

Wavefront sensing for aberration modeling in fluorescence MACROscopy

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Gilbert Engler⁴, Laure Blanc-Féraud⁵, Josiane Zerubia⁵,
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- 2 Forward problem: Characterizing the aberration
- 3 Inverse problem: Wavefront sensing from intensity data
- 4 Take home message and ongoing work

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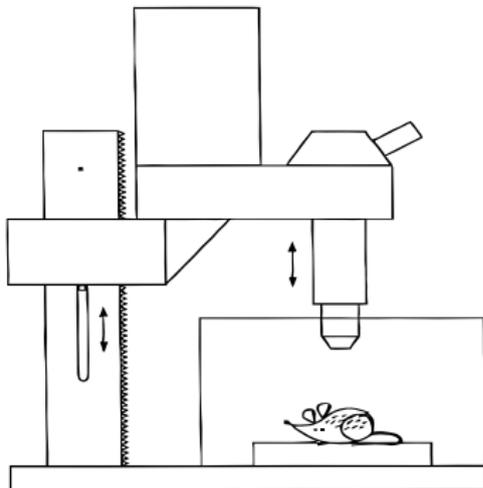
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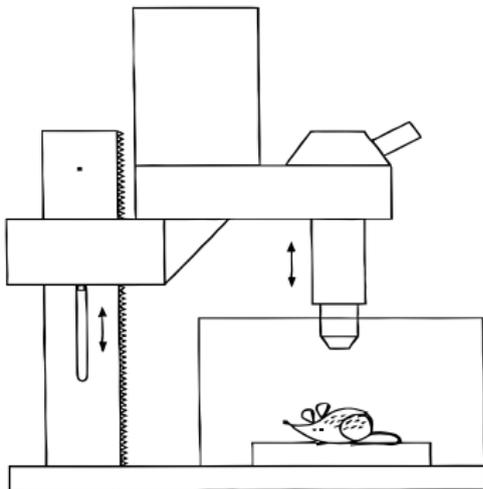
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Central points



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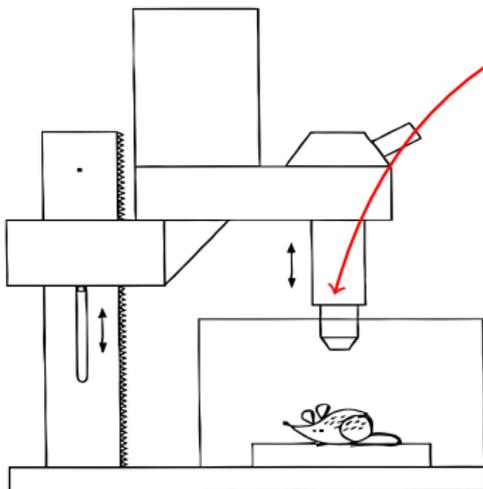
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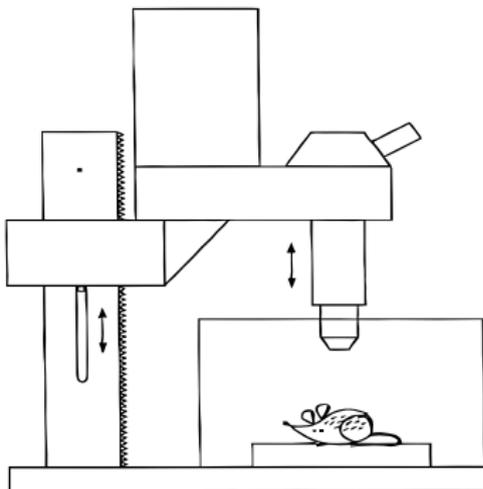
Conclusions

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- ▶ Low magnification objective lens is combined with apochromatic **zoom lens**,



