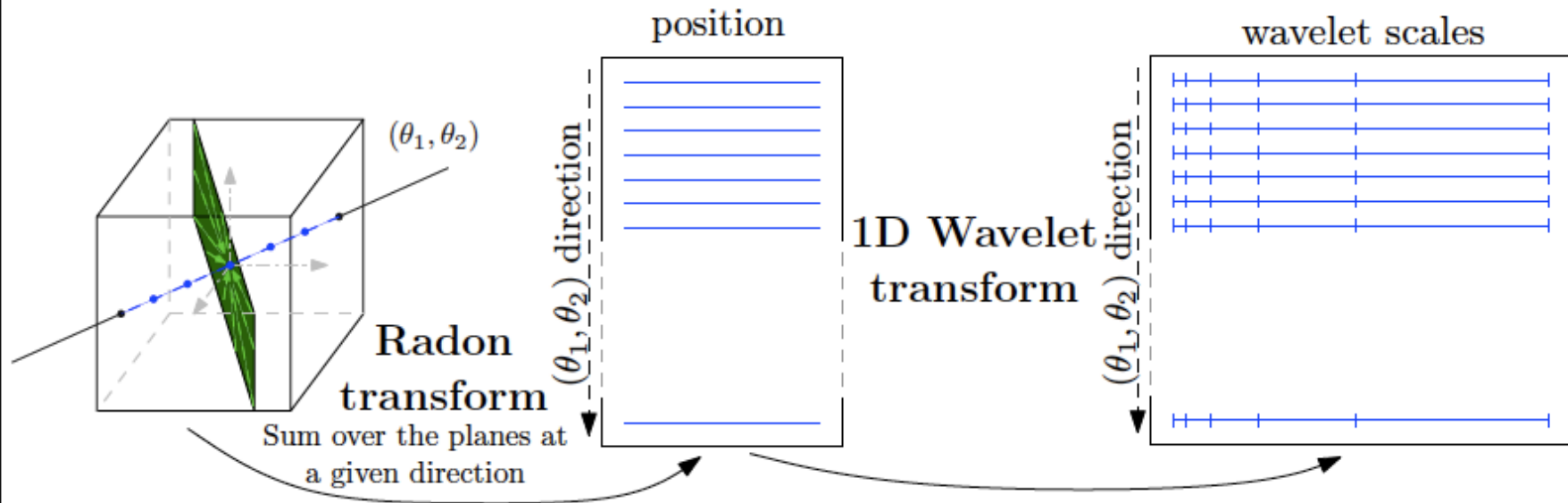
$$\psi_{s,b,\theta_1,\theta_2}^R(\mathbf{x}) = s^{-1/2} \cdot \psi((x \cos \theta_1 \cos \theta_2 + y \sin \theta_1 \cos \theta_2 + z \sin \theta_2 - b)/s).$$

- The 3D ridgelet transform of  $f(\mathbf{x})$ ,  $\mathbf{x} \in \mathbb{R}^3$  is the set of coefficients





# 3D Discrete ridgelet transform



# 3D RidCurvelet transform

- The 3D BeamCurvelet transform of  $f(\mathbf{x})$ :

$$\langle (T_{\mathbf{Q}_{j,\mathbf{k}}})^{-1} w_{\mathbf{Q}_{j,\mathbf{k}}} \Delta_j(f), \psi_{s,b,\theta_1,\theta_2}^R \rangle = \langle f, \Delta_j(w_{\mathbf{Q}_{j,\mathbf{k}}} T_{\mathbf{Q}_{j,\mathbf{k}}} \psi_{s,b,\theta_1,\theta_2}^R) \rangle ,$$

- spatial scale  $j$ ,
- spatial location  $\mathbf{k} = (k_x, k_y, k_z)$ ,
- ridge scale  $s$ ,
- ridge location  $b$ ,
- angular location  $(\theta_1, \theta_2)$ .

A. Woiselle, J.L. Starck and M.J. Fadili, "[3D curvelet transforms and astronomical data restoration](#)", **Applied and Computational Harmonic Analysis**, Vol. 28, No. 2, pp. 171-188, 2010.

# Discrete 3D RidCurvelet transform

**Algorithm:** Fourier-based implementation.

**Data:** A data cube and a block size  $B$ .

**Result:** RidCurvelet transform.

**begin**

Apply a 3D isotropic wavelet transform.

**for**  $j = 1$  **to**  $J$  **do**

Smooth partition of the subband into block cubes of size  $B$ .

**for each block do**

Apply a 3D FFT.

Extract lines passing through the origin at every angle  $(\theta_1, \theta_2)$ .

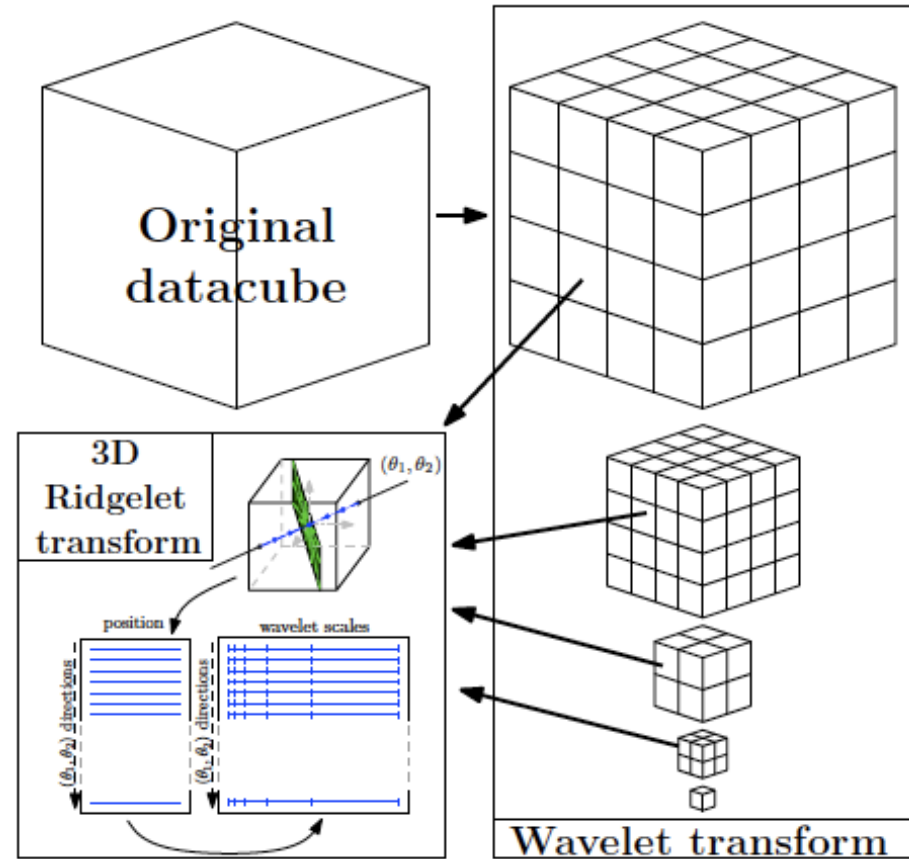
**for each line**  $(\theta_1, \theta_2)$  **do**

Apply an inverse 1D FFT.

Apply a 1D wavelet transform to get the RidCurvelet coefficients.

**if**  $j$  **is odd then** according to the parabolic scaling:  $B \leftarrow 2B$ .

**end**

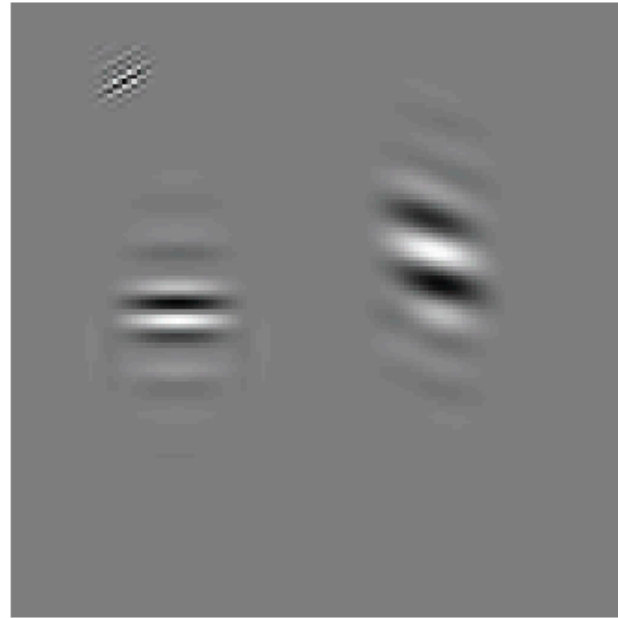
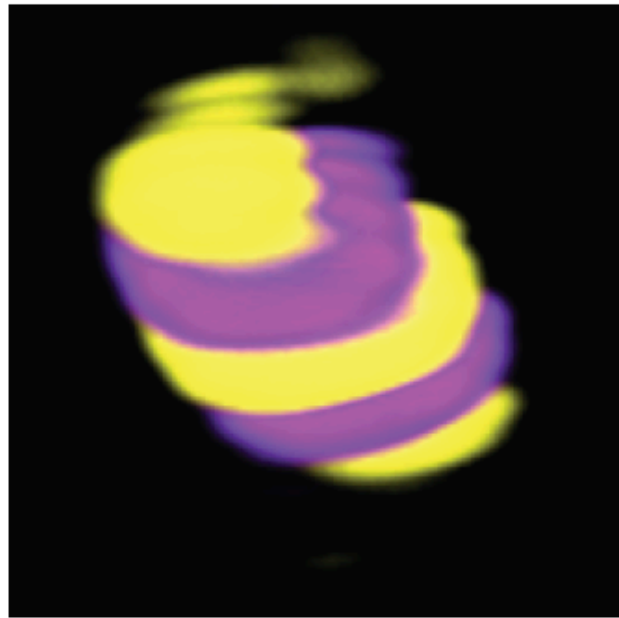


$$\text{Redundancy} \approx 6\rho^3, \rho \in [1, 2].$$

$$\text{Complexity } O(N^3(\log N)^2) \text{ for } N \times N \times N \text{ volume.}$$