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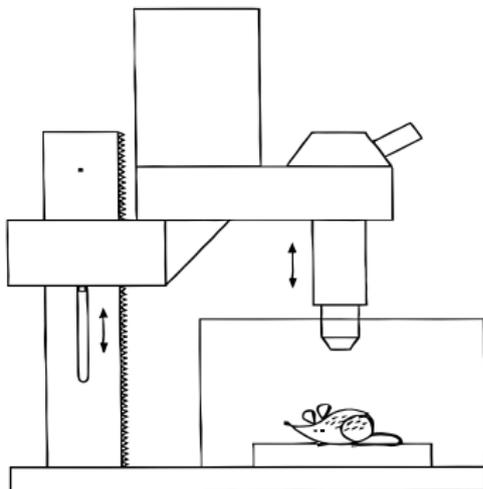
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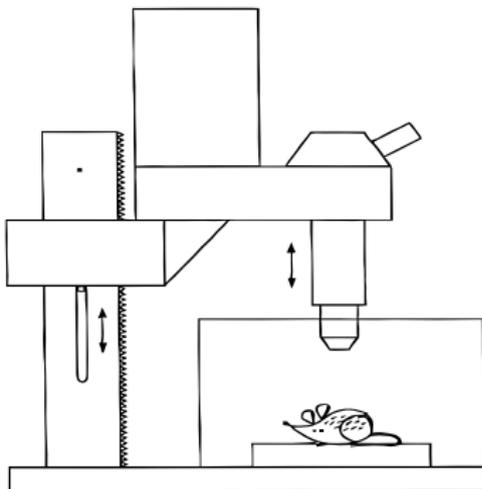
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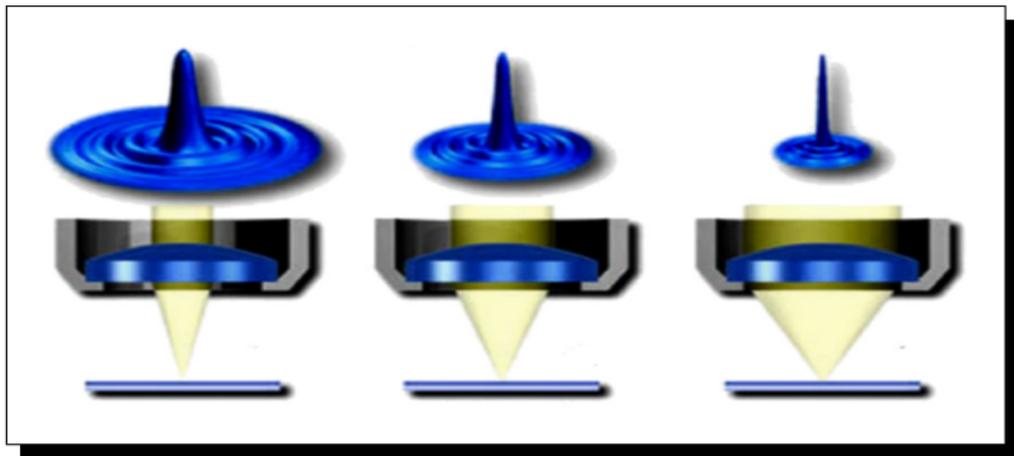
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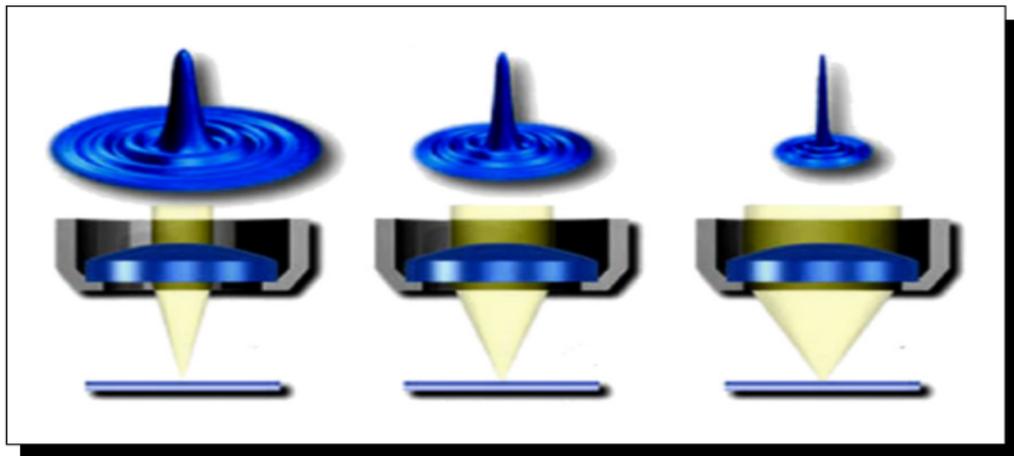
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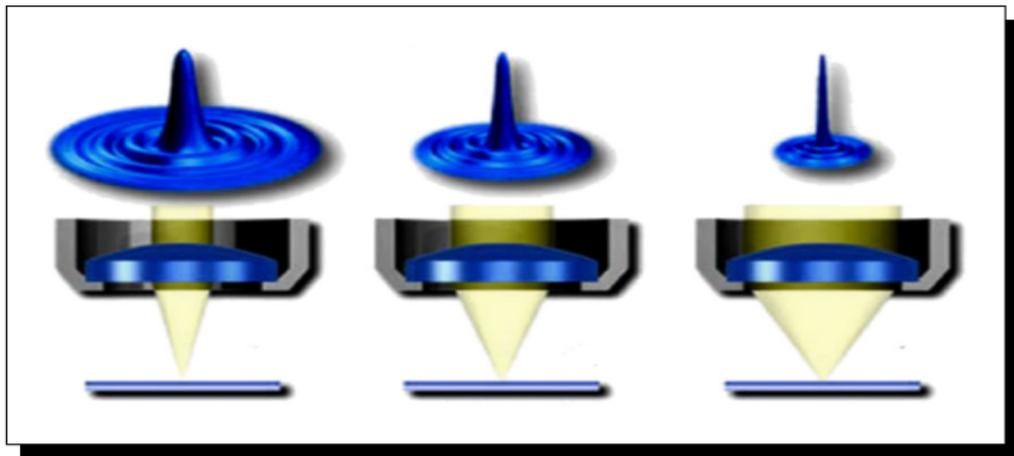
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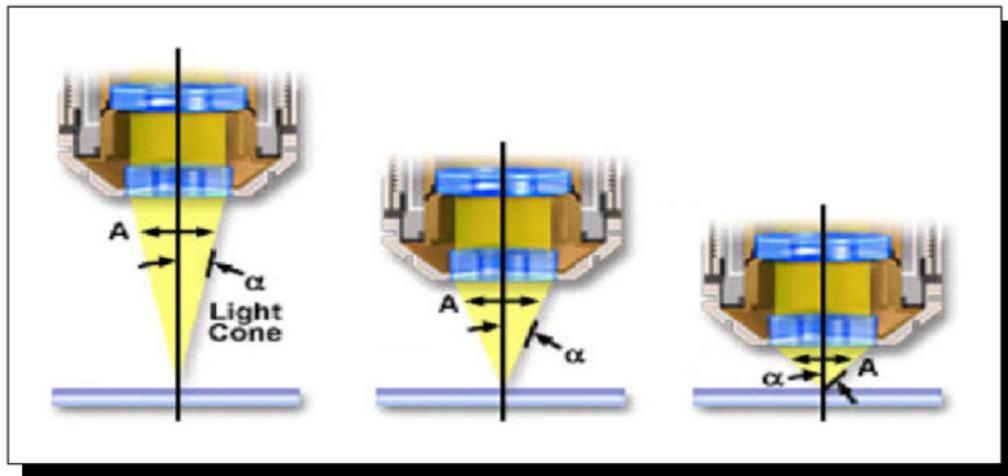


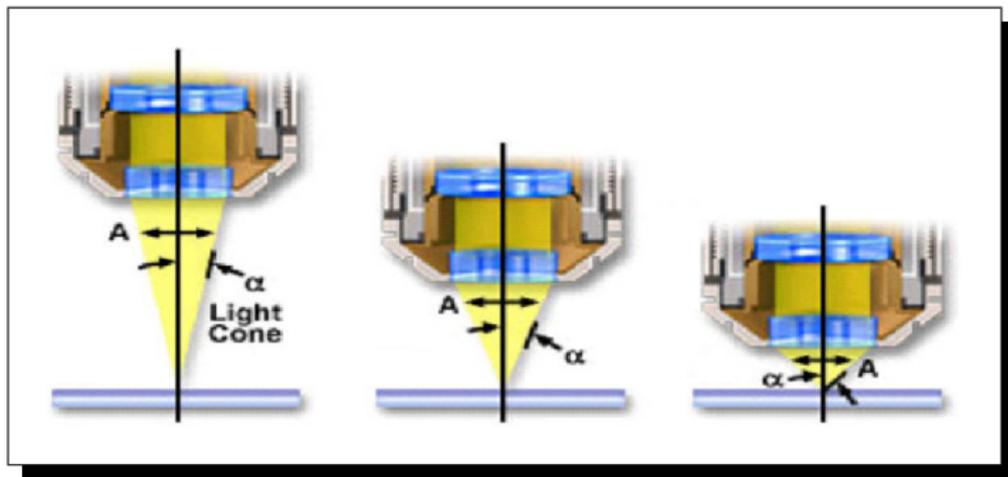
Numerical aperture decreases



Resolution decreases

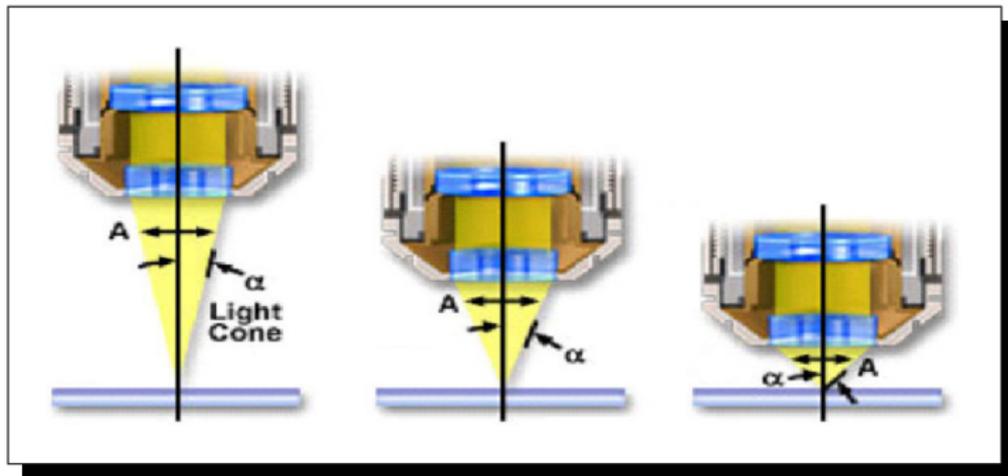






Numerical aperture increases





Numerical aperture increases



Working distance & Field-of-view decreases



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Maximum zoom position

Experimental impulse response

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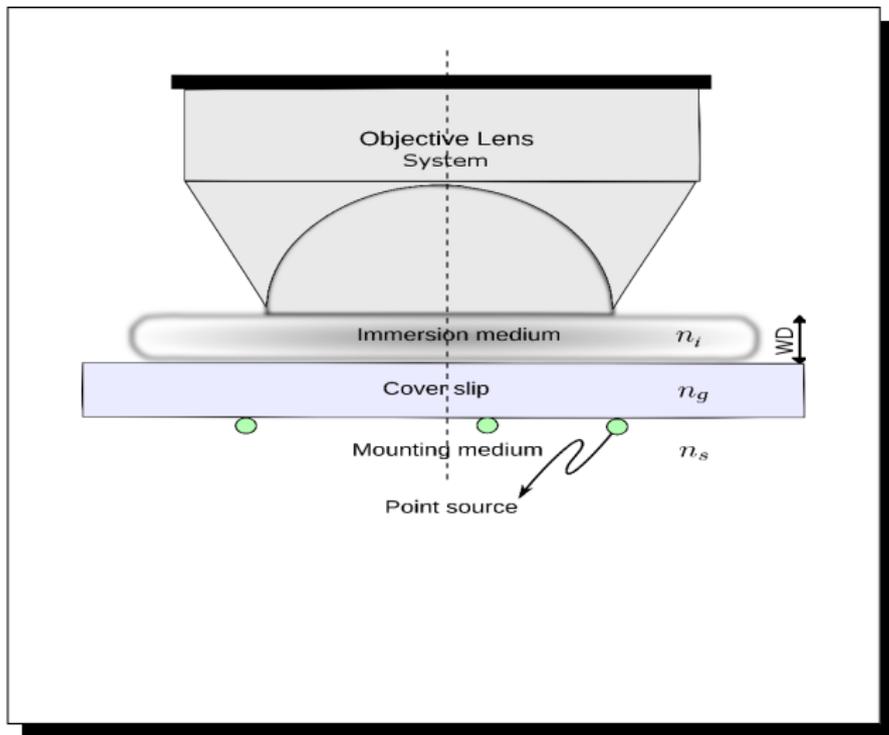
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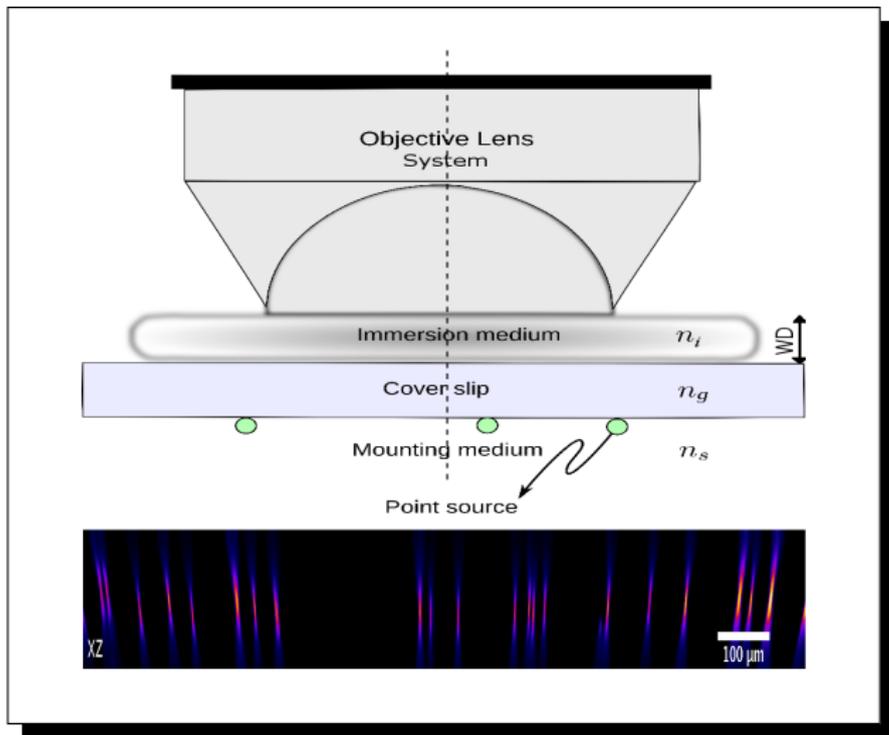