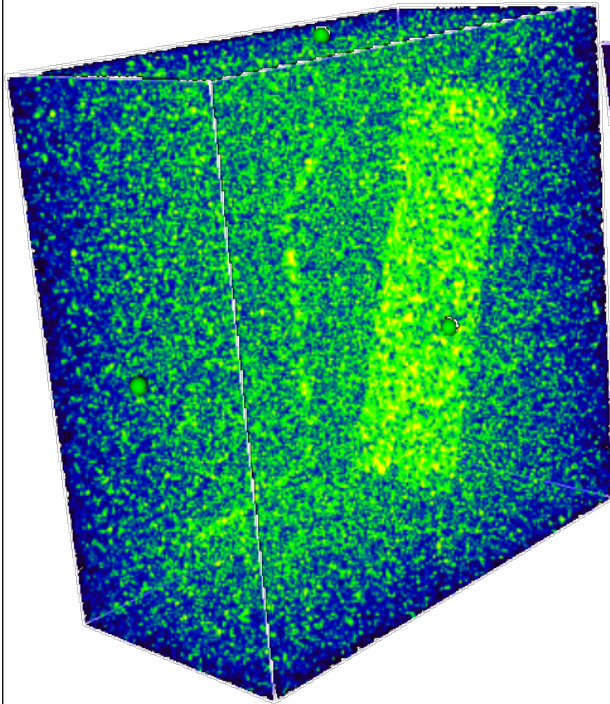
# 3D RidCurvelets

---

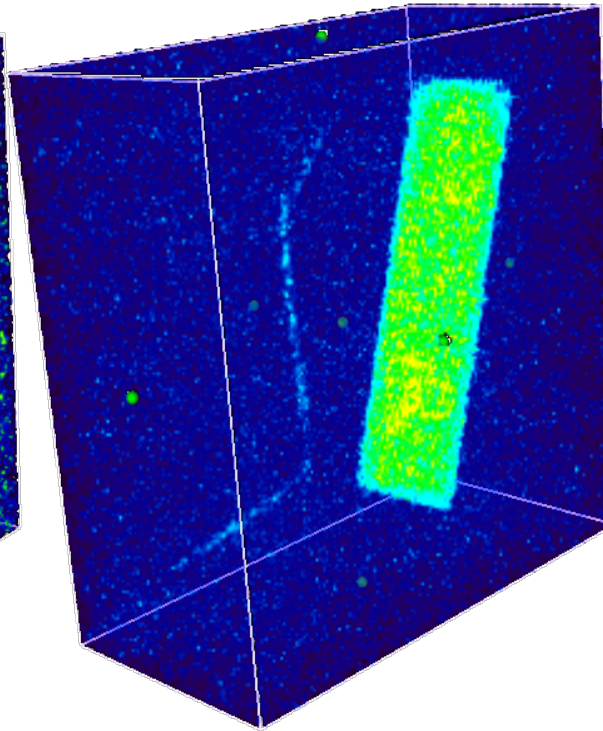


# Denoising

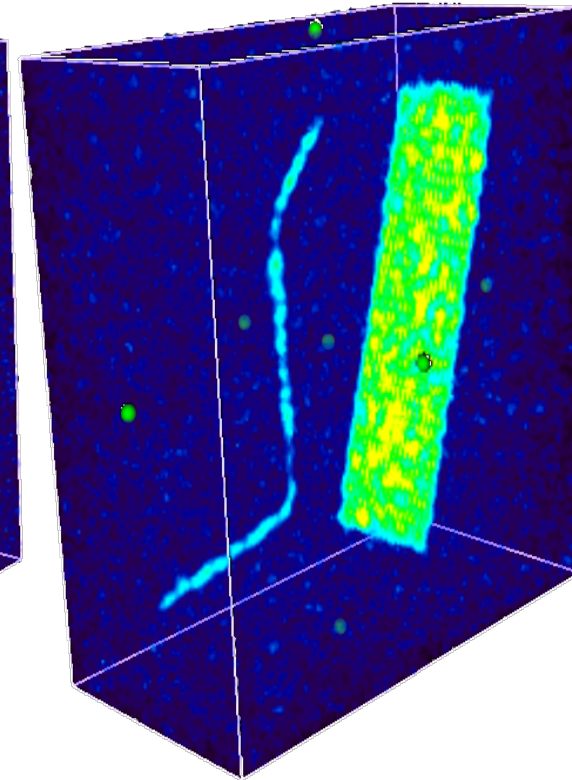
---



Noisy data

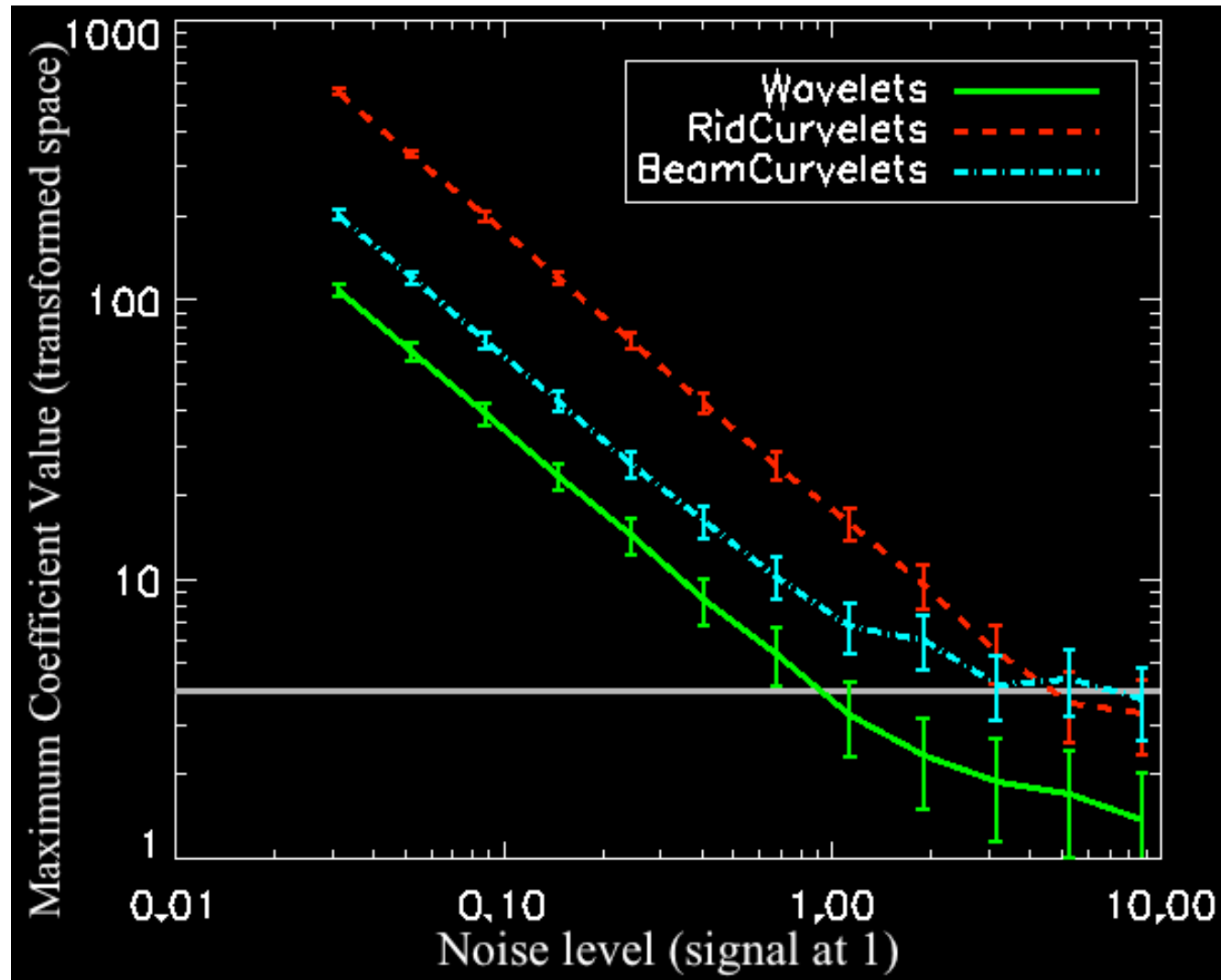


Denoised with RidCurvelets

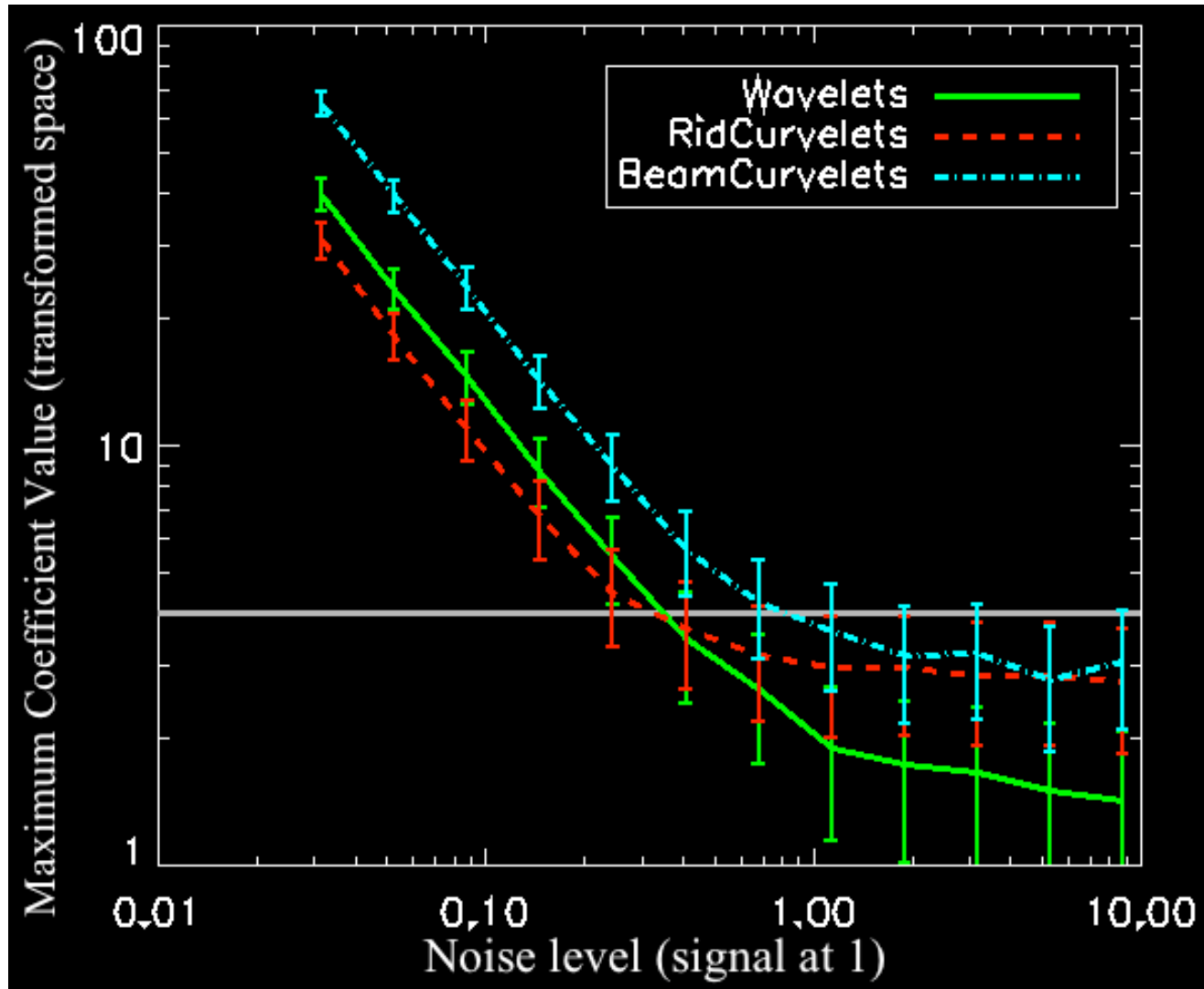


Denoised with BeamCurvelets

# 3D Plane detection level



# 3D Line detection level



# Combined denoising

- Amalgamate several transforms in a single dictionary  $\Phi = [\Phi_1, \dots, \Phi_K]$  to benefit from the best of each transform.
- More flexibility to represent complex geometrical content: the blessing of over-completeness.

- We have to solve

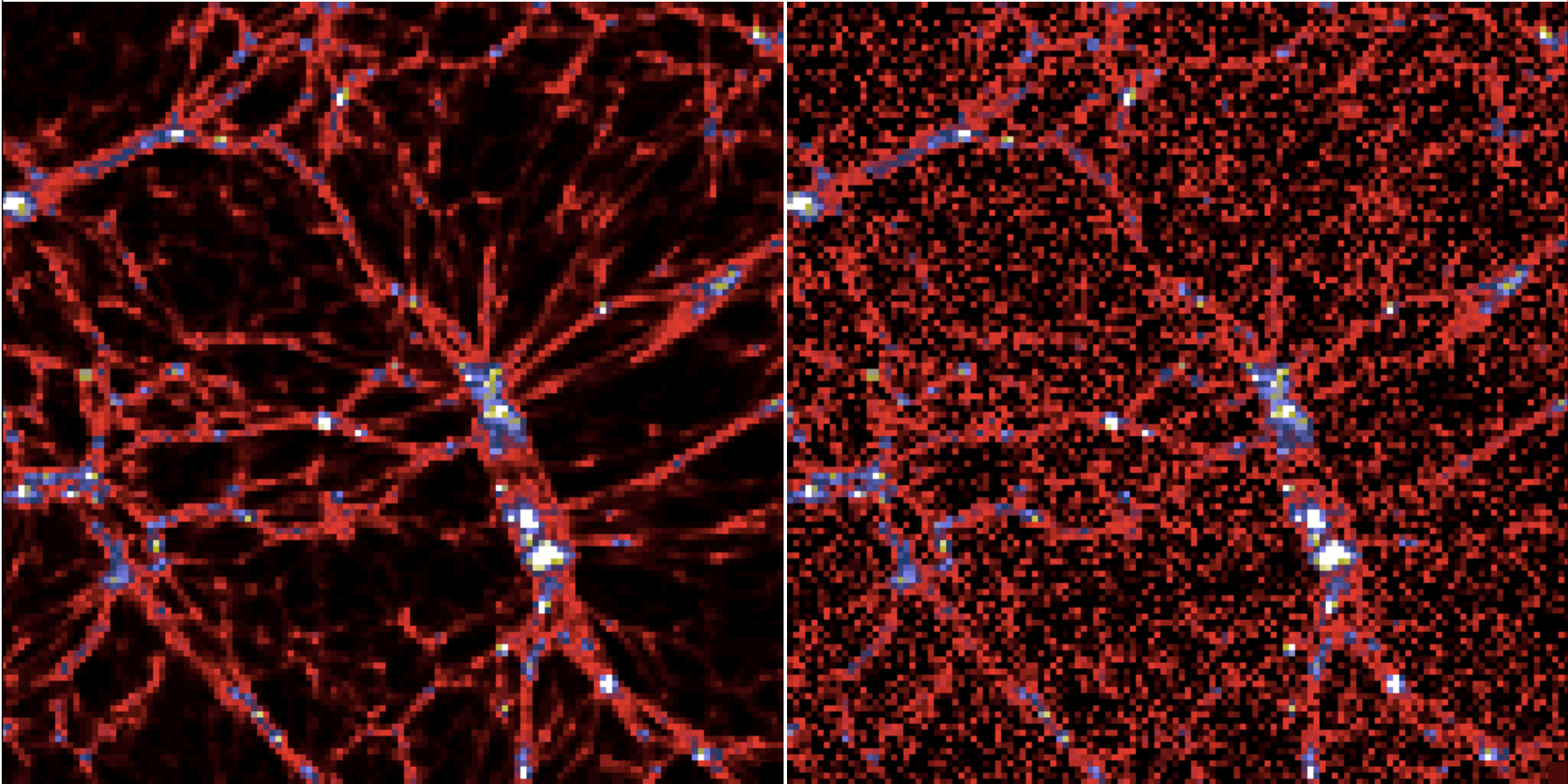
$$\min_{\alpha} \|\alpha\|_p^p \quad \text{s.t.} \quad \|g - \Phi\alpha\|_2 \leq \epsilon(\sigma), \quad 0 \leq p \leq 1.$$

- Solutions by e.g.:
  - Convex optimization (monotone operator splitting).
  - Greedy pursuit.



# Combined denoising results

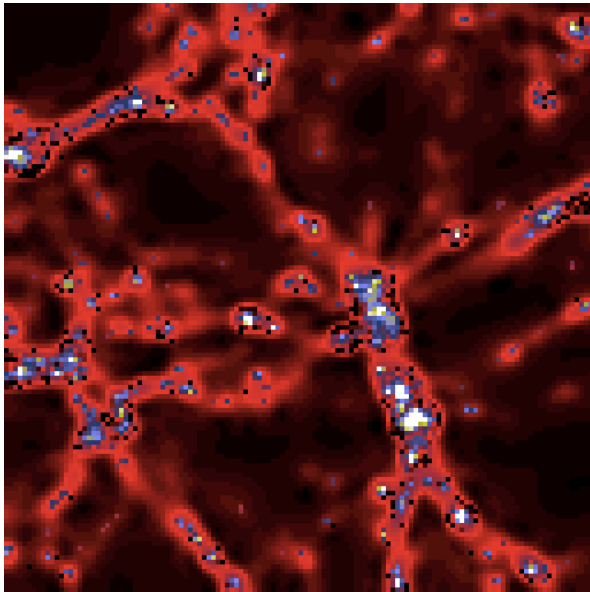
Cold Dark Matter simulations: clusters and filamentary structures with density of the filaments 3 orders of magnitude lower than the clusters.