Figure 1: $2.5\mu\text{m}$ beads imaged using a Leica Widefield MacroFluo™ Z16 APO fit with $5\times$ objective and the $1.6\times$ zoom. ©Herbomel lab, Pasteur Institute.

MACROscopes-Are they really the best of the two worlds?



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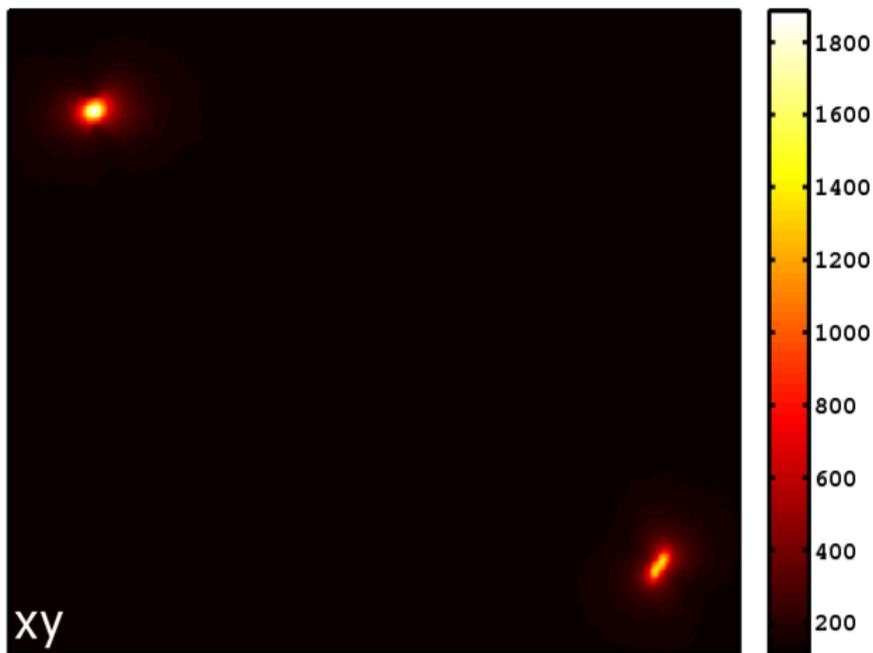


Figure 2: Axial projection of the beads. ©Herbomel lab, Pasteur Institute.



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Forward problem: Characterizing the aberration

- ▶ If $P(k_x, k_y, z)$ is the 2D **complex pupil function** and λ is the wavelength, the amplitude PSF can be calculated by just $2N_z$ number of 2D FFTs as

$$h_A(x, y, z; \lambda) = \int_{k_x} \int_{k_y} P(k_x, k_y, z) \exp(j(k_x x + k_y y)) dk_y dk_x$$

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- ▶ and the **incoherent PSF** is

$$h_{Th}(\mathbf{x}; \lambda_{ex}, \lambda_{em}) = C |h_A(\mathbf{x}; \lambda_{ex})| \times |h_A(\mathbf{x}; \lambda_{em})|$$

- ▶ λ_{ex} and λ_{em} are the excitation and emission peak wavelengths.