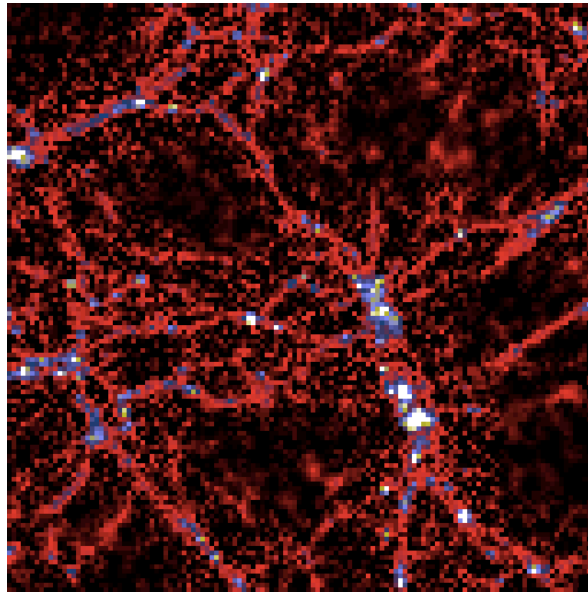
# Combined denoising results

---

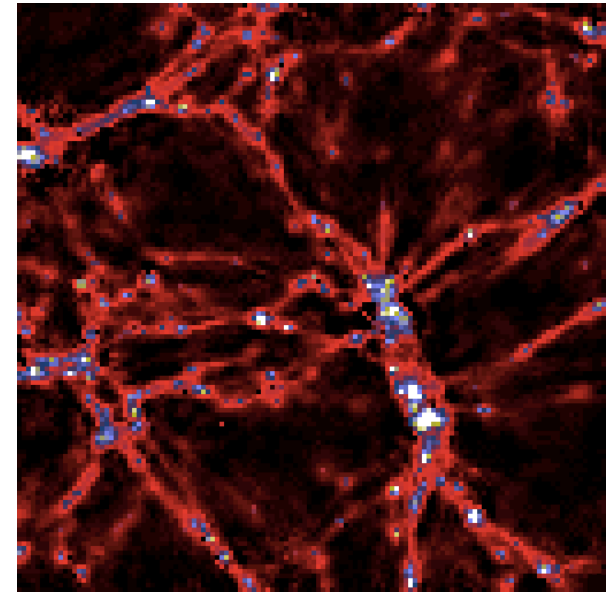
3D UDWT



BeamCurvelets

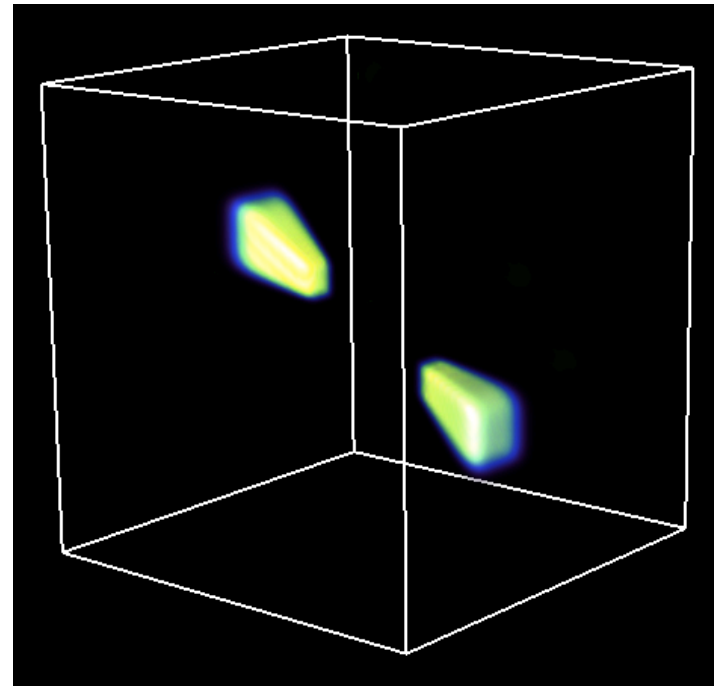
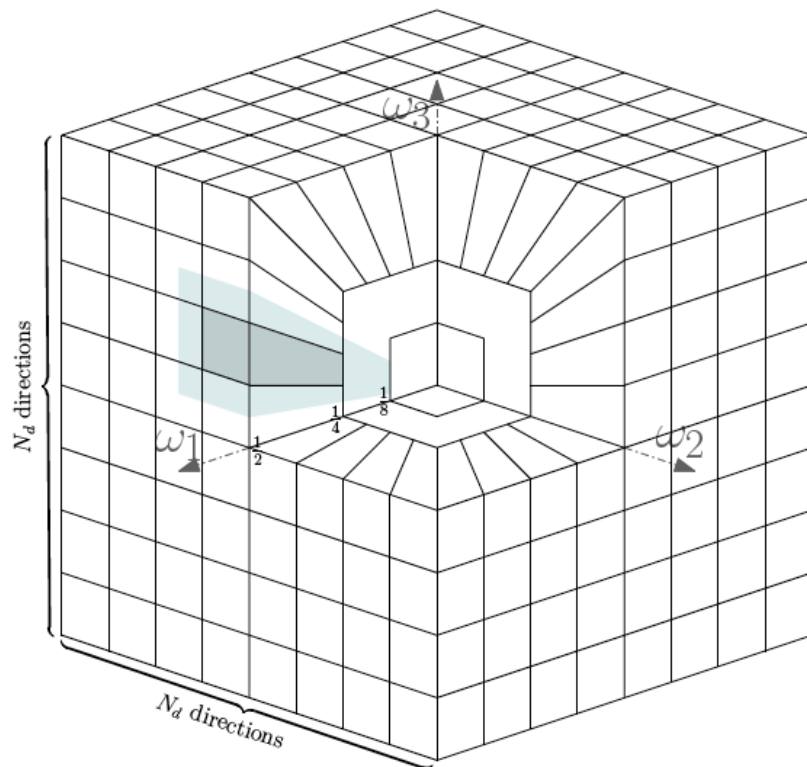


Combined denoising  
BeamCurvelets+3D UDWT



# Fast second generation 3D curvelets

- 3D Fast curvelet transform (Candes et al, 2005)
- Curvelab, a C++/Matlab toolbox available at [www.curvelet.org](http://www.curvelet.org).



Redundancy: 7.11 in 2D and 24.38 in 3D  
with walevets at the finest scale: 3.56 in 2D and 5.42 in 3D

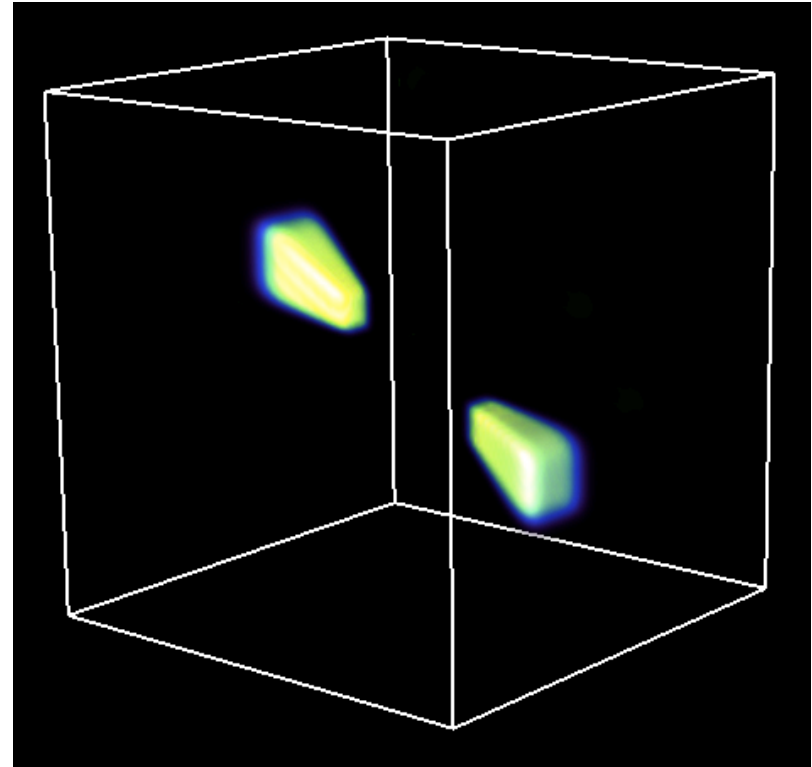
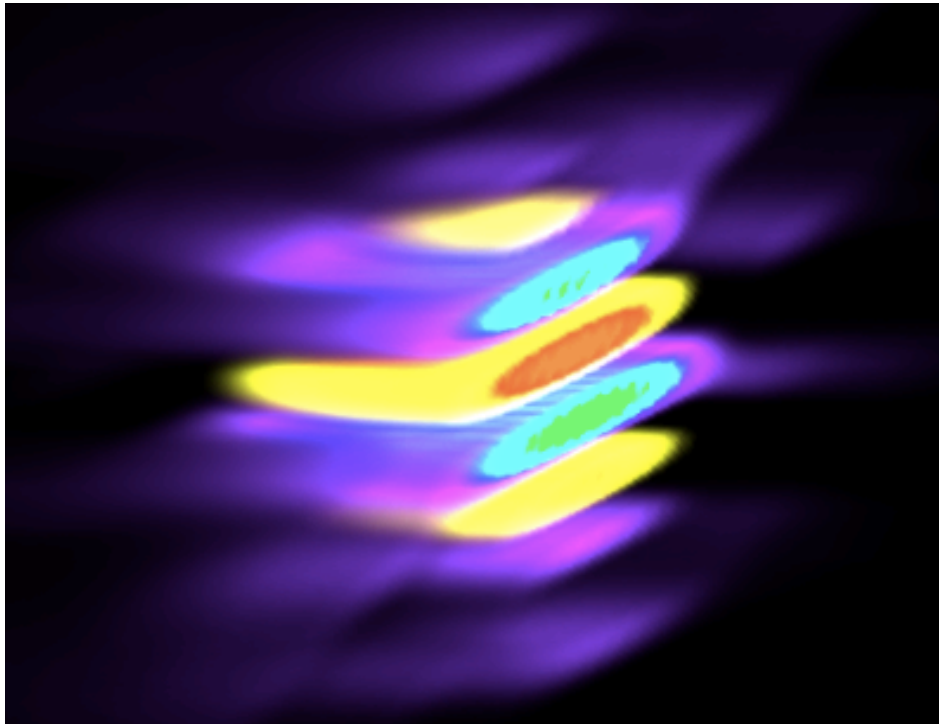
# Fast second generation 3D curvelets

- 3D Fast curvelet transform.
- Main differences with Candès et al. CurveLab :
  - Implementation: e.g. wavelet transform, overlapping angular windows.
  - Much less redundant than Candès et al. (2.3-10.3 instead of 5.4-24.4).
  - Faster in practice.

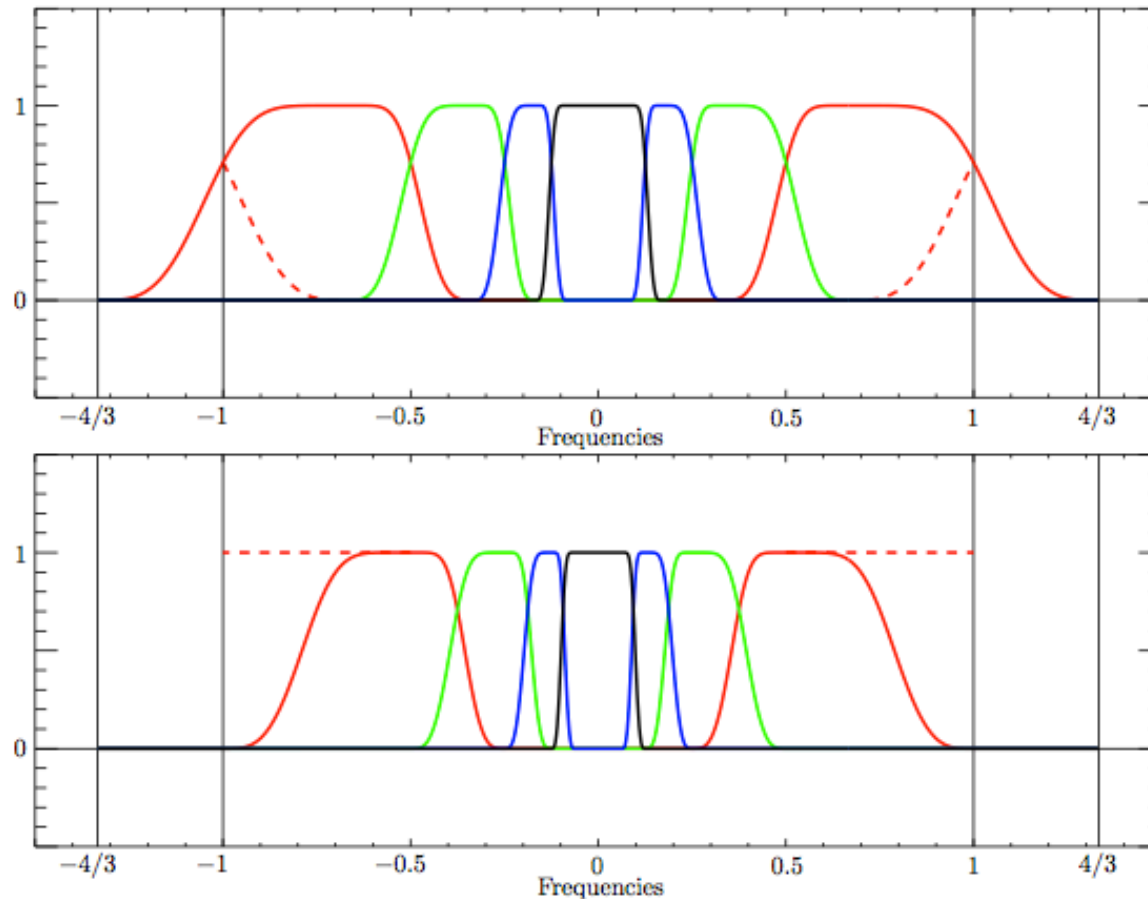
	Original FCT		Proposed FCT	
	C	W	C	W
2-D	7.11	3.56	4.00	2.00
3-D	24.38	5.42	10.29	2.29

# Fast second generation 3D curvelets

---

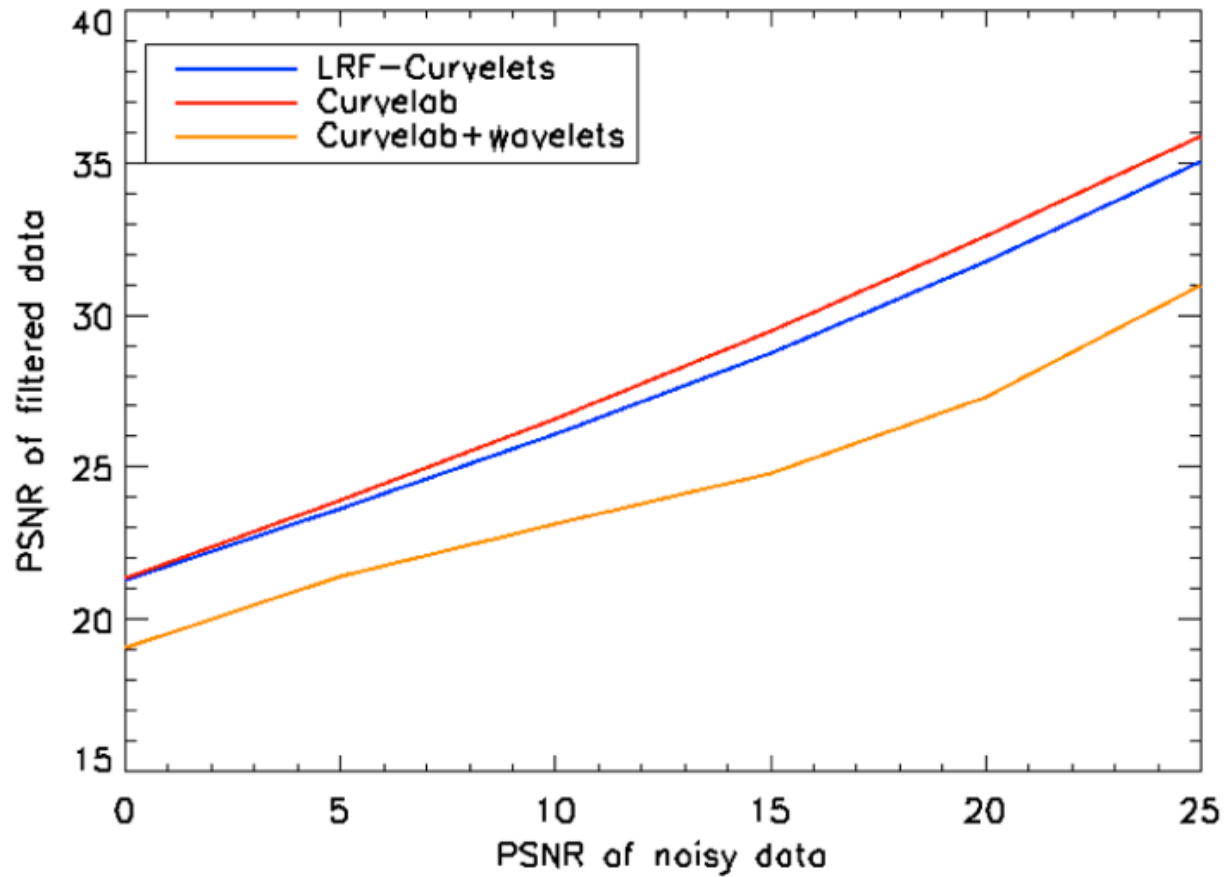


# Fast second generation 3D curvelets



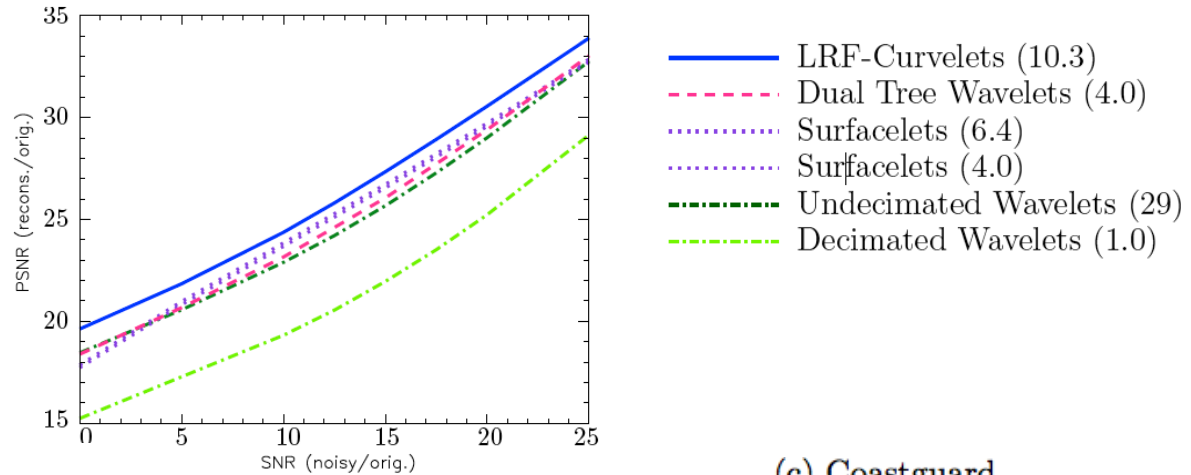
Meyer wavelets functions in Fourier domain. In the discrete case, we only have access to the Fourier samples inside the Shannon band  $[-1/2, 1/2]$ , while the wavelet corresponding to the finest scale (solid red line) exceeds the Shannon frequency band to  $2/3$ . Top : In the Curvelab implementation, the Meyer wavelet basis is periodized in Fourier, so that the exceeding end of the finest scale wavelet is replaced with the mirrored dashed line on the plot. Bottom : In our implementation, the wavelets are shrunk so that they fit in the  $[-1/2, 1/2]$  Shannon band, and the decreasing tail of the finest scale wavelet is replaced by a constant (dashed red line) to ensure a uniform partition of the unity.