Limiting apertures overlapping

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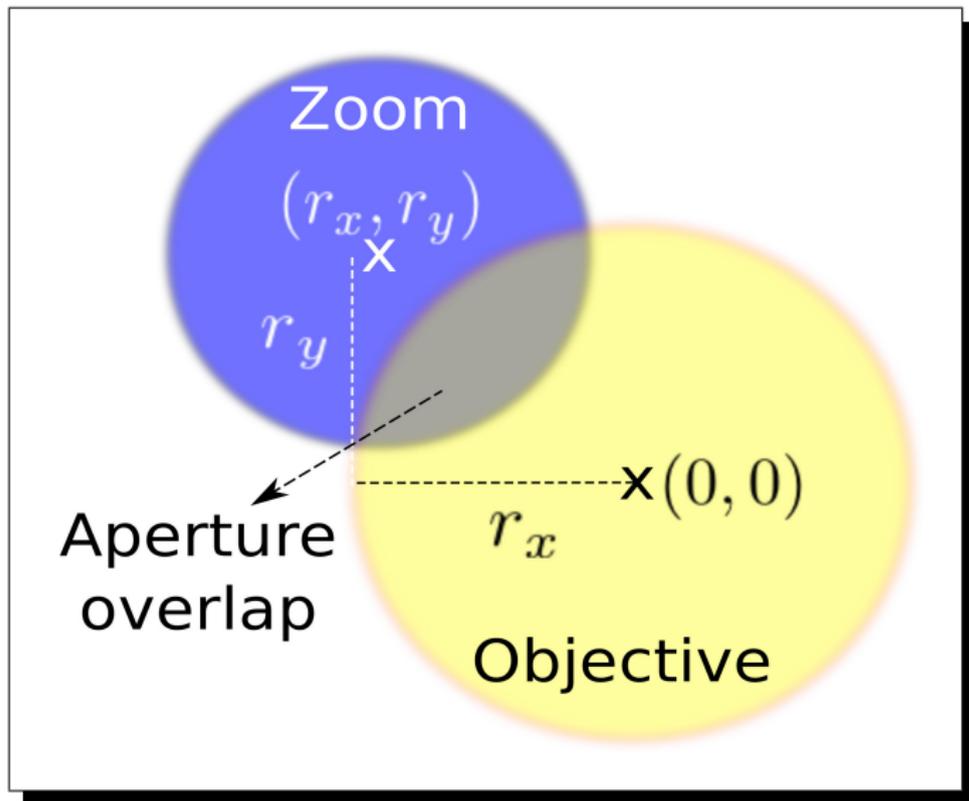
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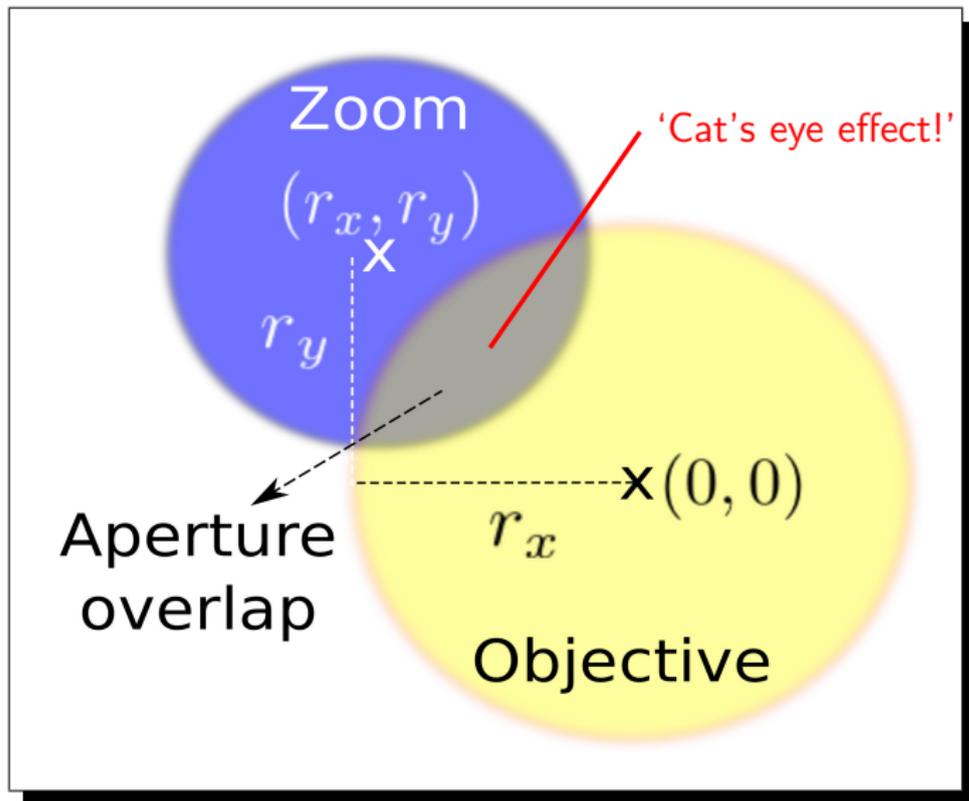
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Figure 3: Lens viewed from the front.
(Photograph by Peter Boehmer.)

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Figure 3: Lens viewed from the front.
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Figure 4: Lens viewed from the side.
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$$P_m(k_x, k_y, z; \lambda) = \begin{cases} e^{jk_0\phi(\theta_i, \theta_s, z)}, & \text{if } \sqrt{k_x^2 + k_y^2} < \frac{2\pi}{\lambda} \text{NA}_{\text{Obj}} \\ 0, & \text{otherwise.} \end{cases}$$



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- ▶ NA_{Obj} and NA_{Zo} are the objective and zoom lens NA;
 (r_x, r_y) are the relative displacements.