# Example: denoising

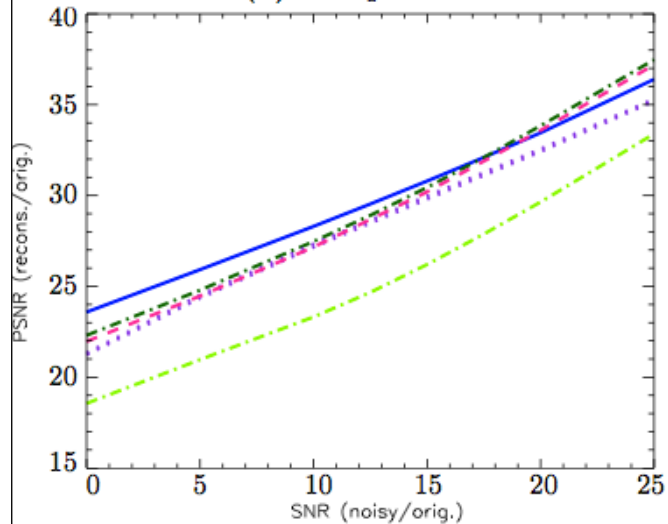


Mean denoising PSNR versus noise level using different FCT implementations. The denoising PSNR was averaged over ten noise realizations and several datasets. The LR-FCT is in blue. Original FCT implementation of Curvelab using curvelets (red) and wavelets (orange) at the finest scale.

# Example: video denoising

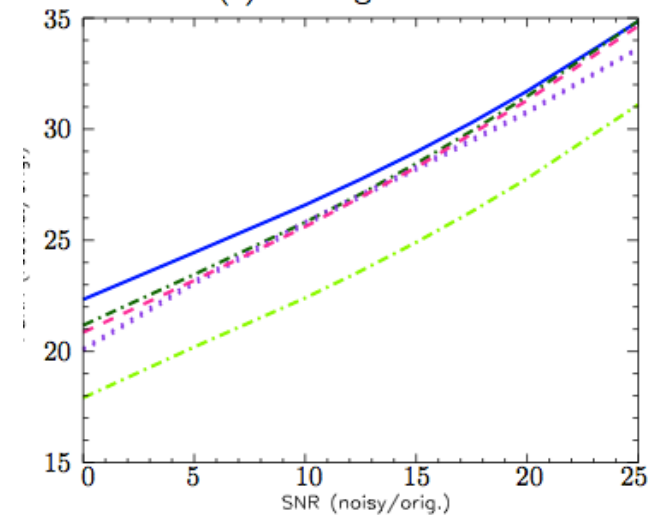


(b) Tempete



(a) Mobile

(c) Coastguard



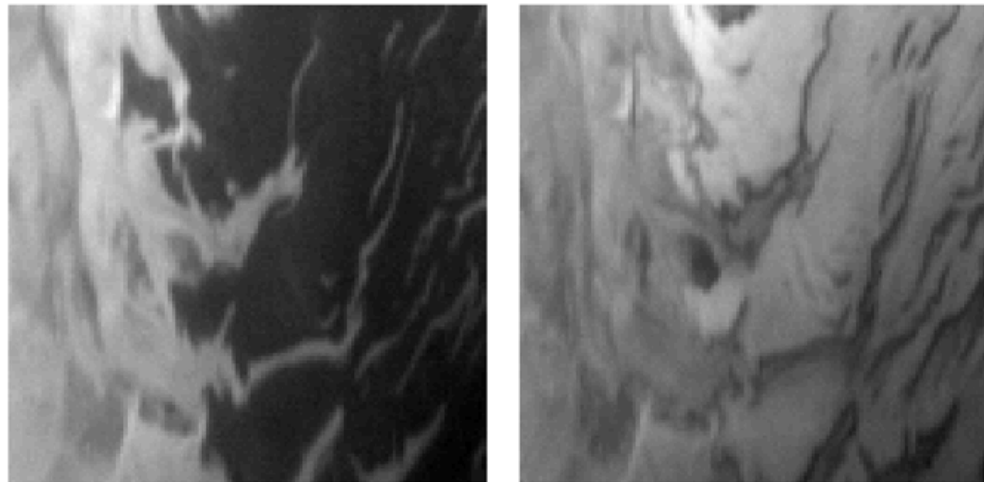
Output PSNR as a function of the input PSNR for three video sequences. (a) mobile, (b) tempete, and (c) coastguard CIF sequence.

Videos available at: [www.cipr.rpi.edu](http://www.cipr.rpi.edu)

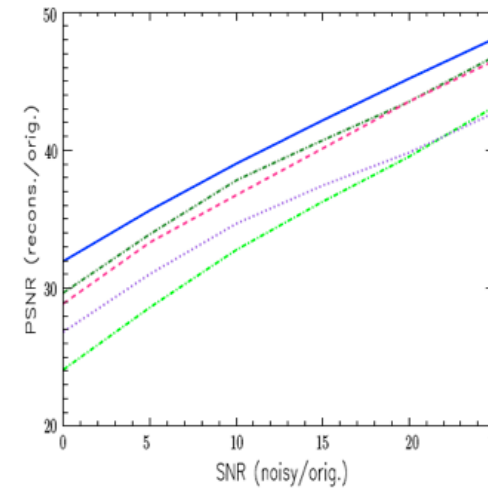
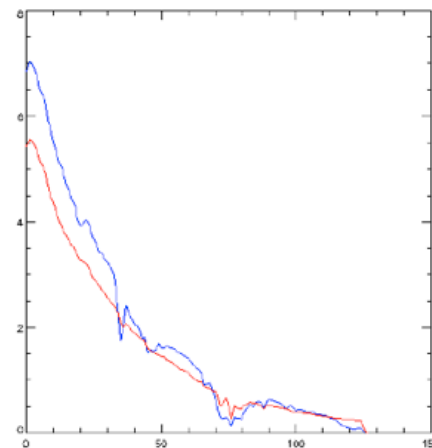


# Example: hyperspectral data denoising

DATA: OMEGA spectrometer on Mars Express ([www.esa.int/marsexpress](http://www.esa.int/marsexpress)) with 128 wavelength from 0.93 $\mu\text{m}$  to 2.73 $\mu\text{m}$ .



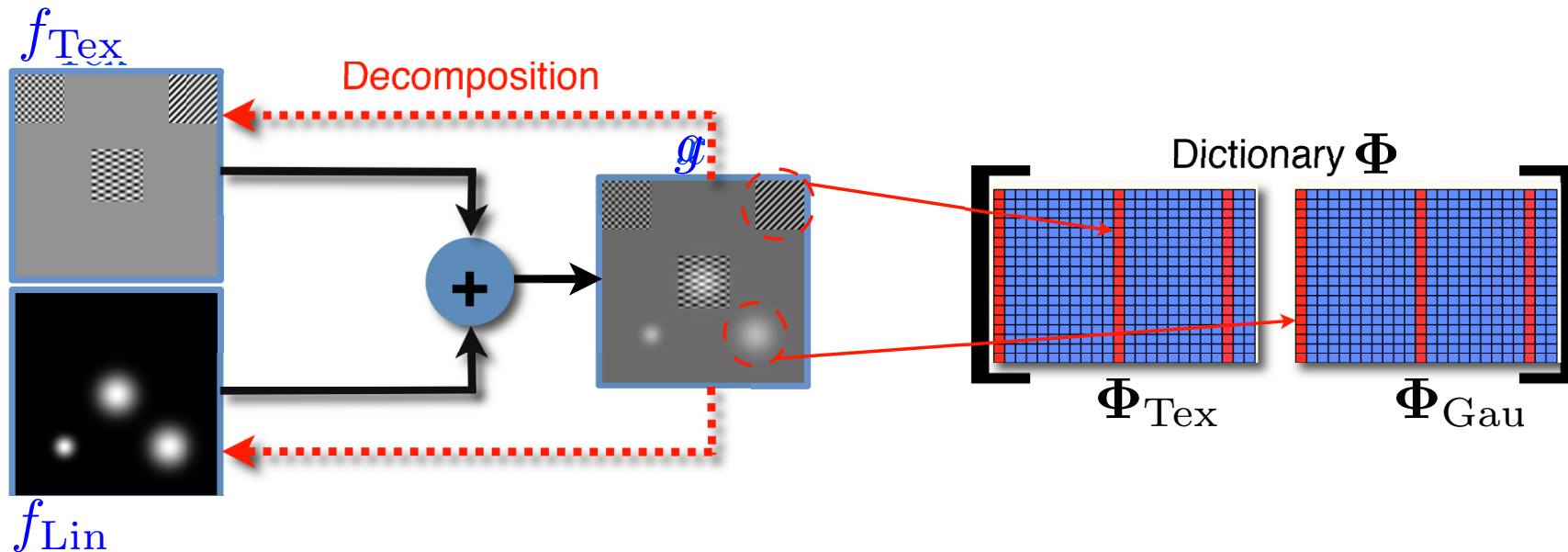
- LRF-Curvelets (10.3)
- - - Dual Tree Wavelets (4.0)
- ⋯ Surfacelets (6.4)
- ⋯ Surfacelets (4.0)
- - - Undecimated Wavelets (29)
- · - Decimated Wavelets (1.0)



Top row : Mars Express observations at two different wavelengths. Bottom-left : two spectra at two distinct pixels.  
Bottom-right : output PSNR as a function of the input PSNR for different transforms

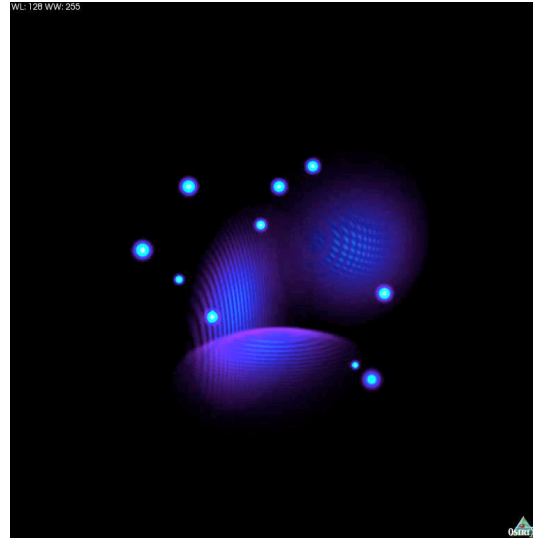
# Sparse component separation

- Separate an image into its morphological components from  $g = \sum_{k=1}^K f_k + \varepsilon = \sum_{k=1}^K \Phi_k \alpha_k + \varepsilon$ , each  $\alpha_k$  is sparse in  $\Phi_k$  but not (or less) sparse in  $\Phi_{k' \neq k}$ .
- A sparse decomposition problem solved by MCA [Starck et al. 2004-2009].