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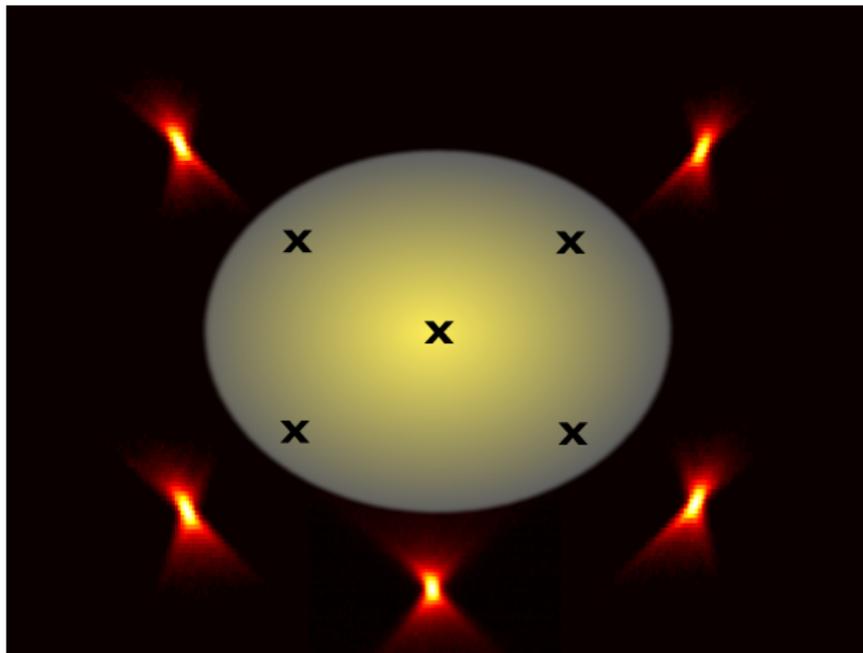


Figure 5: We can characterize the behavior at any position in the lateral field.

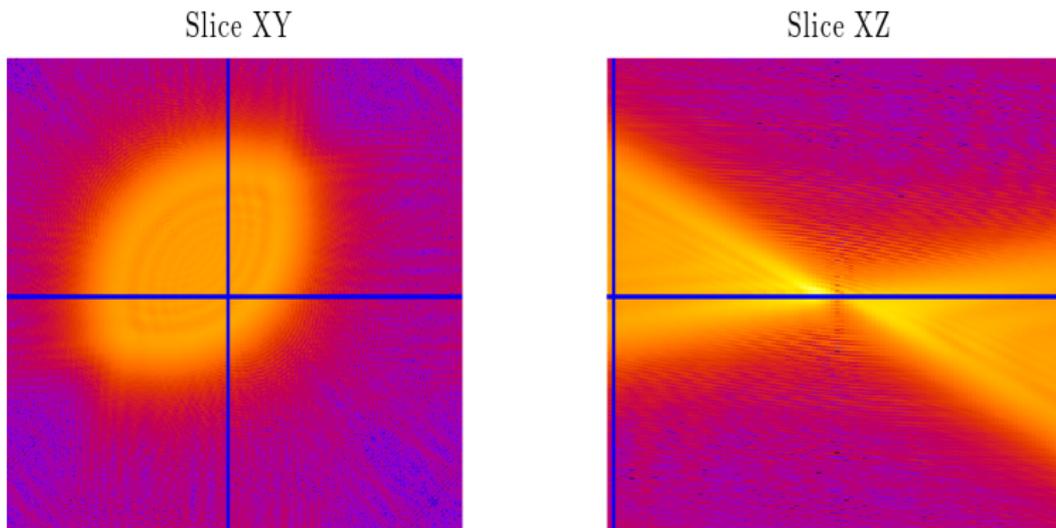


Figure 6: Theoretically calculated MACROscope PSF in log scale. NA= 0.5, lateral sampling 178.33nm, axial sampling 1000nm.

Out-of-focus highlights (OOFH)

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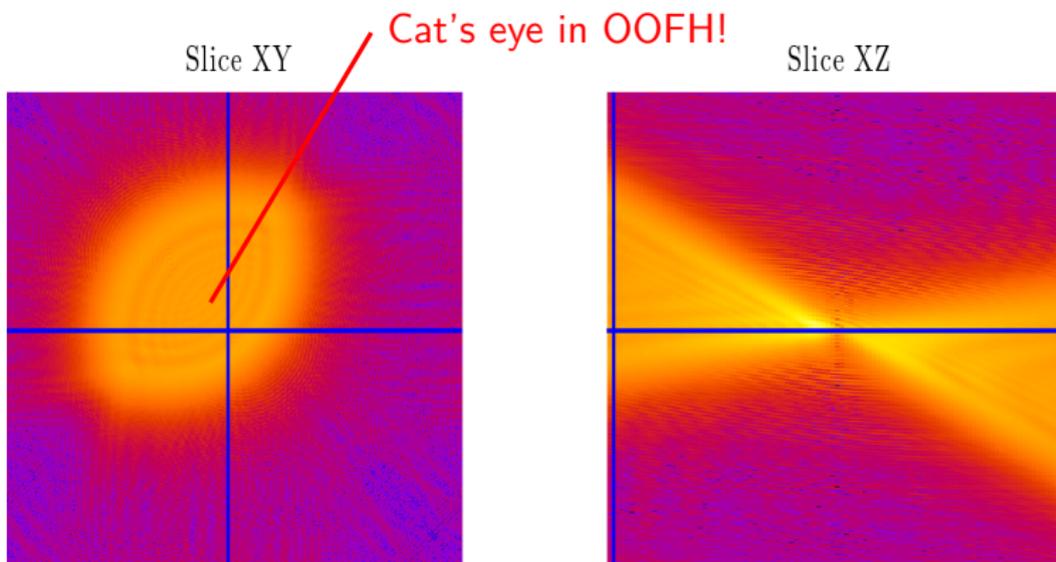


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 - it can be also **retrieved from the observed intensities**,
- ▶ the **aberrations** in the optics of the objective can be determined by studying this phase,
- ▶ the estimated wavefront can be used to **correct** the aberrated optical path.

- ▶ For uncorrelated **low photon count** data the observation is:

$$i(\mathbf{x}) = \mathcal{P}\{|h_A(\mathbf{x})|^2 + b(\mathbf{x})\}, \forall \mathbf{x} \in \Omega_s$$

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- ▶ Considering **Poissonian photon counting statistics**, the likelihood of obtaining image $i(x)$ from a diffraction-limited point source:

$$\Pr(i|h_A) = \prod_{\mathbf{x} \in \Omega_s} \frac{(h_A + b)(\mathbf{x})^{i(\mathbf{x})} \exp(-(h_A + b)(\mathbf{x}))}{i(\mathbf{x})!}$$

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- ▶ From the **Bayes' theorem**, the *a posteriori* is

$$\Pr(h_A|i) = \frac{\Pr(i|h_A) \Pr(h_A)}{\Pr(i)}$$

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- ▶ Estimate the **near-focus amplitude distribution**, \hat{h}_A , by maximizing the *a posteriori* (MAP) or minimizing the cologarithm of the *a posteriori*

$$\hat{h}_A(\mathbf{x}; \varphi_{\text{aberr}}) = \arg \min_{h_A(\mathbf{x})} -\log[\text{Pr}(h_A|i)], \text{ s. t. } k_{\text{MAX}} < \frac{2\pi \text{NA}}{\lambda_{\text{ex}}}$$

- k_{MAX} is the **pupil support**,
 - ▶ this can be solved by using a **fixed-point iterative algorithm**.

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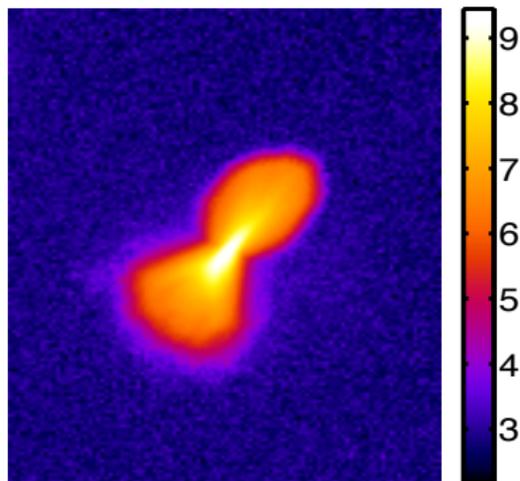


Figure 7: Radially projected $2.5\mu\text{m}$ observed intensity volume. ©Imaging Center, IGBMC, France.

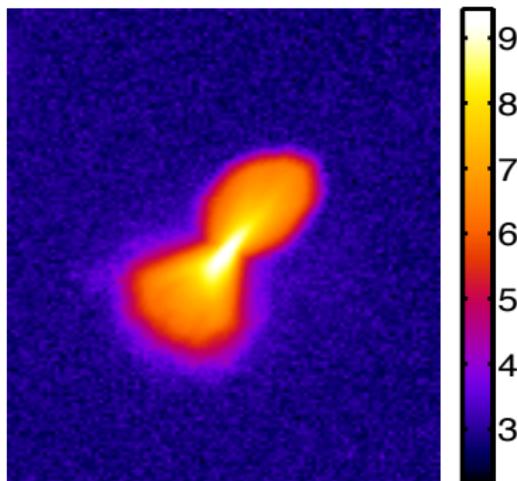


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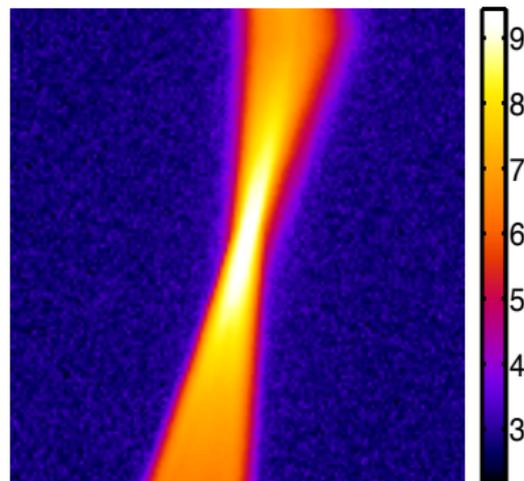


Figure 8: Axially projected $2.5\mu\text{m}$ observed intensity volume. ©Imaging Center, IGBMC, France.

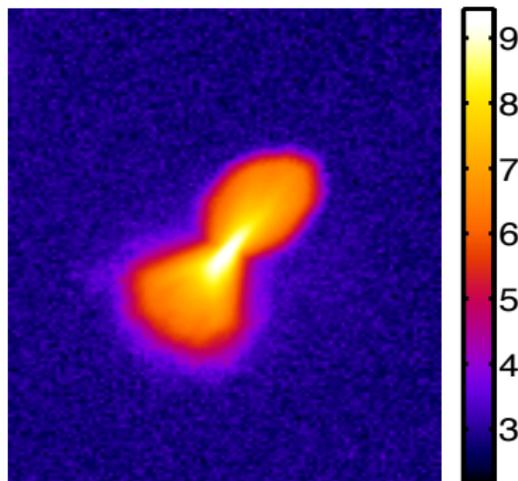


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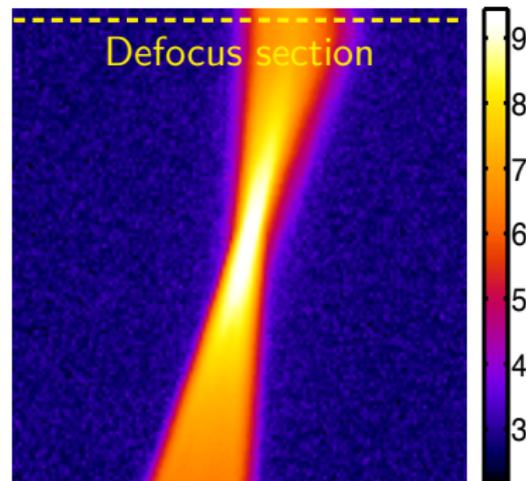


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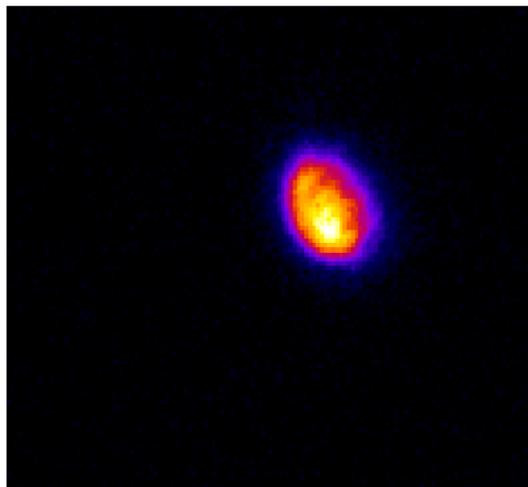


Figure 9: OOFH radial section of the observed volume, $z = -57\mu\text{m}$.