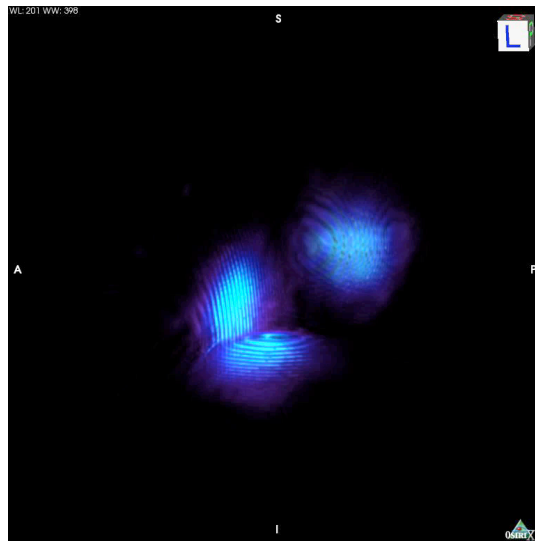
# Results

Original (3D shells + Gaussians)

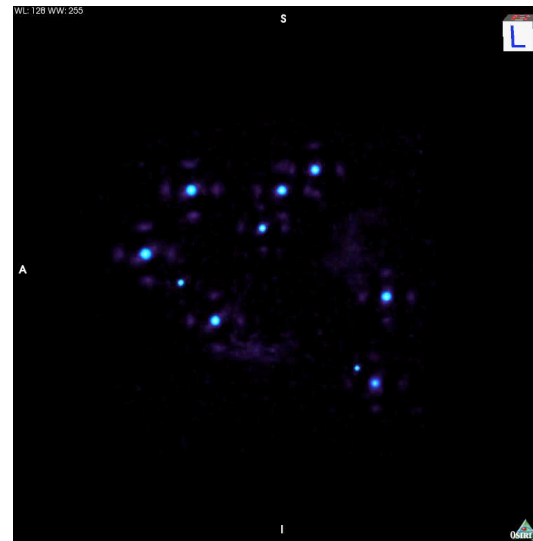


Dictionary  
RidCurvelets + 3D UDWT.

Shells



Gaussians



# Inpainting

---

- Restore an image from its degraded version with missing samples:

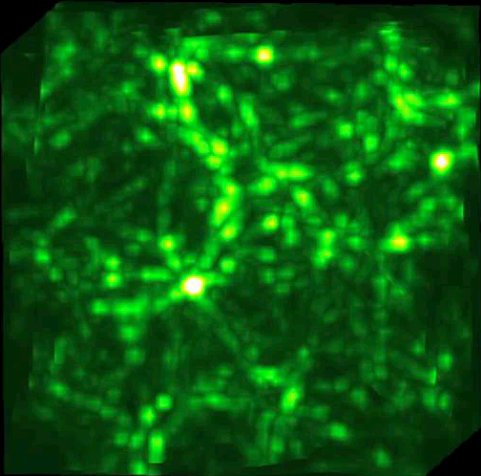
$$g = \mathbf{M}\Phi\alpha + \epsilon ,$$

- An instance of CS reconstruction.

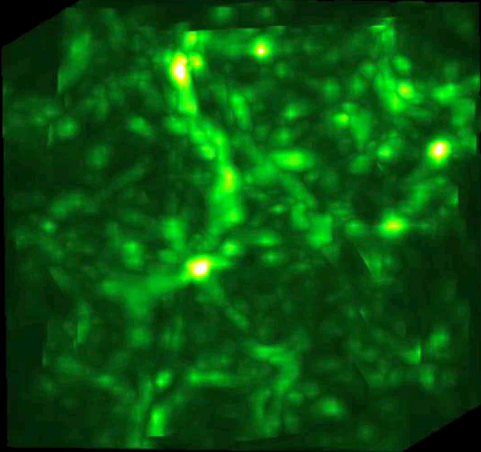
- Solve

$$\min_{\alpha} \|\alpha\|_p^p \quad \text{s.t.} \quad \|g - \mathbf{M}\Phi\alpha\|_2 \leq \epsilon .$$

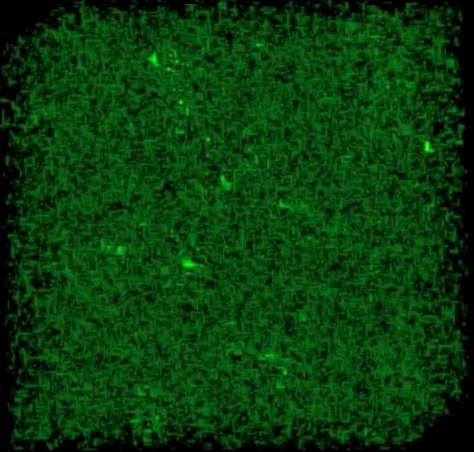
- Can be solved by several algorithms, here we use an adaptation of MCA.



Original



Inpainted



Mask

WL: 220 WW: 360

R

S

I



Dictionary  
BeamCurvelets

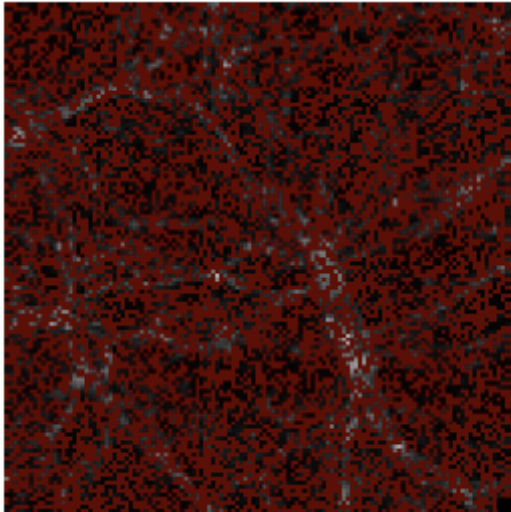
R

L

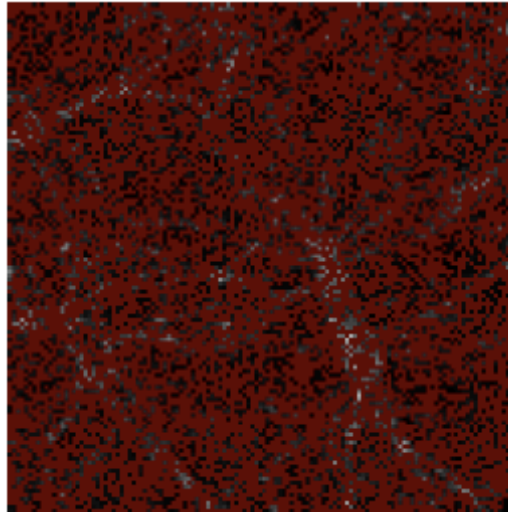
R



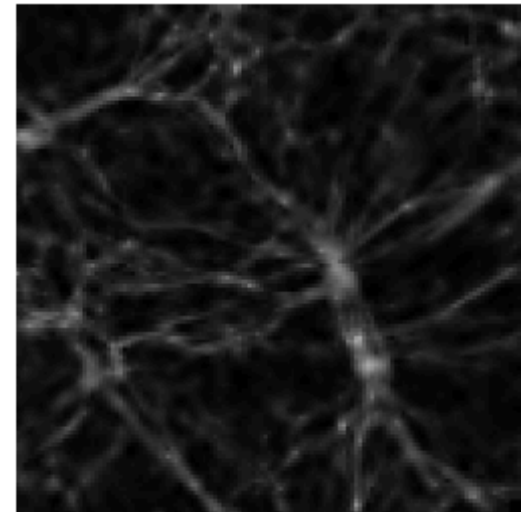
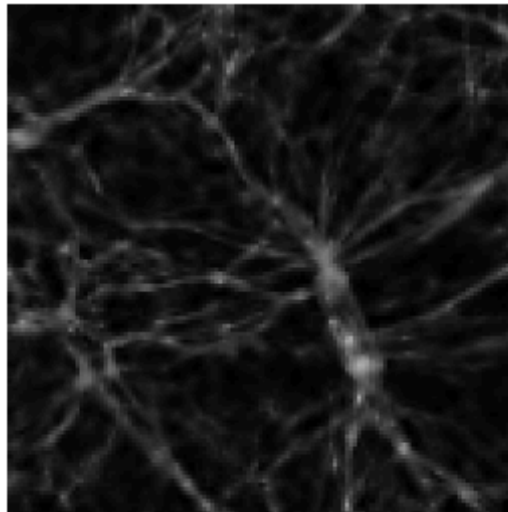
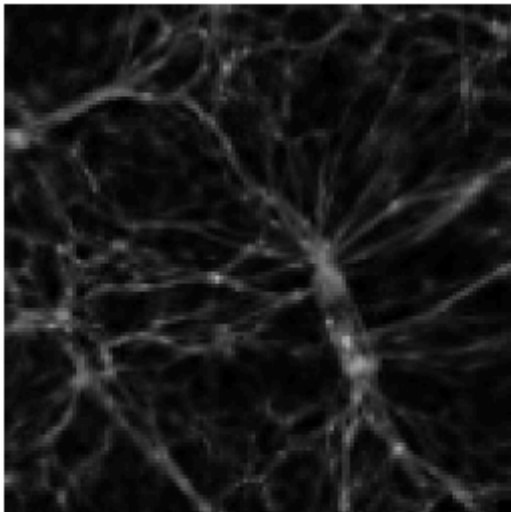
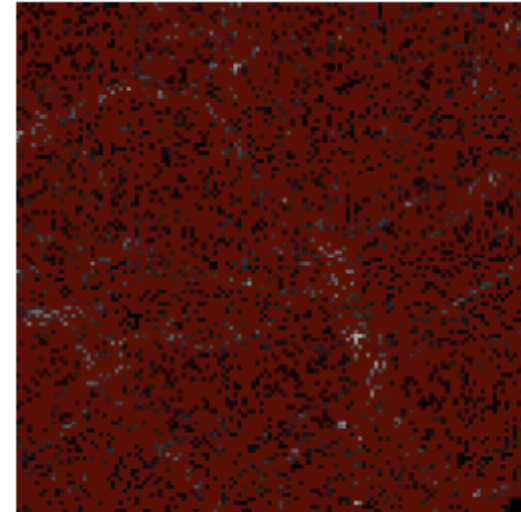
Masked (20%)



Masked (50%)

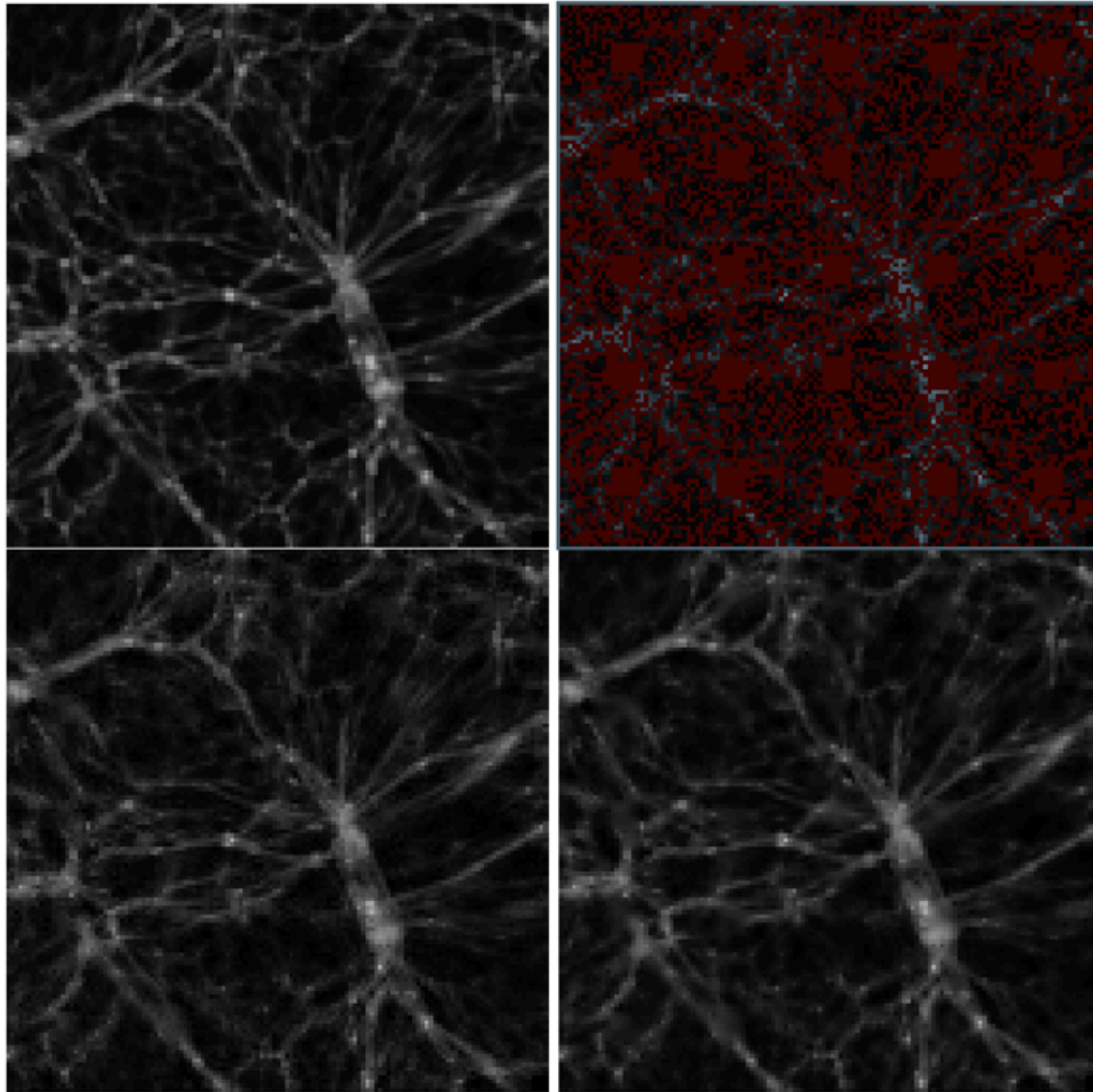


Masked (80%)



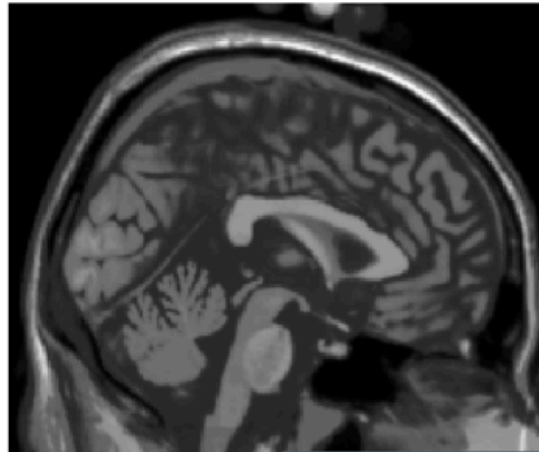
Central slice of the masked CDM data with 20, 50, and 80% missing voxels, and the inpainted maps. The missing voxels are dark red.





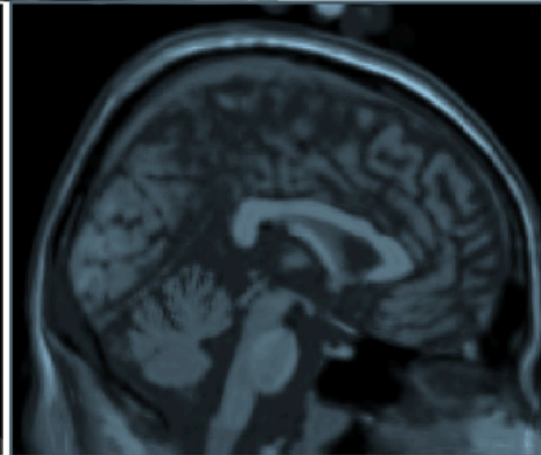
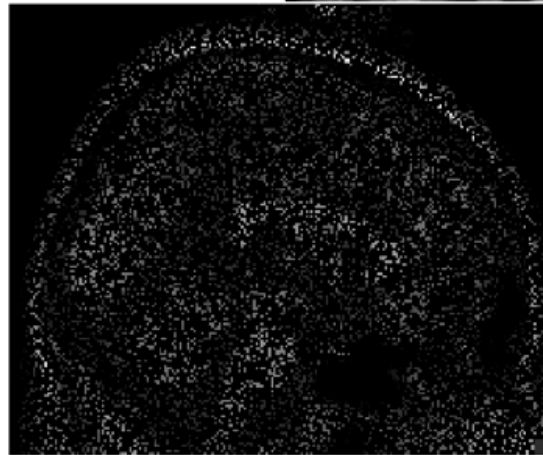
First row : original central frame of the CDM data cube, and degraded version with missing voxels in red.  
Bottom row : the filtered results using the RidCurvelets (left) and the BeamCurvelets (right).

A sagittal ((y; z)) slice of the original synthetic MRI volume from BrainWeb.

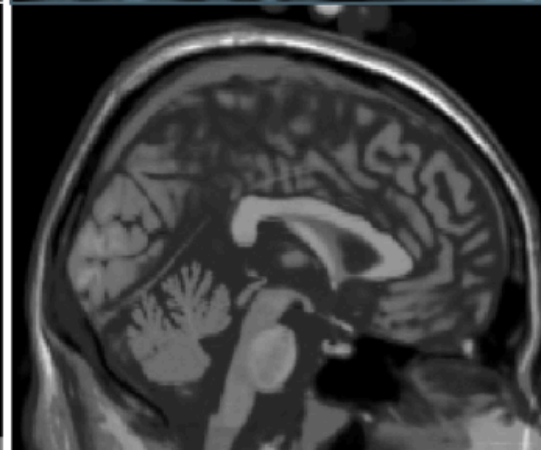


inpainting results with a FCT +UDWT dictionary.

random 80% missing voxels,



10% missing z slices.





# Conclusions

---

- Several 3D multiscale oriented representations.
- Adapted to sparsify several geometrical structures: filamentary and planar segments.
- Fast analysis and synthesis algorithms (FFT-based): parallel implementations.
- A wide variety of applications.



Jean-Luc Starck  
Fionn Murtagh

# Astronomical Image and Data Analysis

Second Edition



 Springer

Jean-Luc Starck  
Fionn Murtagh  
Jafal Fadili



# SPARSE IMAGE and SIGNAL PROCESSING

Wavelets, Curvelets,  
Morphological Diversity

CAMBRIDGE