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Costantini, M. (1998). A novel phase unwrapping method based on network programming. *IEEE Tran. on Geoscience and Remote Sensing*, 36, 813-821.

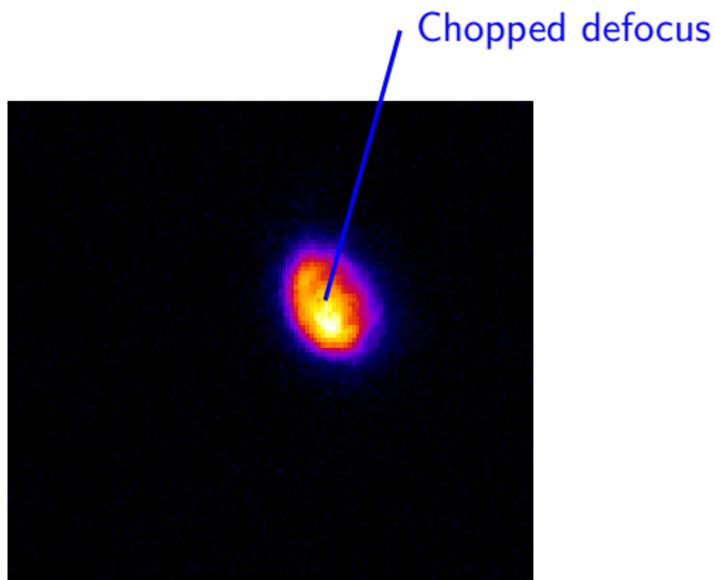


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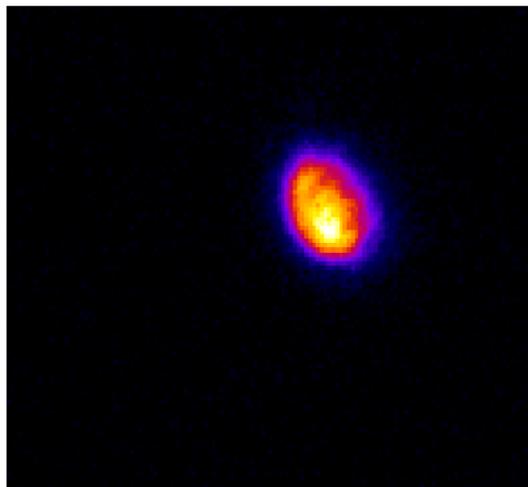


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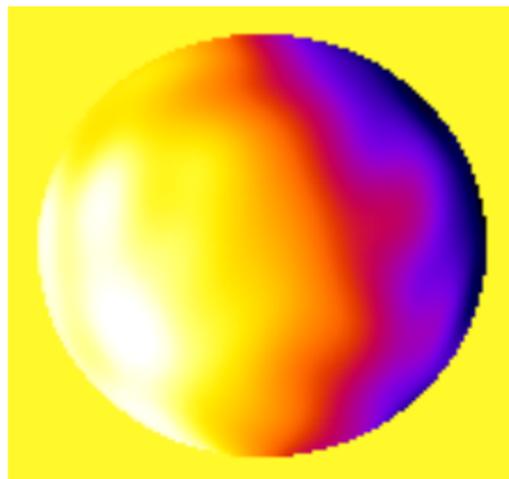


Figure 10: Retrieved **unwrapped pupil phase** from the intensity images $\tau = 0.9$ and the maximum number of iteration is 40.

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Chopped pupil

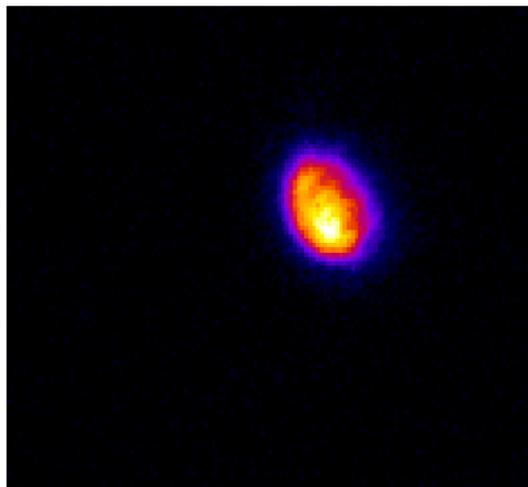


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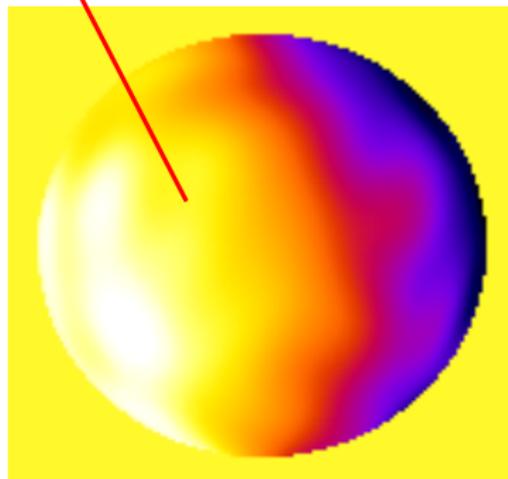


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- ▶ ongoing work: **restore the images** by correcting for the field aberration and also the diffraction effects.

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